

Weak Interaction Physics at *TwinSol*

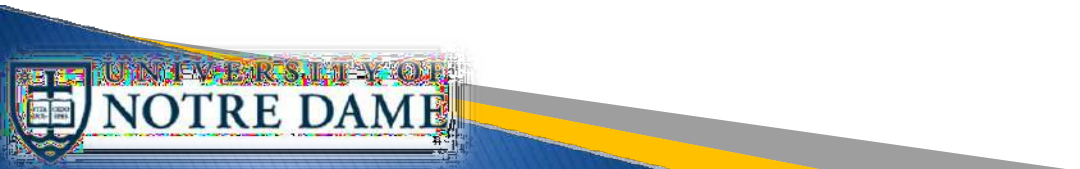
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*Emeritus Professor

**Associate Professor

P.I. : Ion-Trap Project



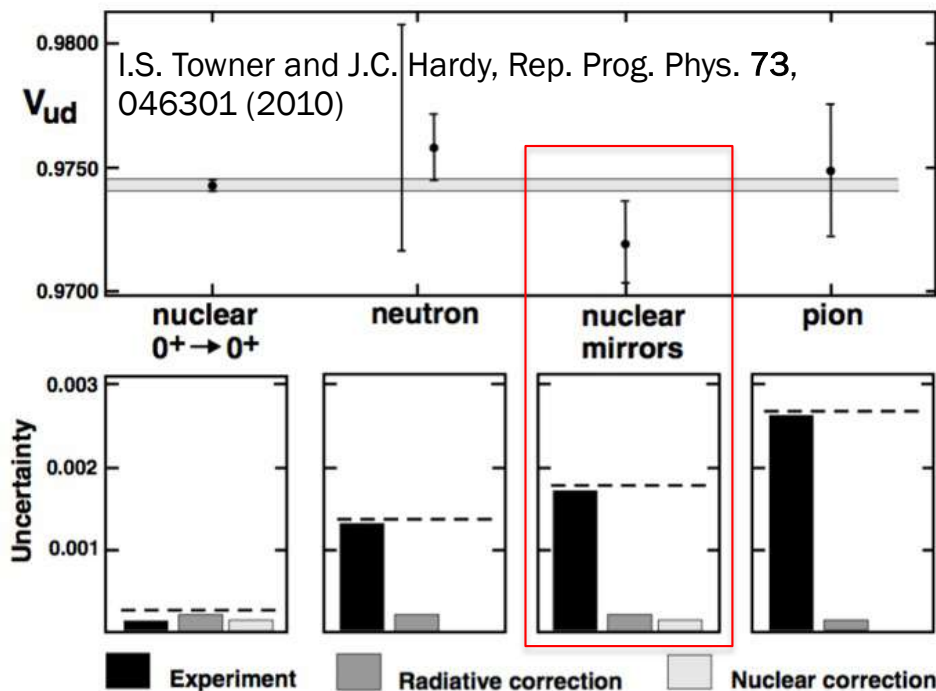
Precision Tests of the Standard Model

Via the CKM matrix unitarity test.

$$\sum_i |V_{ui}|^2 = |V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 = 1 \quad \begin{pmatrix} |d_w\rangle \\ |s_w\rangle \\ |b_w\rangle \end{pmatrix} = \begin{pmatrix} V_{ud} & V_{us} & V_{ub} \\ V_{cd} & V_{cs} & V_{cb} \\ V_{td} & V_{ts} & V_{tb} \end{pmatrix} \begin{pmatrix} |d_s\rangle \\ |s_s\rangle \\ |b_s\rangle \end{pmatrix}$$

↘ largest element; *involves nuclear physics*

- Non-unitarity: **new physics** (or uncover erroneous data/theoretical corrections)



- For a reliable CKM matrix unitarity test, V_{ud} needs to be both **precise** and **accurate**.
- Accuracy** can be tested by a determination of V_{ud} using multiple systems.
- Pure Fermi β -decay: most precise at this moment.

V_{ud} from mixed-mirror transitions

- Another avenue to extract V_{ud} , test CVC, and test δ_c .



- Only 5 nuclei so far.
- Parameters needed:

✓ Half-life



✓ Branching ratios

✓ Q-values

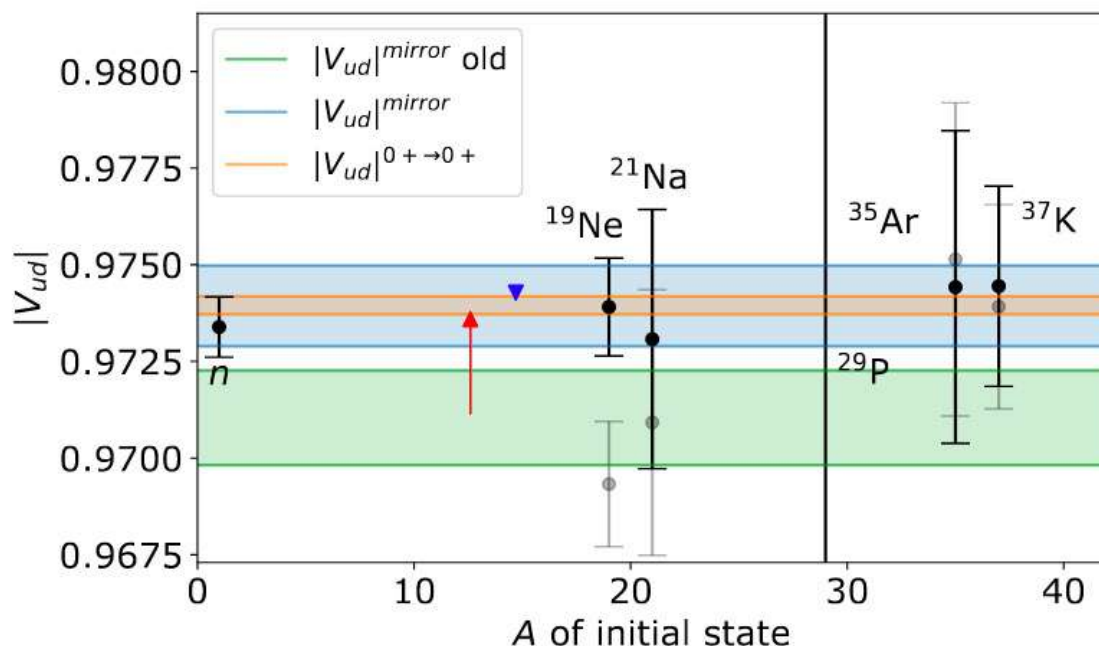
✓ Fermi-to-Gamow

Teller mixing ratio ρ



ρ determined by measuring either:

- β asymmetry parameter A_β
- ν asymmetry parameter B_ν
- β - ν angular correlation $a_{\beta\nu}$



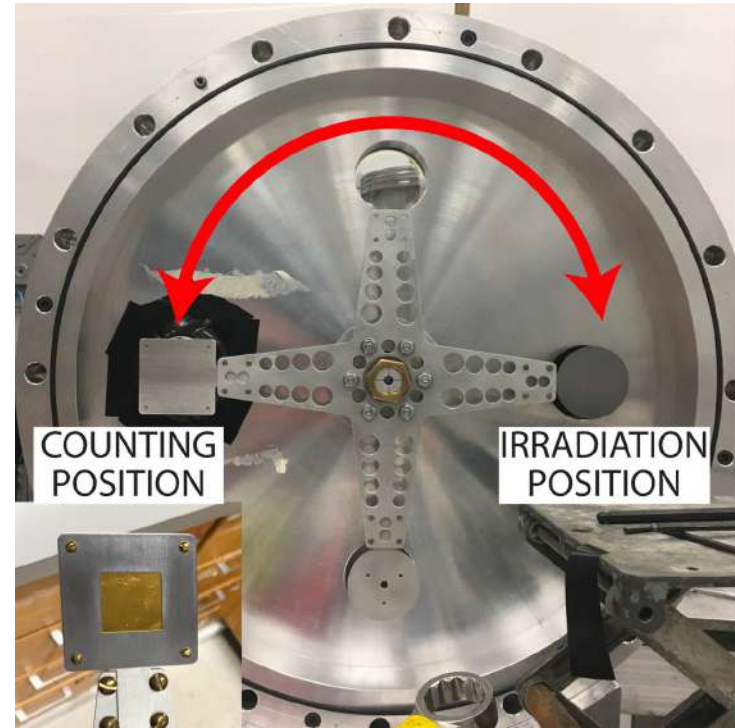
V_{ud} from mirror transitions is currently*:

- ~3.5 times less precise than pure Fermi.
- In agreement with pure Fermi, after re-evaluation using new theoretical corrections.

* <https://arxiv.org/pdf/2102.03458.pdf>

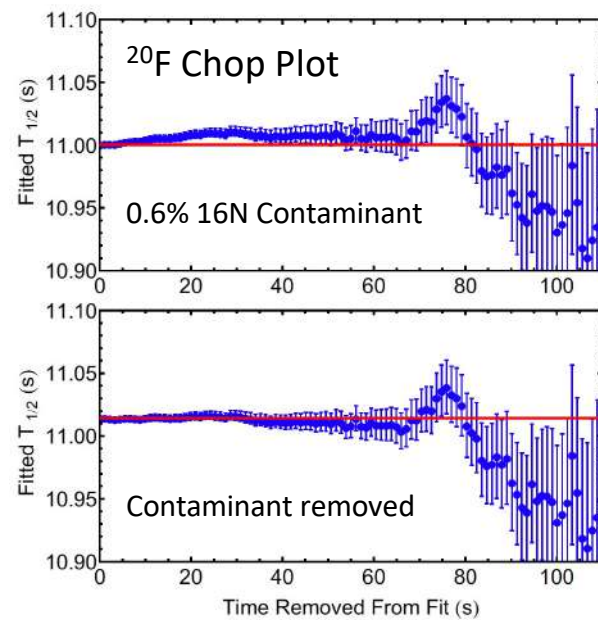
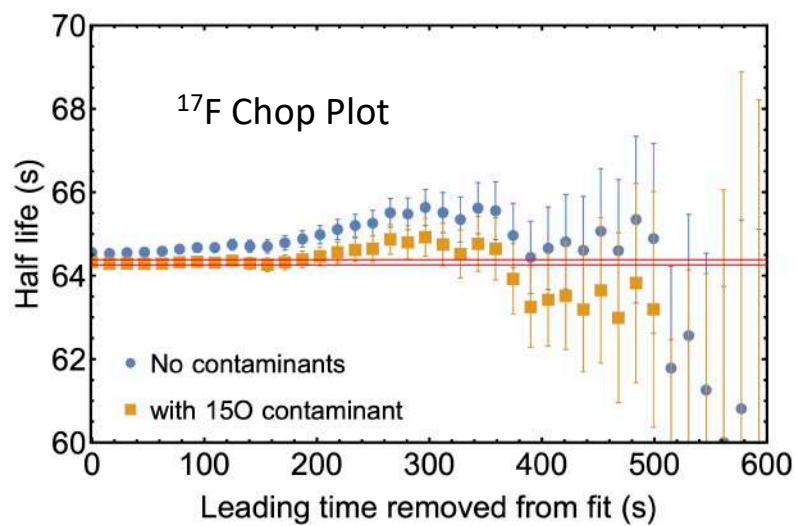
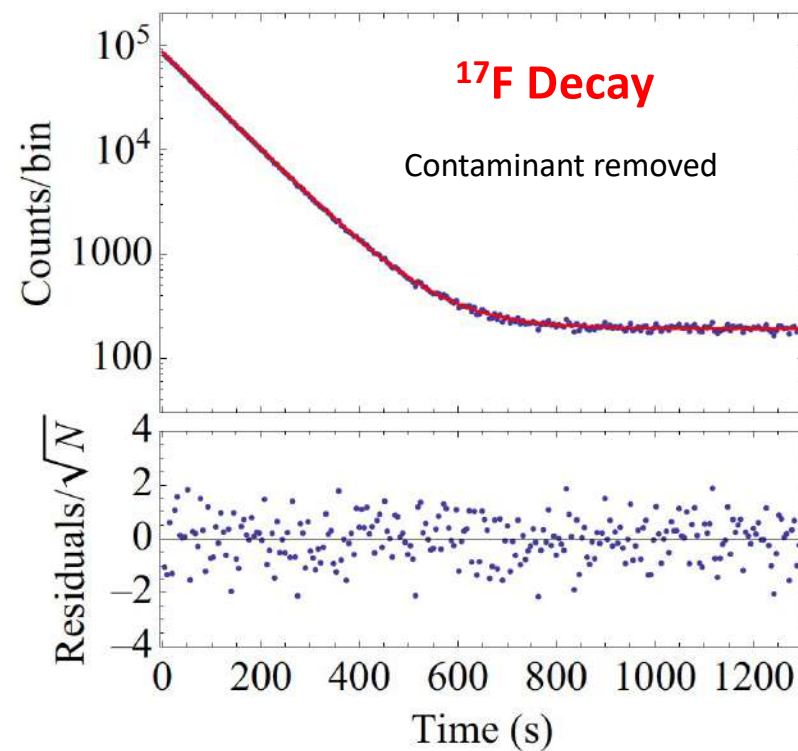
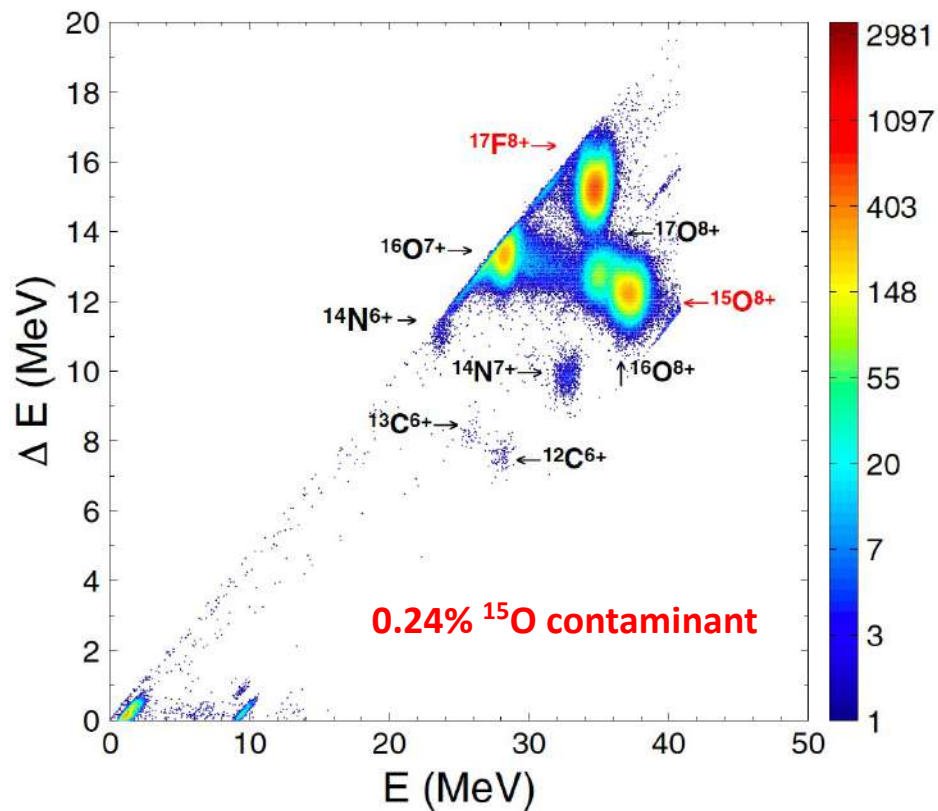
Experimental Considerations

- (1) Device to **rotate** the stopper away from the beam location.
- (2) **Gold stopper**: reduces contaminants formed by the beam incident on Ta impurities.
- (3) **Detector**: 1 mm plastic scintillator on a light guide: reduces interference from γ -radiation.
- (4) **Quartz-window PMT**: reduces background from the window by about a **factor of 10**.
- (5) Beam removed during counting period using a **pre-tandem electrostatic deflector**.

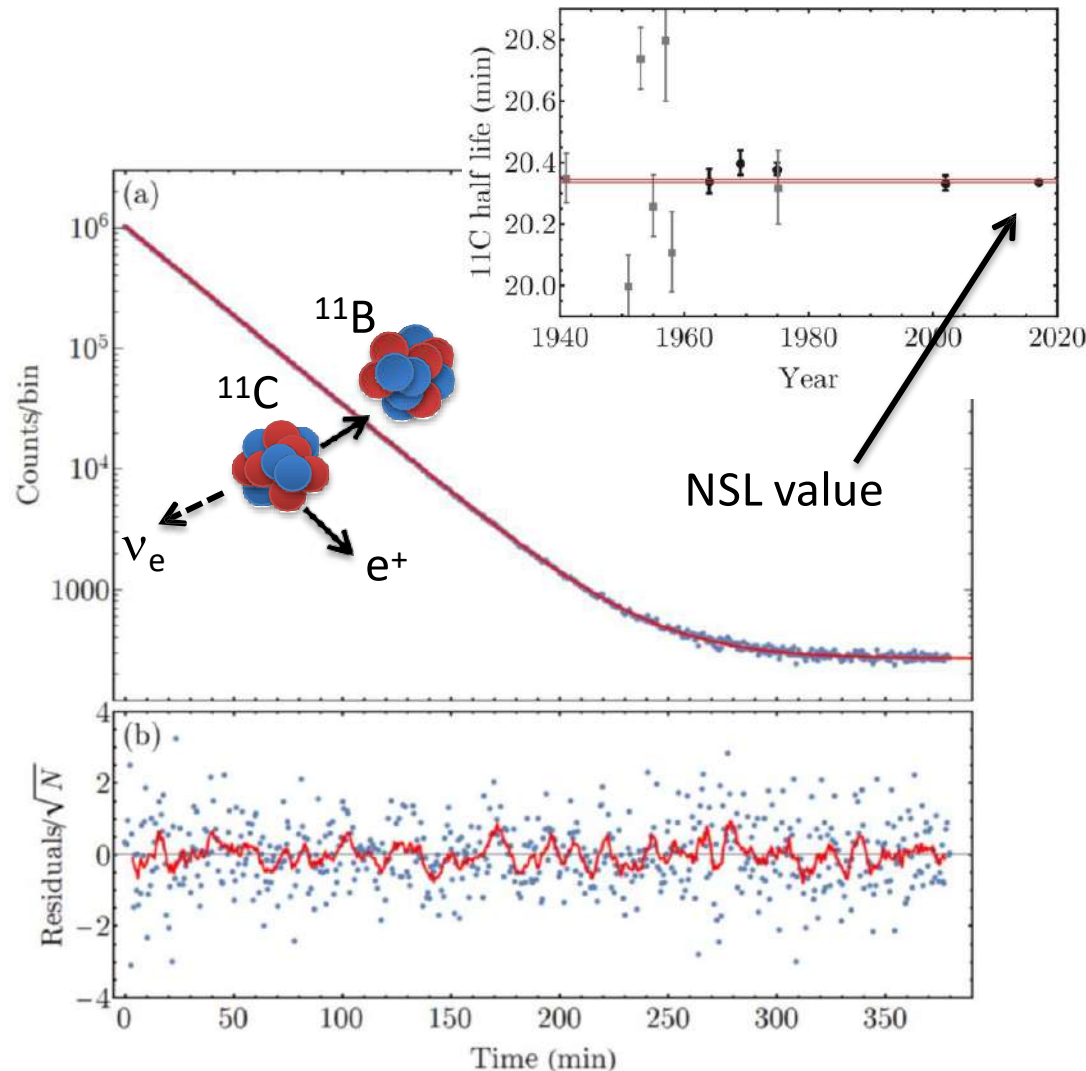


Beam-direction view of the apparatus.

The cross structure (meant for simultaneous irradiation while counting) was not used in these experiments. Note the gold stopper.

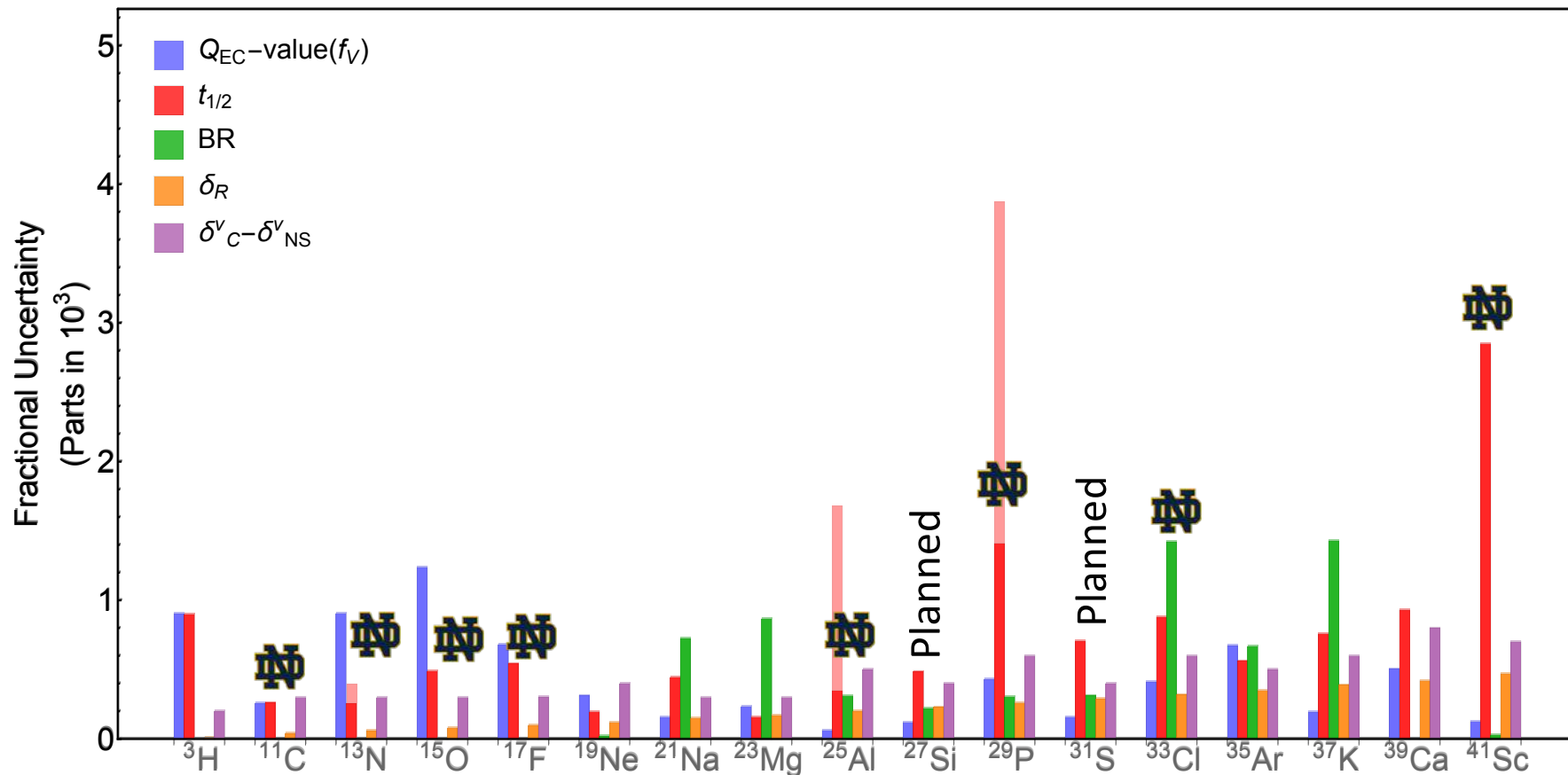


^{11}C half-life measurement at the NSL



- Superaligned mixed mirror decays \rightarrow precision tests of SM through V_{ud}
- Measurement of ^{11}C $t_{1/2}$ @ UND \rightarrow **most precise** of all SA mixed decays
- **Lightest** of all mirror transitions: to test CVC
- Measurement of mixing ratio using the future ion trap is planned to extract V_{ud} .

Ft-values of Mirror Transitions



^{11}C : A. Valverde *et al.*, PRC **97**, 035503 (2018)

^{15}O : D. Burdette *et al.*, PRC **101**, 055504 (2020)

^{17}F : M. Brodeur *et al.*, PRC **93**, 025503 (2016)

^{25}Al : J. Long *et al.*, PRC **96**, 015502 (2017)

^{29}P : J. Long *et al.*, PRC **101**, 015501 (2020)

^{13}N : J. Long *et al.*, in preparation

^{33}Cl : P. O'Malley *et al.*, in preparation

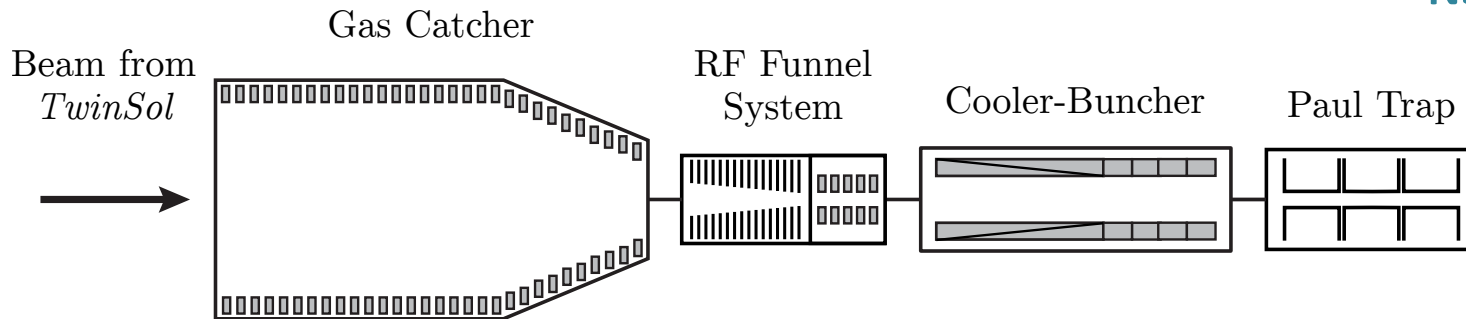
^{41}Sc : B. Liu *et al.*, in preparation

Others: ^{20}F : D. Burdette *et al.*, PRC **99**, 015501 (2019)

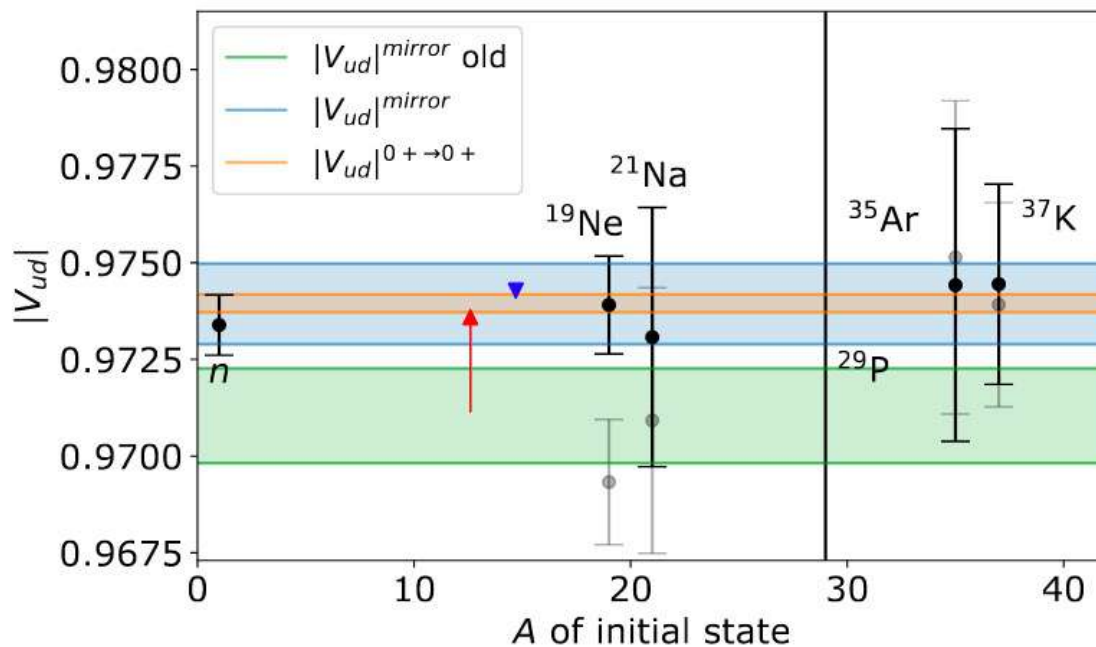
^{28}Al : B. Liu *et al.*, in preparation

“St. Benedict” to measure ρ

NSF: MRI PHY-1725711



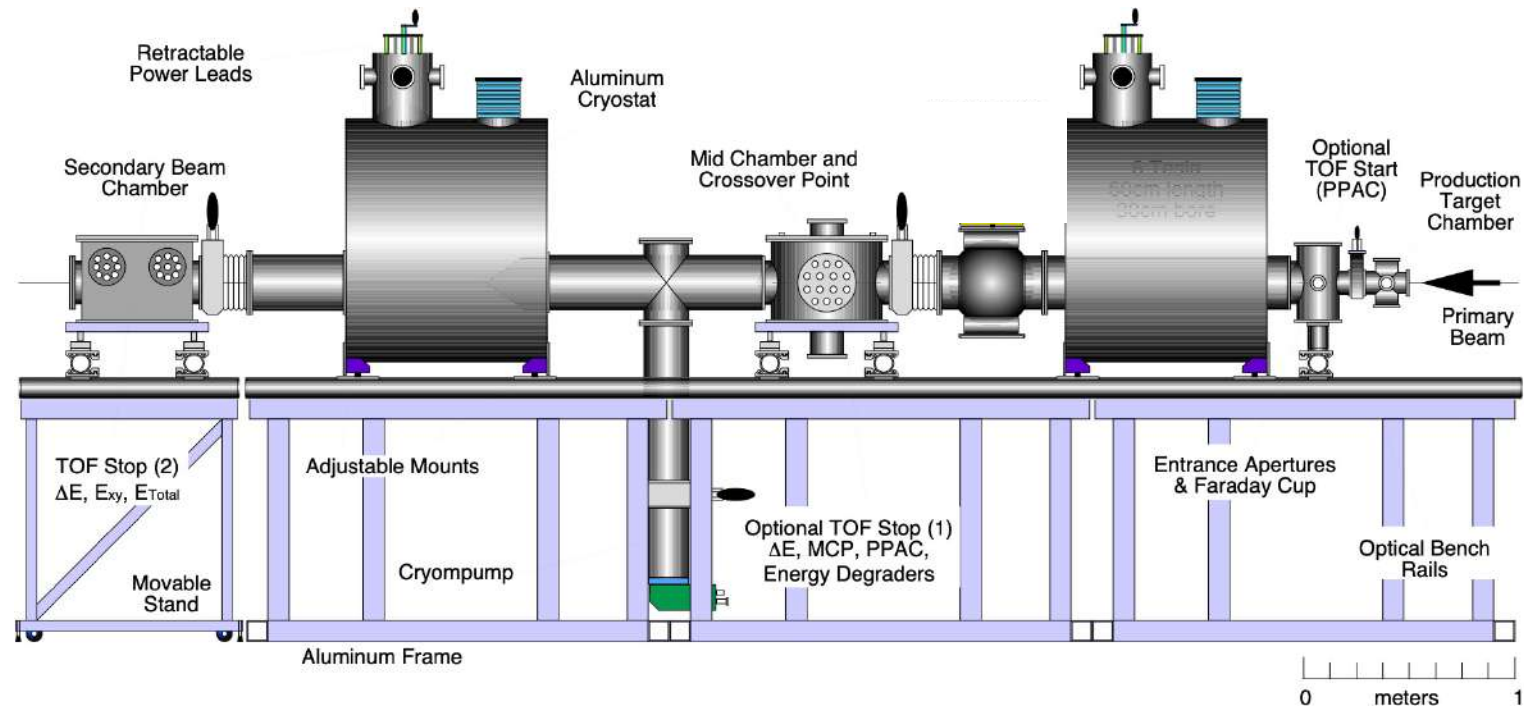
Gas catcher and cooler-buncher are completed.
RF carpet commissioning is presently underway.



N. Severjins *et al.*, PRL **63**, 1050 (1989)*

- *Also ^{17}F {but low precision}
- ^{17}F is one of the most plentiful *TwinSol* beams ($\sim 2 \times 10^6$ pps)
- This will likely be the first trap experiment.
- Next: measure lighter nuclei (^{11}C to ^{15}O) since they can also probe for scalar/tensor currents.

TwinSol (original and mods)

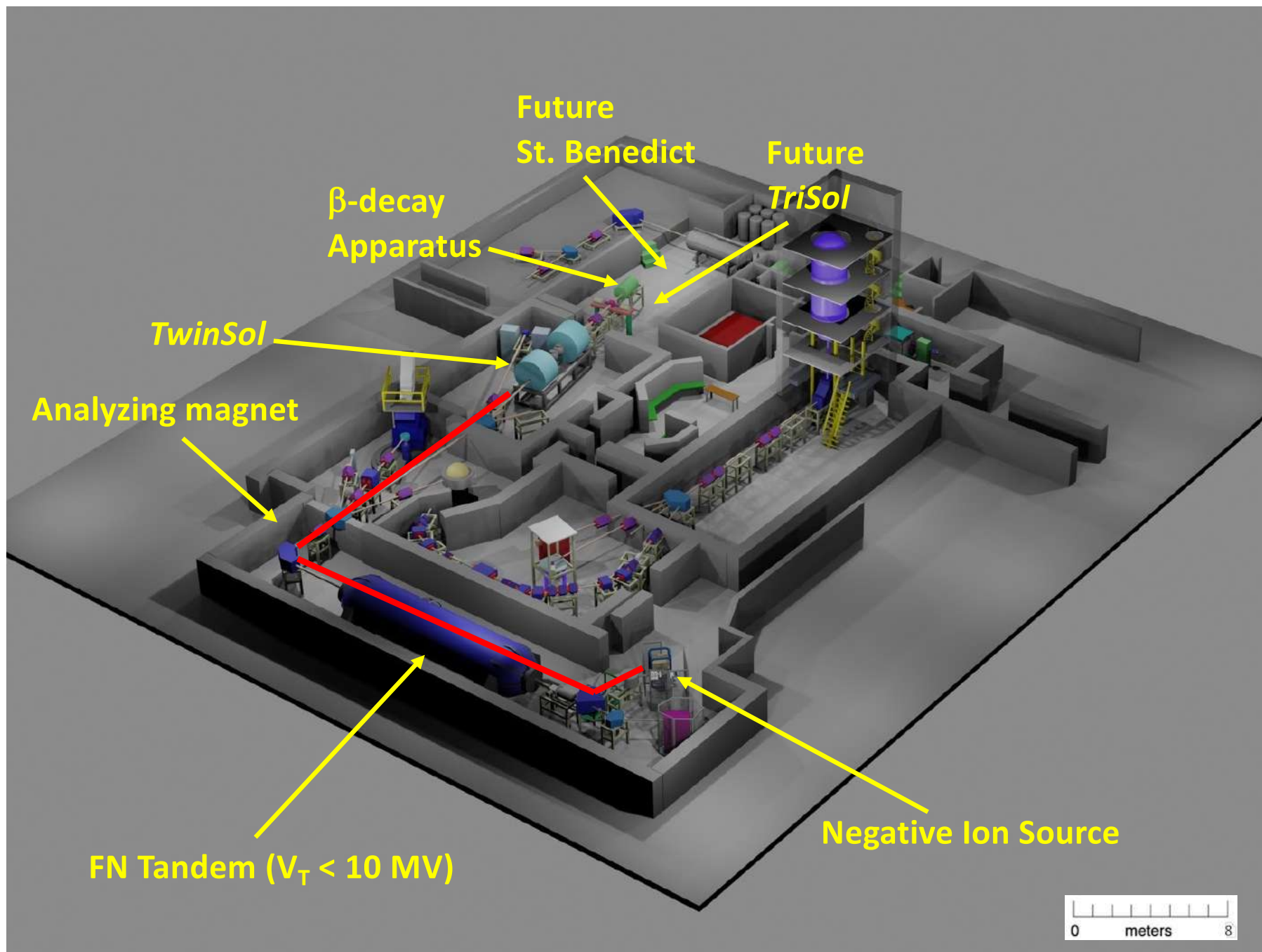


Modified Target Chamber



Gas Cells





LISE++ to Calculate Settings

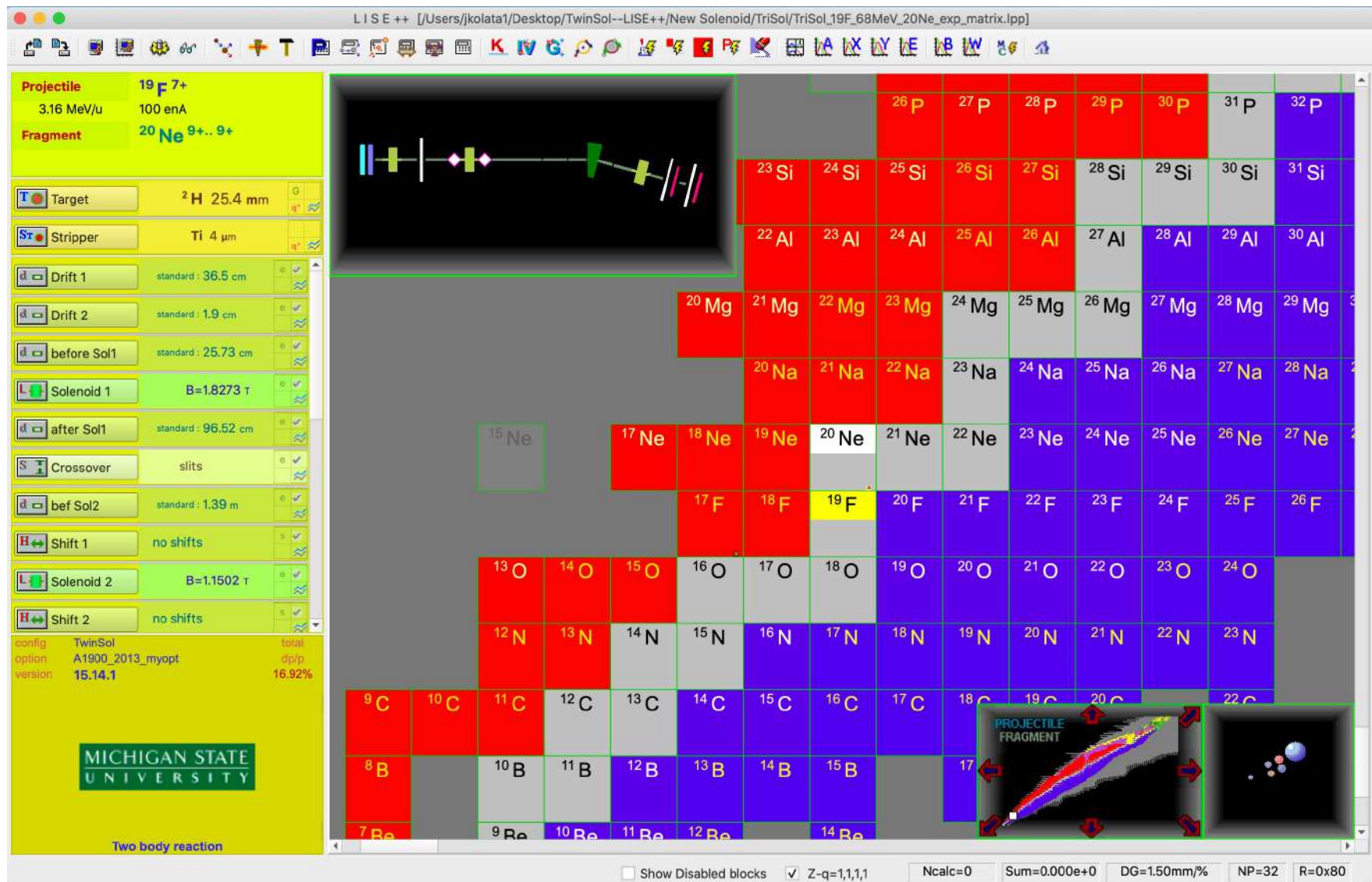
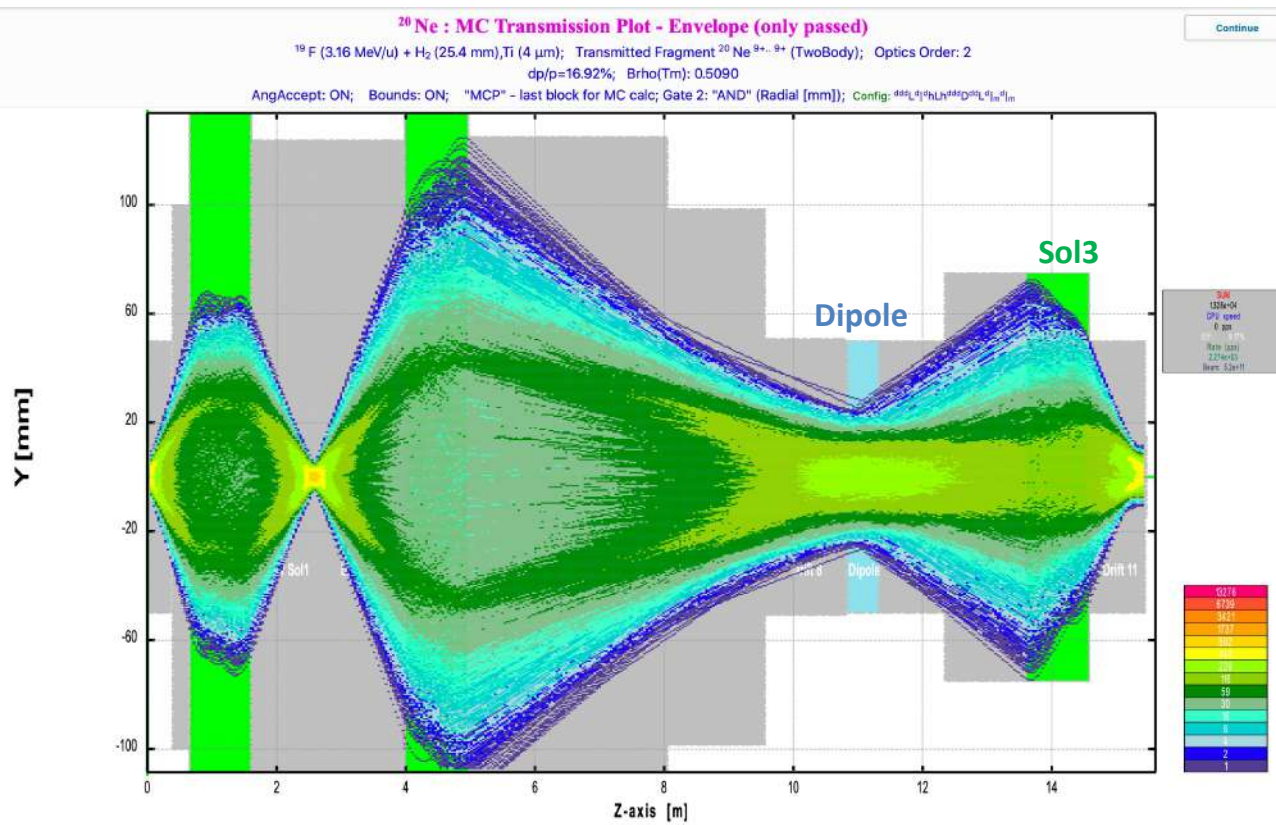


Illustration: ^{20}Ne production

Calculated Beam Envelope



Using the ¹⁹F(d,n) reaction at 68 MeV:

8.5 MV on the FN terminal will produce 50 MeV ²⁰Ne. Energies up to 60 MeV might be possible if we're willing to push the terminal voltage.

Calculated beam properties:

2.3x10³/s for 100 enA ¹⁹F (7⁺)
10 mm FWHM
ΔE = 1.65 MeV FWHM
±1.75° angle

Purity should be high since the (d,n)
Q-value is > +10 MeV. Only ¹⁷O has a
similar Q-value.

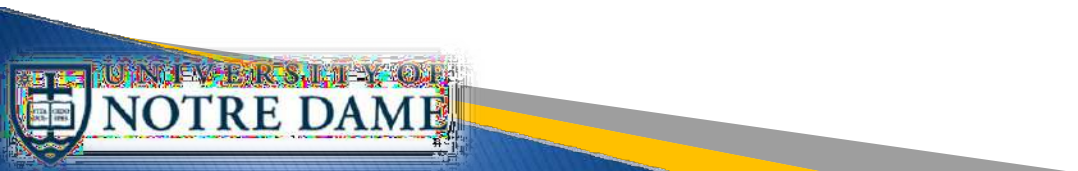
Why ^{20}Ne ?

Submission to FRIB PAC-1:

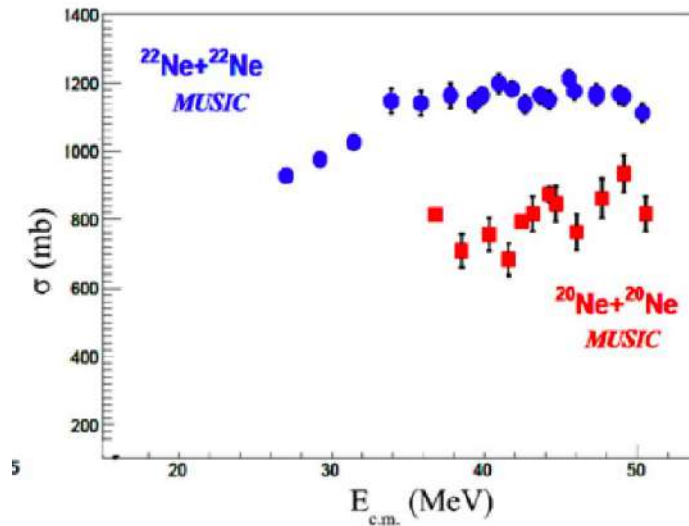
Measurement of $^{26}\text{Ne}+^{20}\text{Ne}$ fusion cross sections with the AT-TPC to constrain pycnonuclear burning in accreting neutron stars.

Jaspreet S. Randhawa, Rahul Jain
February 12, 2021

As part of this proposal, it was suggested that a comparison with the fusion of lighter systems, such as $^{20}\text{Ne}+^{20}\text{Ne}$, would be necessary.

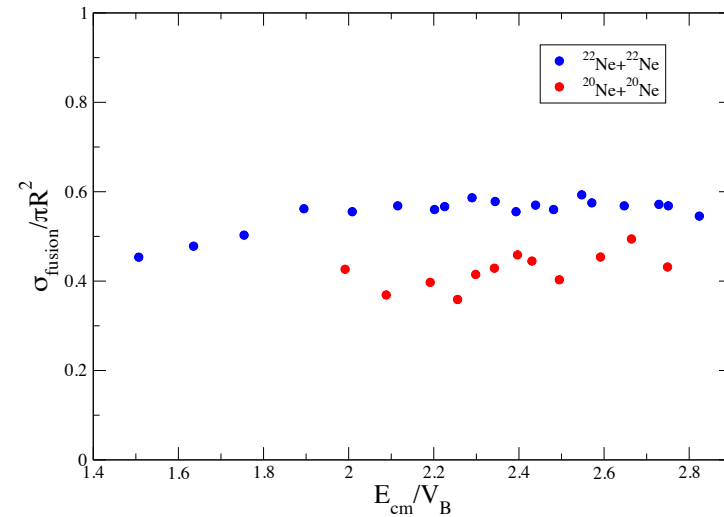


Recent Ne+Ne Fusion Data



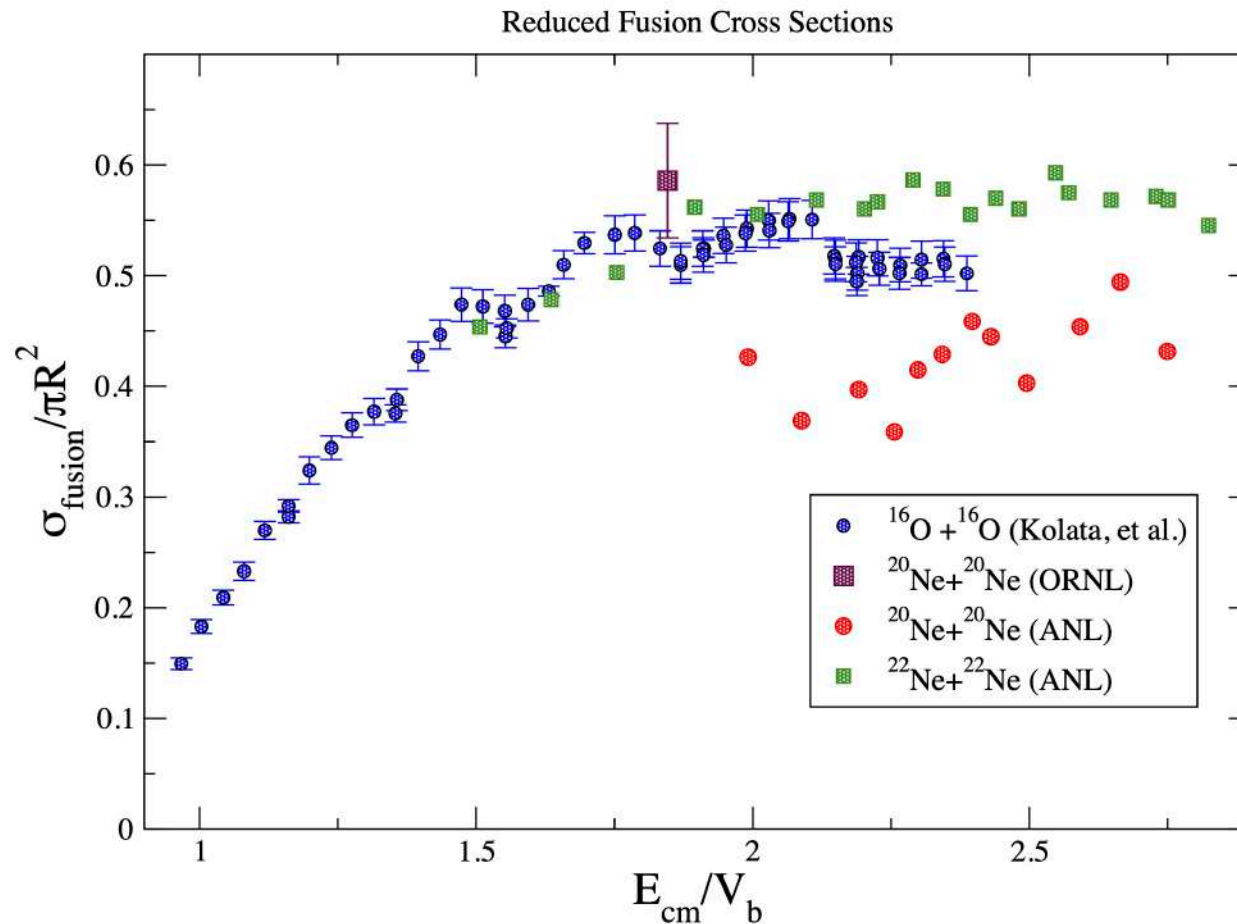
Avila, et al. ; MUSIC chamber
Argonne National Lab. [2016]
(Preliminary analysis.)

(Note oscillations in ^{20}Ne data)



“Reduced” data: corrected for
nuclear radius. The big cross
section difference remains.

Additional Fusion Data (Reduced)

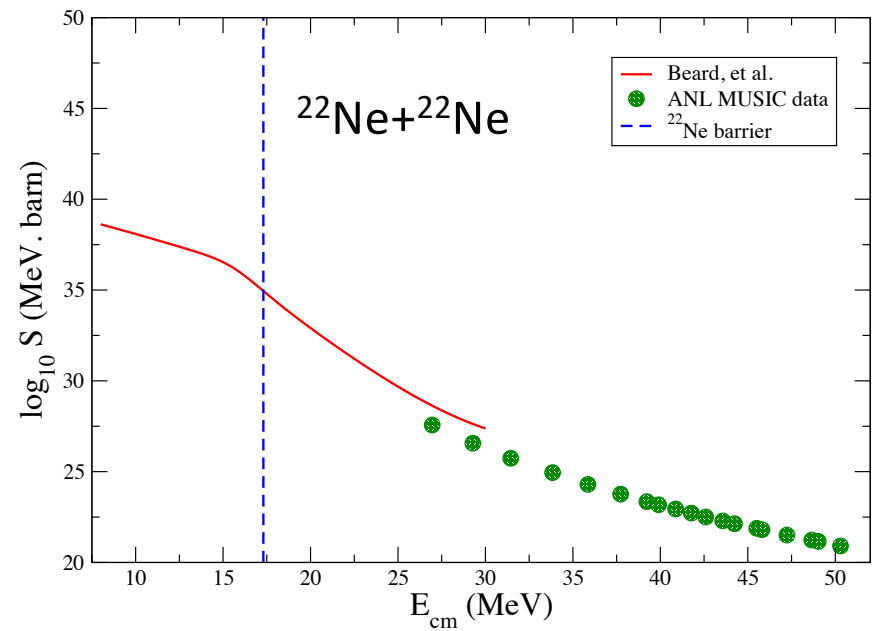
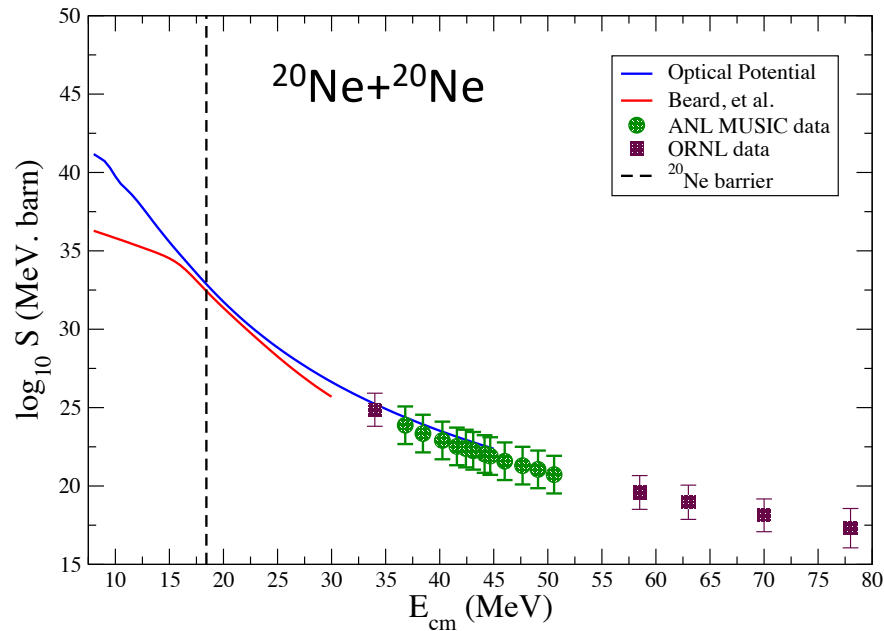


ORNL: PRC 28, 1148 [1983].
Kolata, *et al.*: PRC 16, 891 (1977).

Oscillations in $^{16}\text{O} + ^{16}\text{O}$ occur due to anti-symmetrization in this identical-particle system, which eliminates the odd partial waves. Does $^{20}\text{Ne} + ^{20}\text{Ne}$ show this effect?

The observed absolute magnitude of the ANL $^{20}\text{Ne} + ^{20}\text{Ne}$ data disagrees substantially with the additional data. The big effect of neutron excess disappears. Which is correct?

Astrophysical S-factor

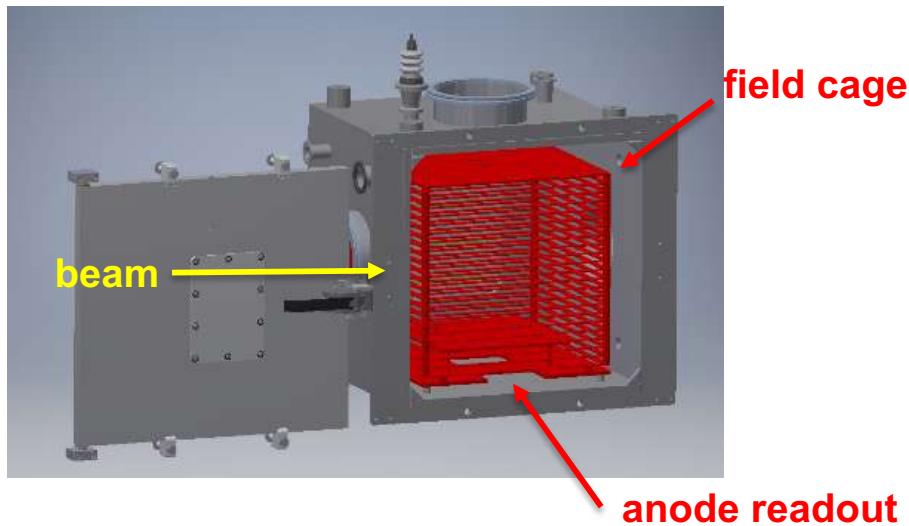


S-factor predictions from M. Beard, *et al.*, At. Data Nucl. Data Tables 96, 541 (2010) using the Sao Paulo potential. (These predictions, based on a polynomial approximation, become increasingly unrealistic above $E_{cm} = 30$ MeV).

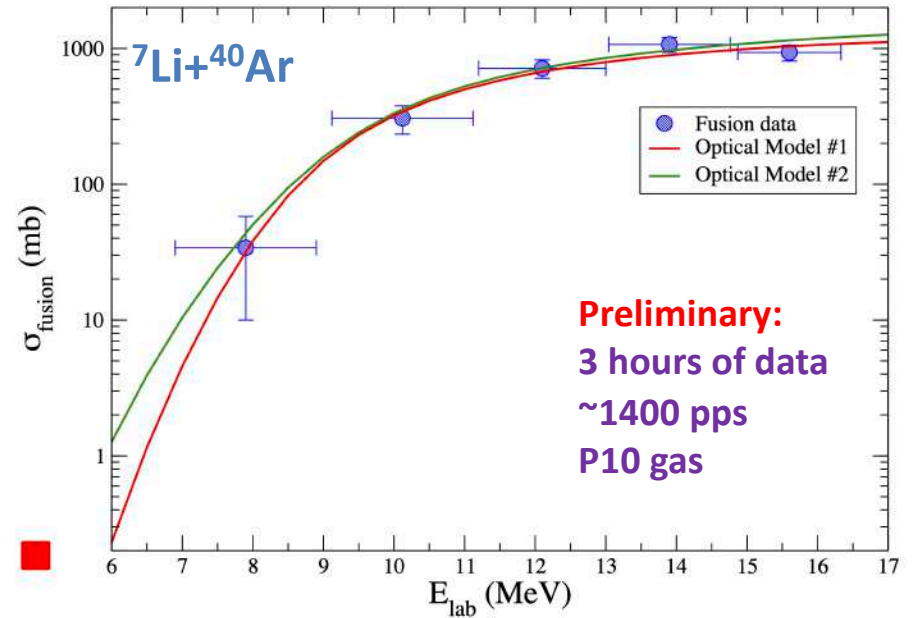
Note that these predictions actually suggest that the $^{20}\text{Ne}+^{20}\text{Ne}$ experimental cross sections are too high while the $^{22}\text{Ne}+^{22}\text{Ne}$ cross sections are too low, i.e., the observed neutron-excess effect may even have been under-estimated.

An experiment starting at $E_{cm} = 25$ MeV should easily reach the region below 16 MeV where there is increasing disagreement between Beard, et al. and optical-model predictions.

ND-CUBE



- Rectangular geometry ; perpendicular drift
- 40 cm x 40 cm x 40 cm chamber, up to 1 atm
- Micropattern gas detector (MPGD)
- 20 cm x 30 cm, active area ; **1000 anode pads**
- Micromegas with thick GEM amplification



Optical Models:

#1: K.W. Potthast, et al., NPA614, 95 (1997)

#2: J. Cook, et al., NPA388, 153 (1982)

Collaborators

Weak Interaction Physics

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Patrick O'Malley

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Patrick O'Malley
Jaspreet Randhawa

Sydney Coll (undergrad)

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