

# Nuclear Science at the University of Notre Dame



Patrick O'Malley  
on behalf of the Nuclear Science Laboratory  
University of Notre Dame



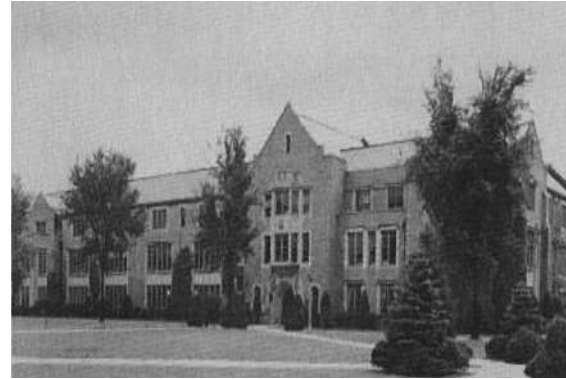
HINPw6 , May 14, 2021



# Historical Development of NSL



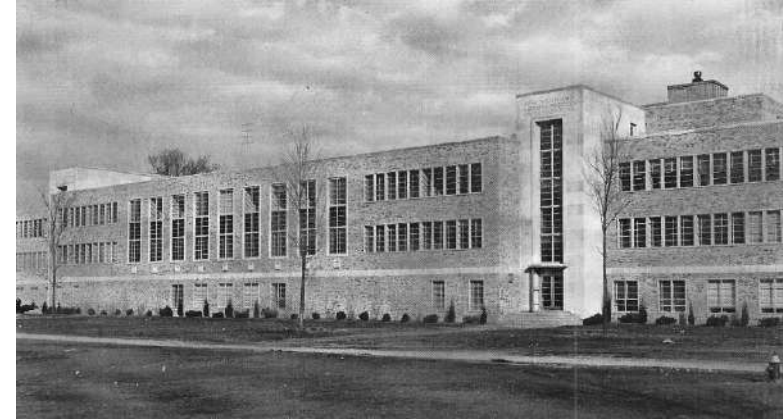
1936-1938 Construction of 1<sup>st</sup> Accelerator  
Home built 1 MV open air machine for electron beams.



1940- 1942 Construction of 2<sup>nd</sup> Accelerator  
a home built 8MV (4MV) VDG machine. It served the Manhattan (Hanford) project until 1950.



# 1953 Nieuwland Science Hall 1968 NSL Tandem Accelerator



Continuous funding of Nuclear Physics at Notre Dame since 1946 through the Office of Naval Research and from 1953 by the National Science Foundation founded in 1950.

The first accelerator in Nieuwland Science Hall was a 4 MV machine, complemented in 1968 by an FN tandem from HV Engineering (\$2.5M). The construction started in 1966, the building was complete in winter 1967, the accelerator was delivered in early 1968. Operation started in 1969. Upgrade to Pelletron system in 1995.



# Can never have enough accelerators!



Santa Ana (5U) accelerator



Sept 2016



Oct 2016



9S First Beam  
Dec 2017



# Can never have enough accelerators!



Santa Ana (5U) accelerator



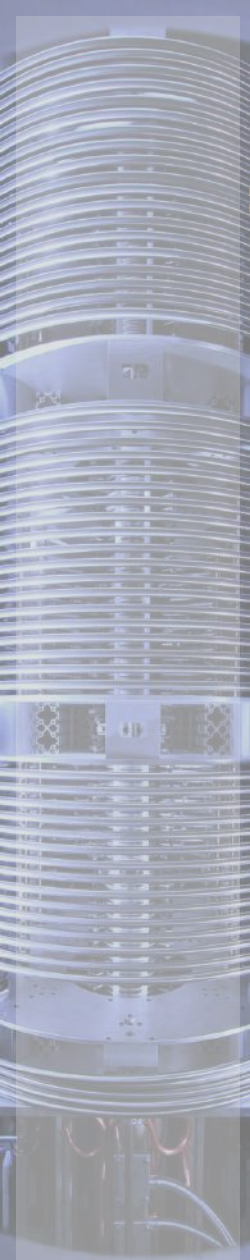
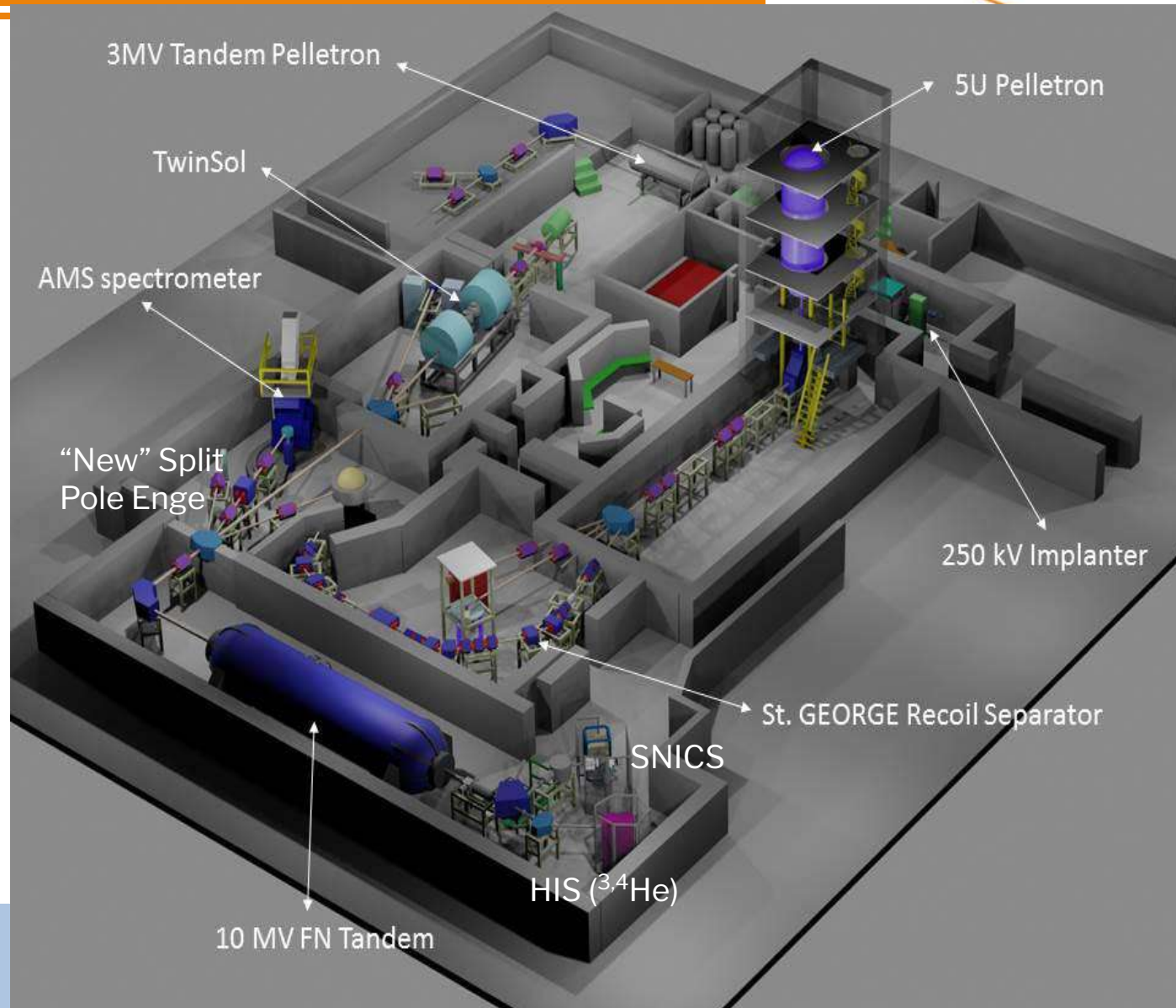
Oct 2016



9S First Beam  
Dec 2017



# Nuclear Science Laboratory



# Solid Target Line



- (a,n) cross sections on light elements are needed for a wide range of applications
- Past measurements are nearly all total cross sections. Partial cross sections are needed!
- A cost effective and efficient solution for (a,n) measurements
  - Deuterated liquid scintillators for prompt neutron detection
  - HPGe or LaBr<sub>3</sub> for secondary gamma ray detection



James DeBoer  
R-matrix  
development



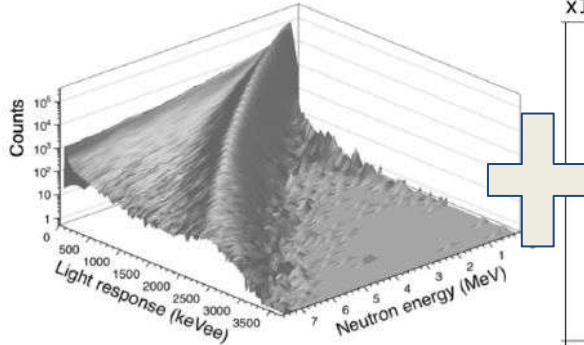
Michael  
Wiescher,  
Nuclear  
Astrophysics &  
Nuclear  
Applications



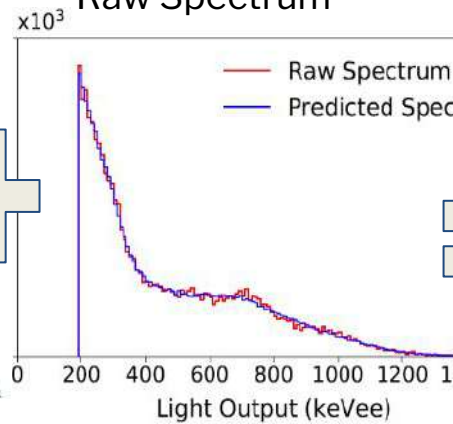
# ORNL Deuterated Spectroscopic Array - ODeSA



Light Response Matrix



Raw Spectrum



neutron energy spectrum

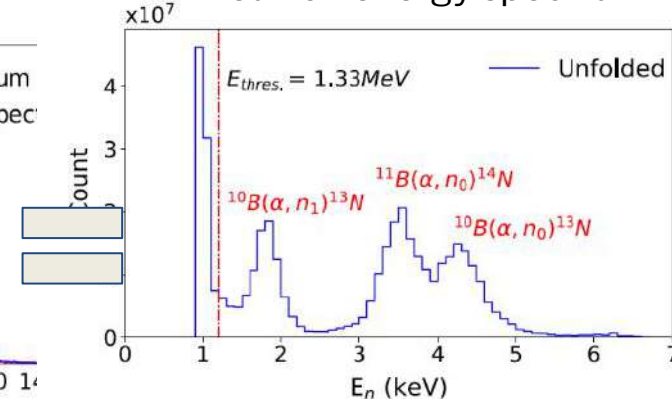
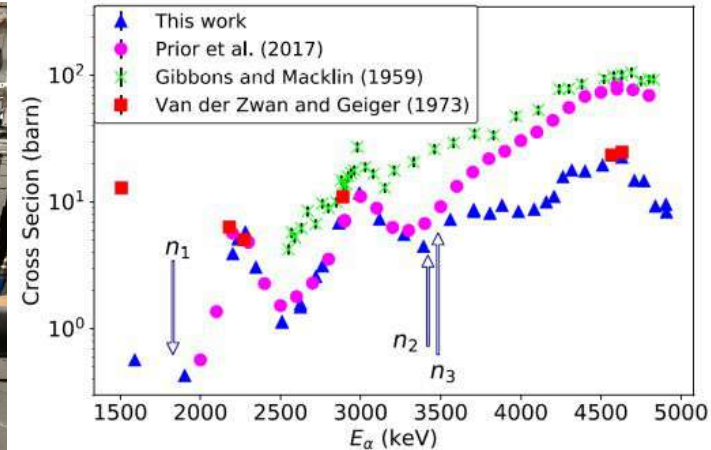
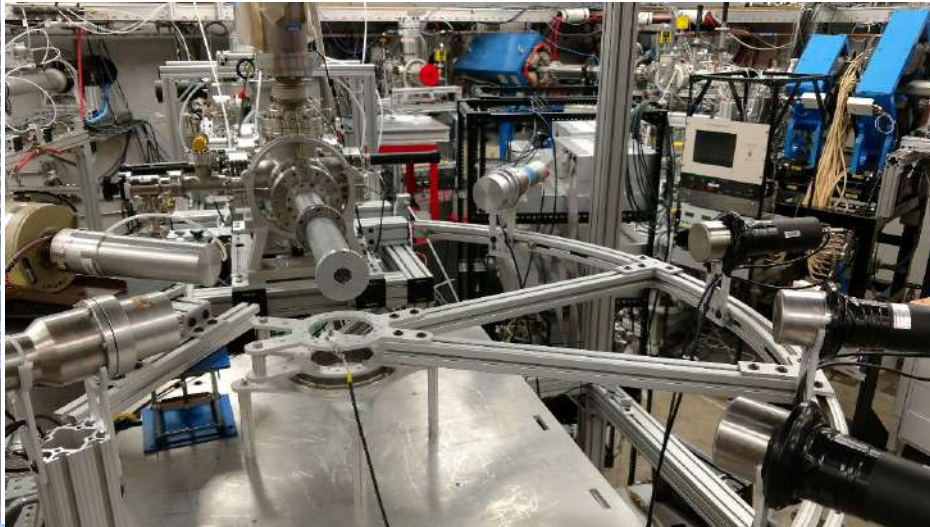


Fig. 9. Response matrix generated using a broad energy neutron source from a thick target  $^{27}\text{Al}(d, n)$  reaction at  $E_d = 7.44$  MeV [12].



Febbraro et al. (2019)

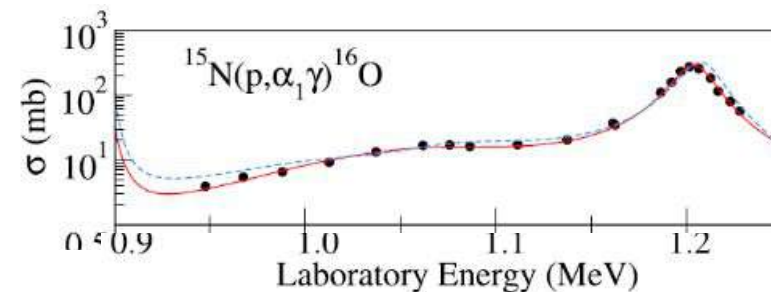
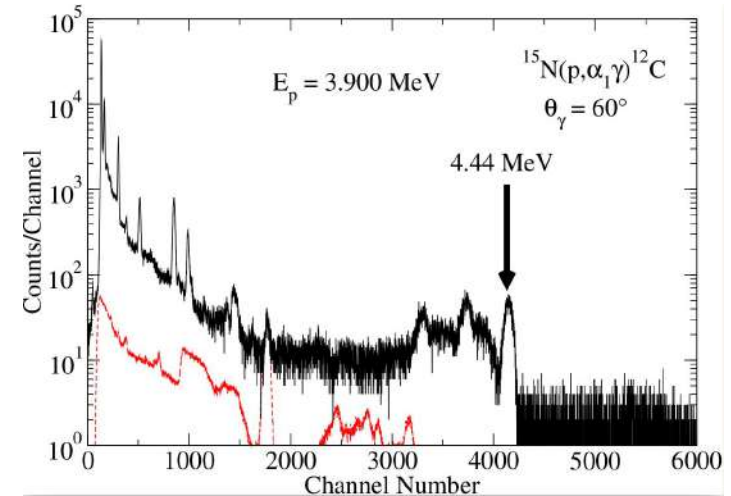
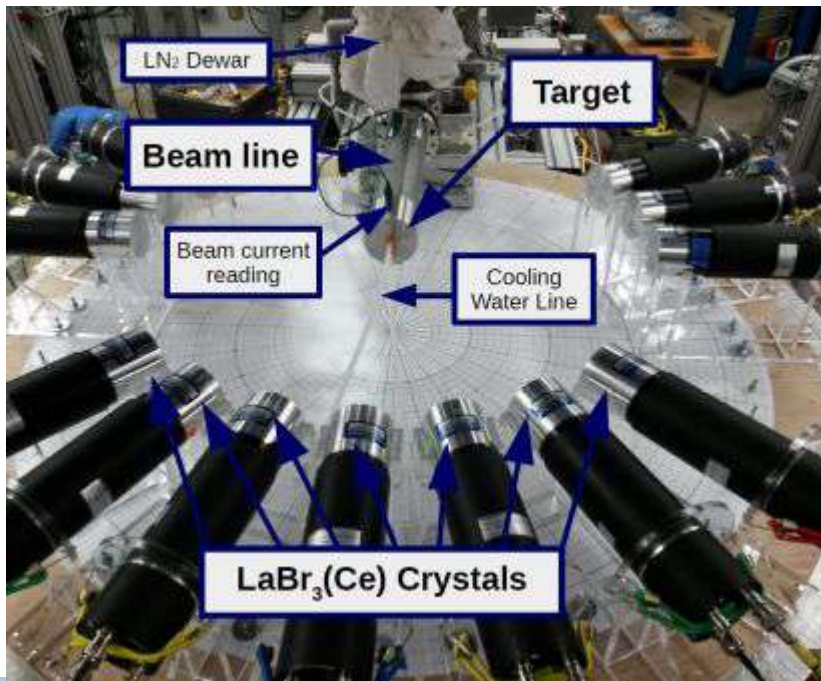




# HAGRiD



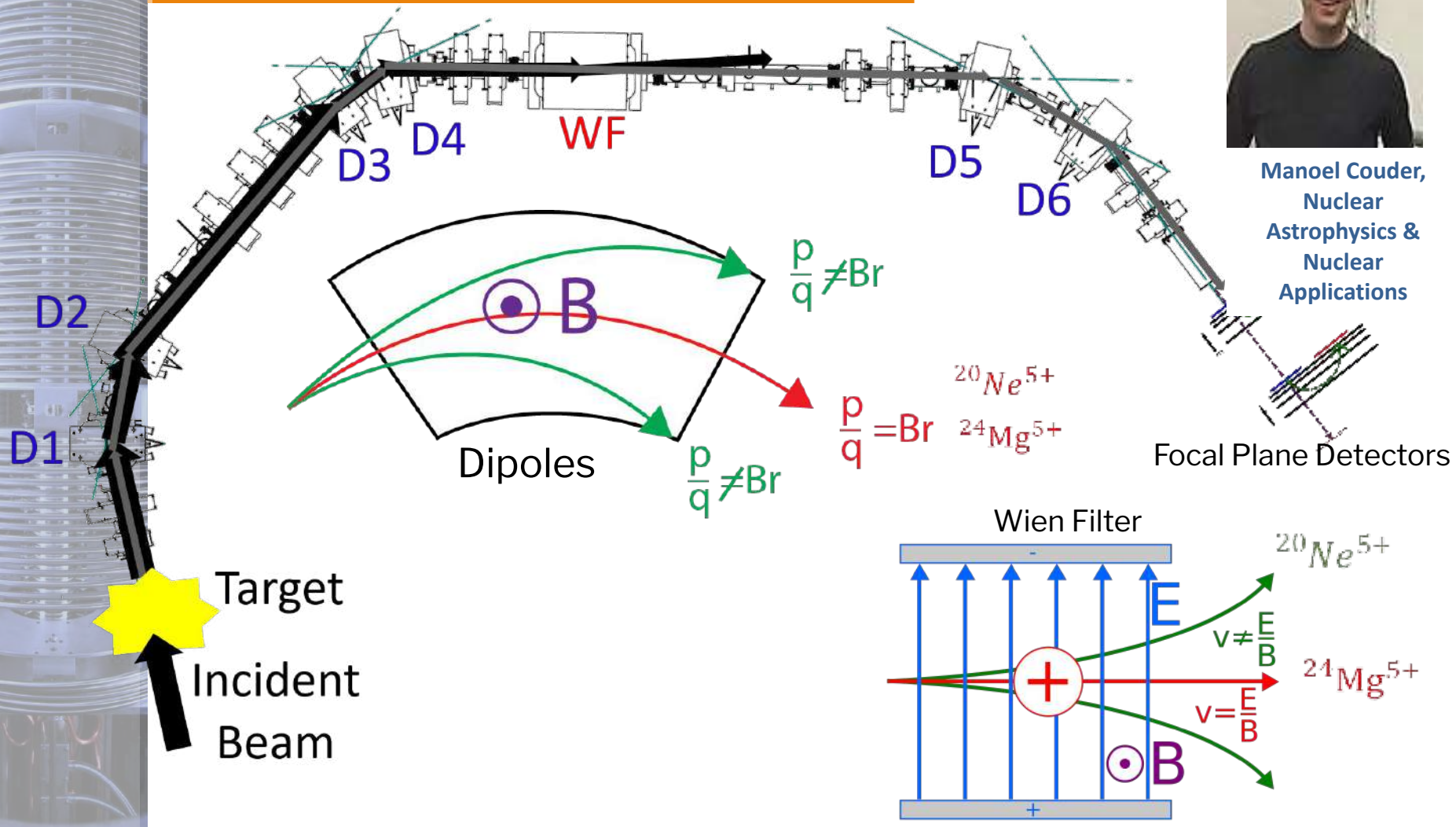
- HAGRiD is an array of LaBr<sub>3</sub> detectors from Kate Jones at UTK
  - relatively high efficiency at high gamma ray energies
- Future (α,nγ) measurements include
  - <sup>13</sup>C(α,nγ)<sup>16</sup>O
  - <sup>17</sup>O(α,nγ)<sup>20</sup>Ne
  - <sup>18</sup>O(α,nγ)<sup>21</sup>Ne
  - <sup>25</sup>Mg(α,nγ)<sup>28</sup>Si

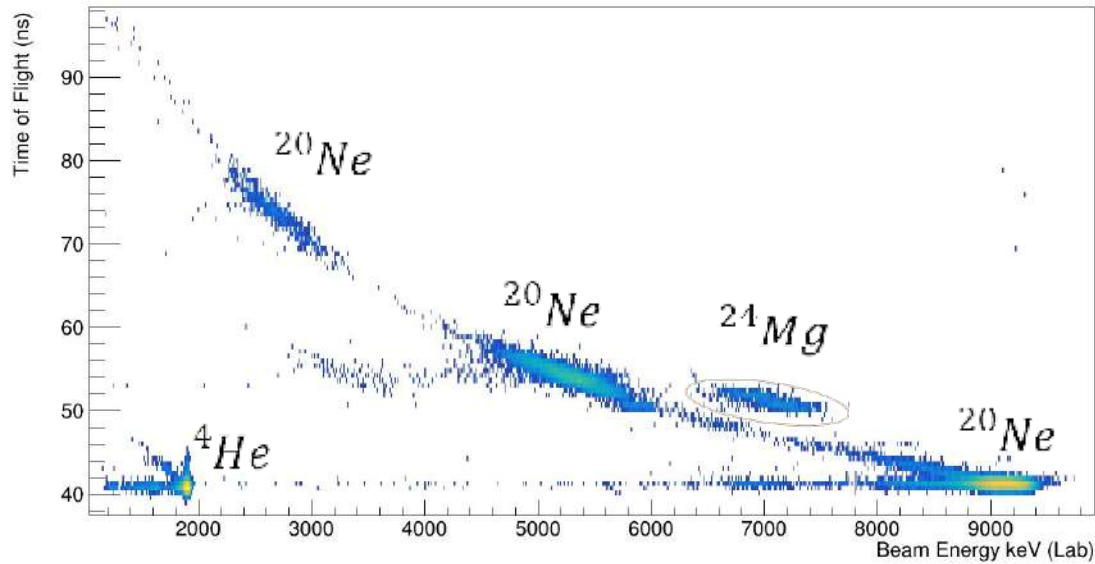
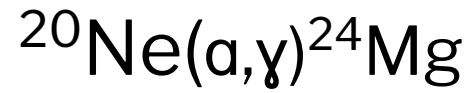


# Inverse Kinematics with St. George



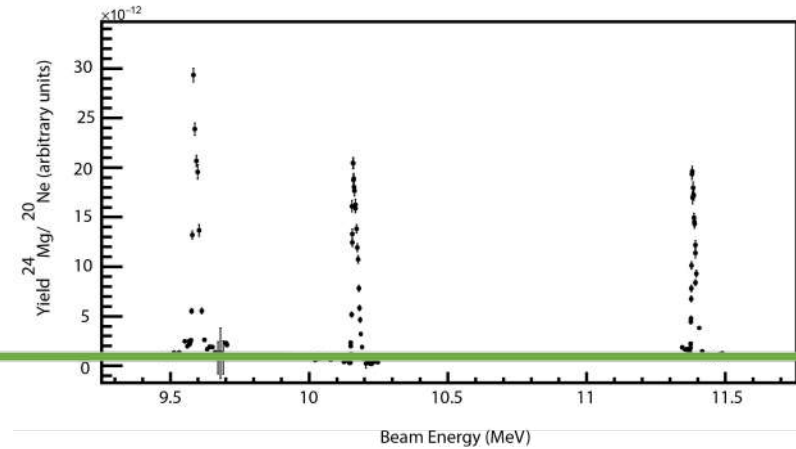
Manoel Couder,  
Nuclear  
Astrophysics &  
Nuclear  
Applications



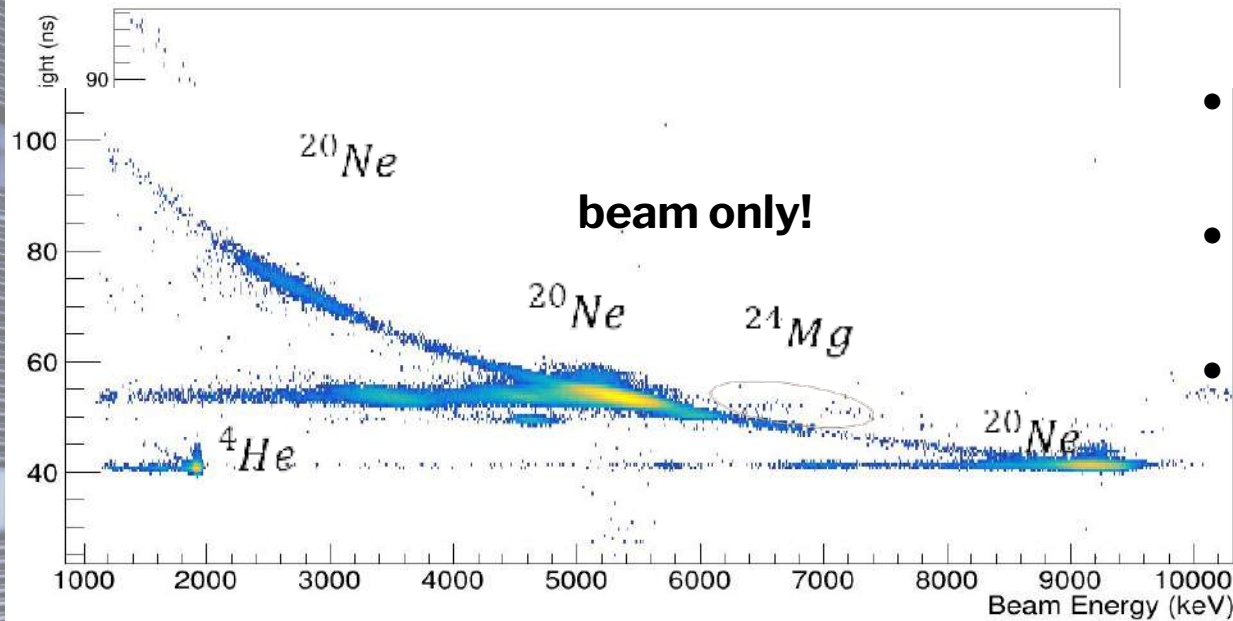


- Commissioning run for St. George run by student **Luis Morales**
- identified a background from the beam
- installing new Wien filter on upstream beamline

Beam Contamination  $^{24}\text{Mg}$  background

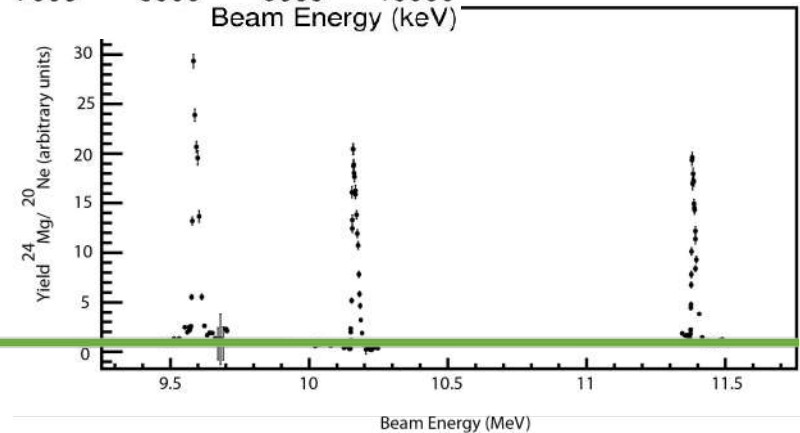


# $^{20}\text{Ne}(\alpha,\gamma)^{24}\text{Mg}$



- Commissioning run for St. George run by student **Luis Morales**
- identified a background from the beam
- installing new Wien filter on upstream beamline

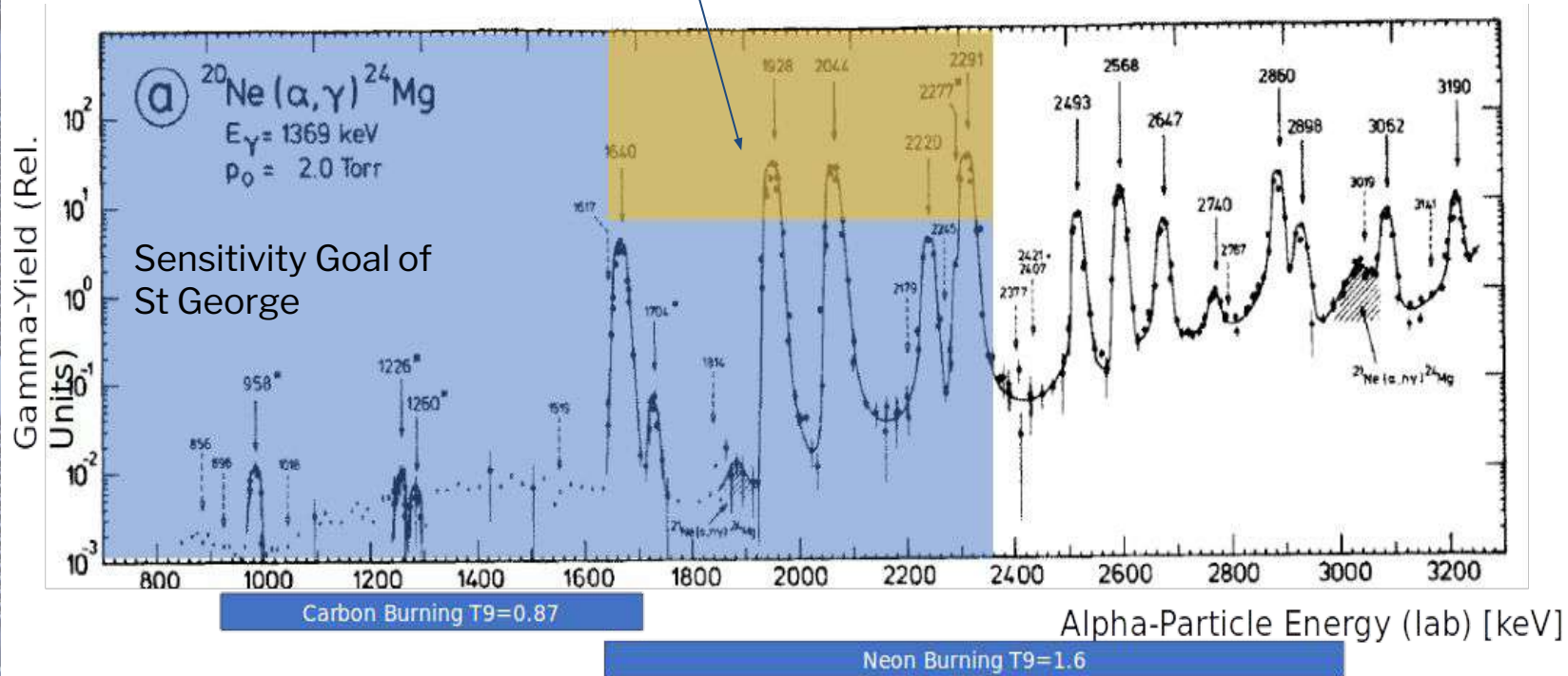
Beam Contamination  $^{24}\text{Mg}$  background



# Status of St. George



Current sensitivity



P. Schmalbrock, Stellar reaction rate of  $^{20}\text{Ne}(\alpha, \gamma)^{24}\text{Mg}$ , 1983 [https://doi.org/10.1016/0375-9474\(83\)90488-8](https://doi.org/10.1016/0375-9474(83)90488-8)



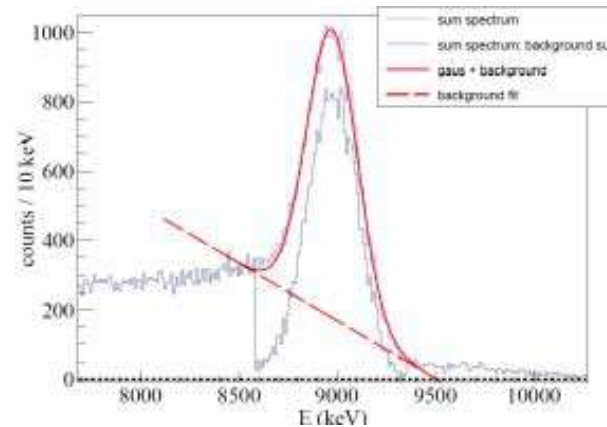
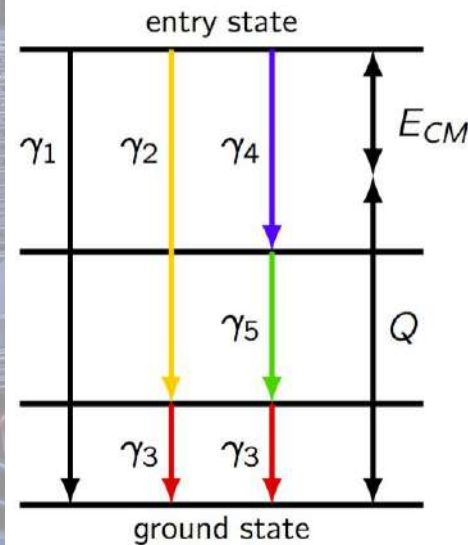
# HECTOR



- High Efficiency Total absorption spectrometer
- 16 separate segments, each read by two PMTs
- crystal size: 4x8x8 inch,
- 1 mm Al casing surrounding each crystal
- 12 mm Al on the outer walls of the array
- 60 mm borehole



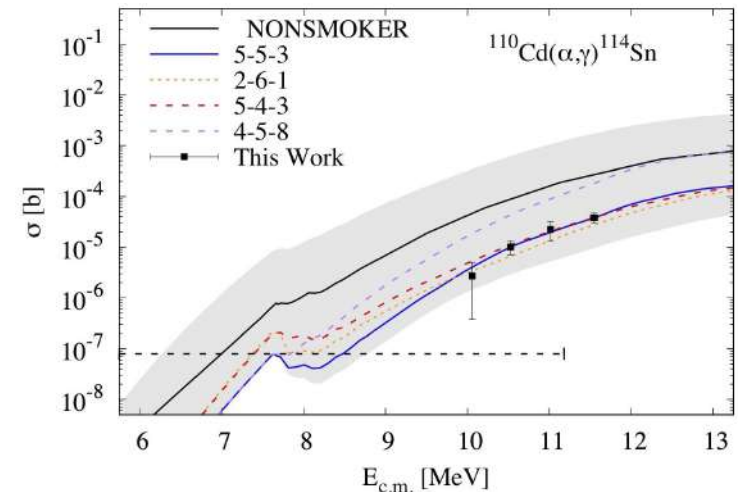
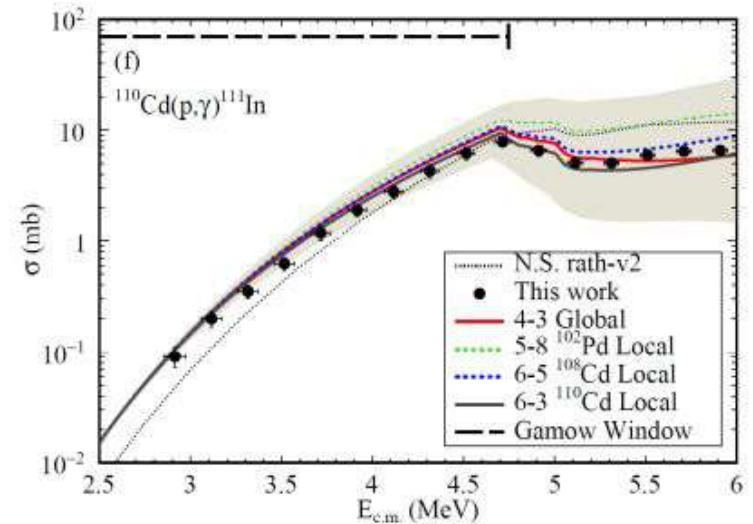
**Anna Simon,**  
Nuclear  
Astrophysics &  
Nuclear  
Reactions



# HECTOR



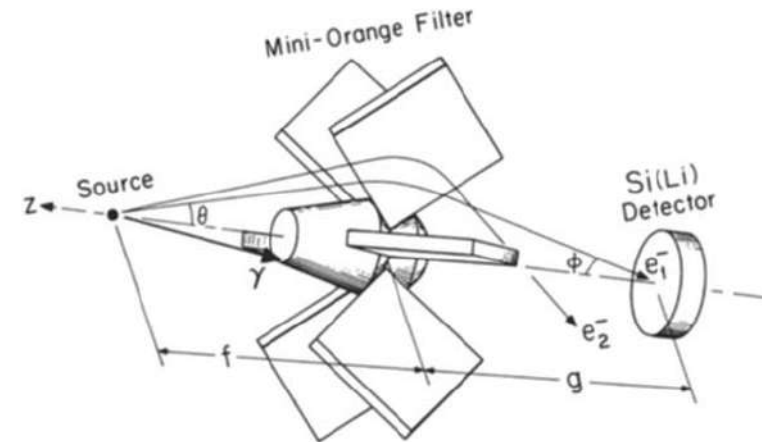
- $^{111}\text{In}$  had a poorly-constrained temperature of the  $(\gamma,p)/(\gamma,n)$  branching point, where the  $(\gamma,p)$  becomes dominant over the  $(\gamma,n)$  reaction.
- Constrained the  $(\gamma,p)/(\gamma,n)$  branching point to  $2.71 \pm 0.05$  GK.
  - This may have a significant impact on the predicted p-nuclei abundances.
- The reaction flow in the astrophysical  $\gamma$ -process, the  $(\alpha,\gamma)$  cross sections were measured for  $^{90}\text{Zr}$ ,  $^{102}\text{Pd}$ ,  $^{108}\text{Cd}$ , and  $^{110}\text{Cd}$
- Goal is to pinpoint the nuclei at which the  $(\gamma,\alpha)$  become dominant over the  $(\gamma,n)$  reactions and redirect the nucleosynthesis path towards lighter nuclei
- As a result a new branching point at  $^{114}\text{Sn}$  was identified for  $T9 < 1.7$  GK.



# Putting ICEBALL to the Fire



- ICEBall consists of six annular Si(Li) detectors that are 5 mm thick and have an active area of  $750 \text{ mm}^2$  each.
- ICEBall has a measured absolute efficiency of 15% at 356 keV and a resolution of 3-5 keV FWHM for electrons.



## **fIREBall ('fabulous' Internal converSion Electron Ball Array) is funded**

- New Si(Li) detectors
- New designs of magnetic filters
- Modern digitizer readout
- Modern Geant4/COMSOL simulations
- BGO shields coupled with HPGe detectors (Georgina)



Ani Aprahamian  
Nuclear  
Astrophysics &  
Nuclear Structure  
Physics

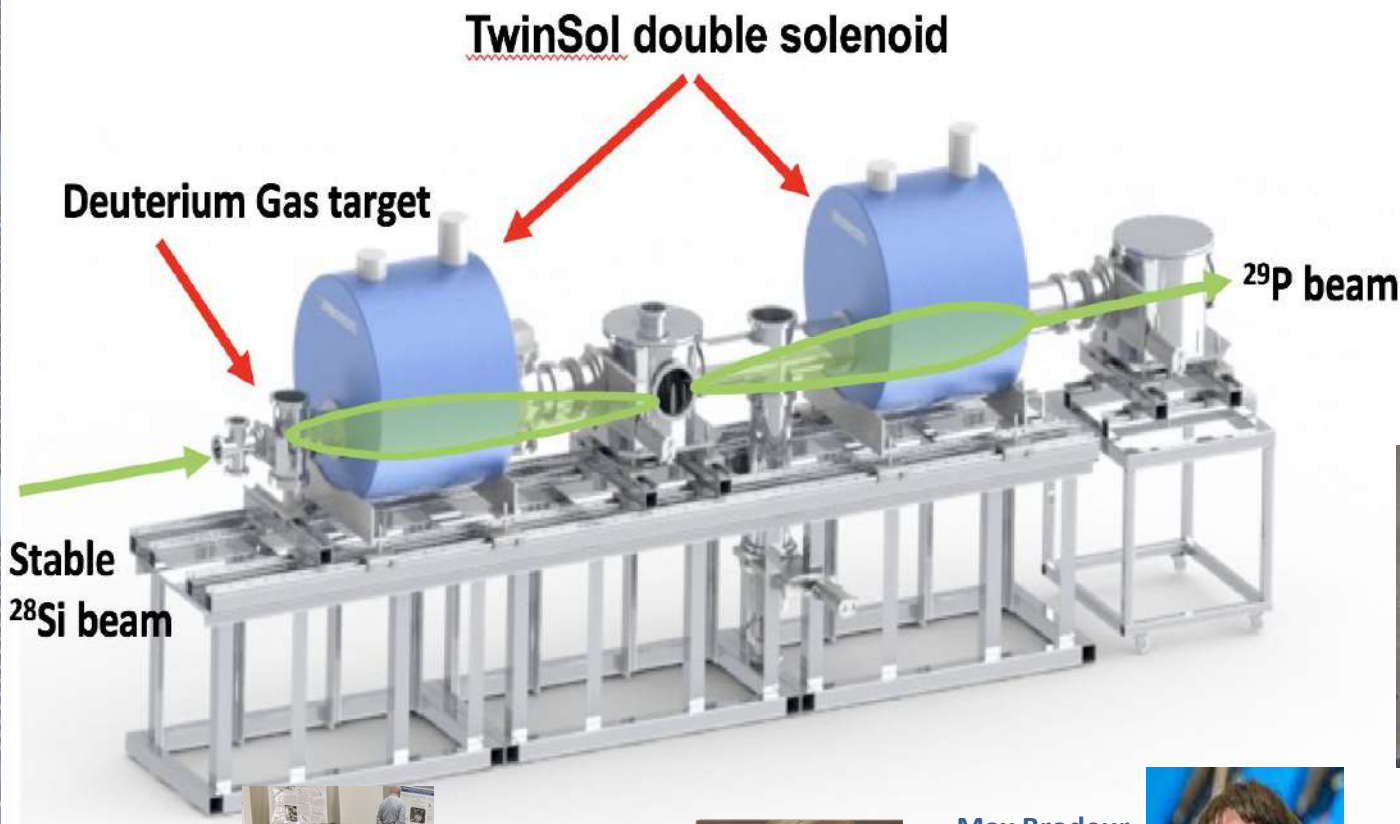


Wanpeng  
Tan  
Nuclear  
Structure  
and  
User  
Support

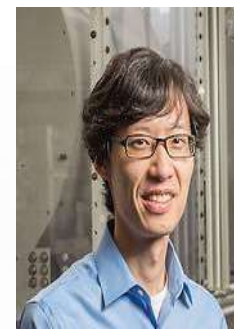




# The Sol-ful Folk



**Dan Bardayan,**  
Nuclear  
Astrophysics



**Tan Ahn,**  
Nuclear  
Astrophysics &  
Nuclear  
Structure

**Patrick O'Malley**  
TwinSol/St.  
Benedict



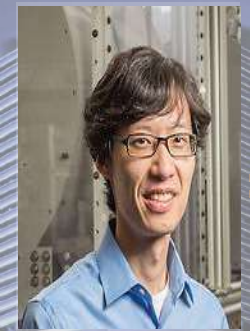
**James Kolata,**  
Fusion Barrier  
Studies  
Emeritus Faculty



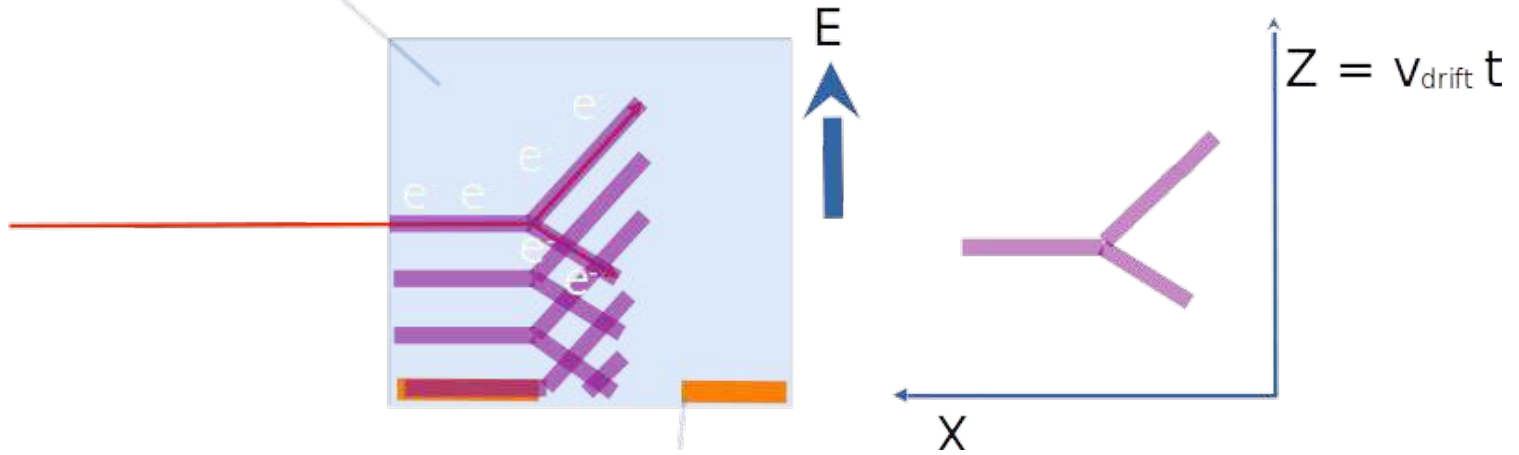
**Max Brodeur,**  
Nuclear  
Astrophysics &  
Fundamental  
Symmetries



# ND-CUBE- Active Target TPC

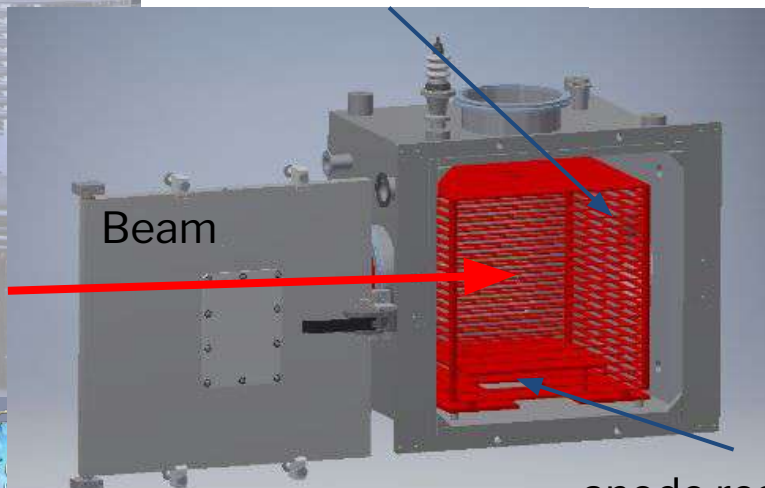


tracking medium

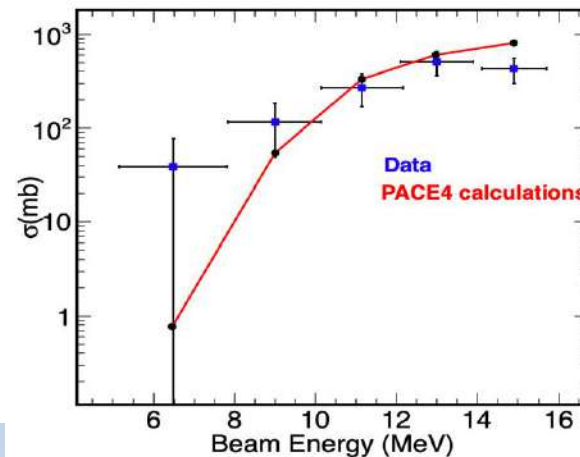


electron amplification/readout

Field cage



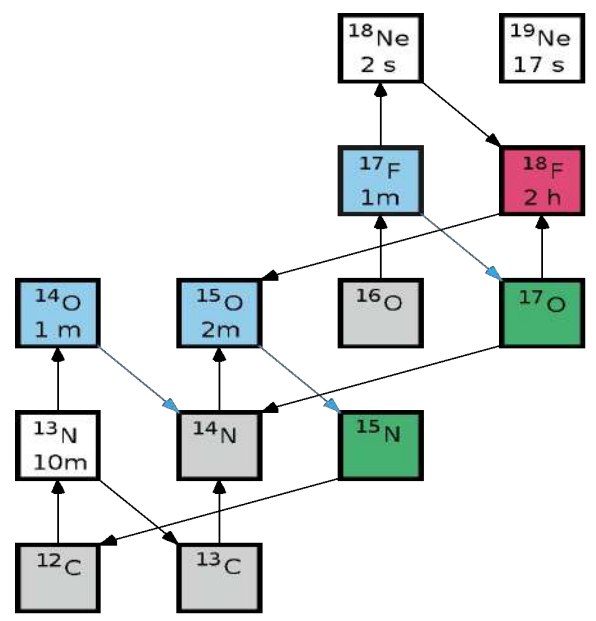
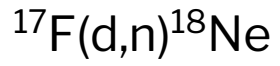
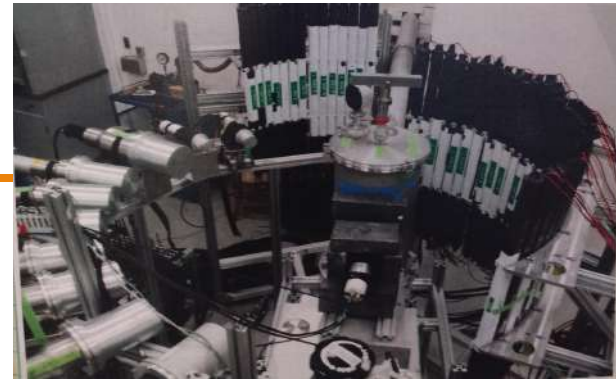
One hour data



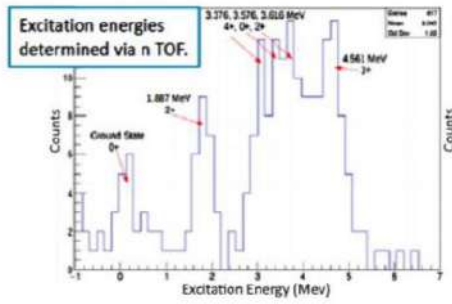
${}^7\text{Li} + {}^{40}\text{Ar}$   
fusion cross  
section for  
commissioning  
run



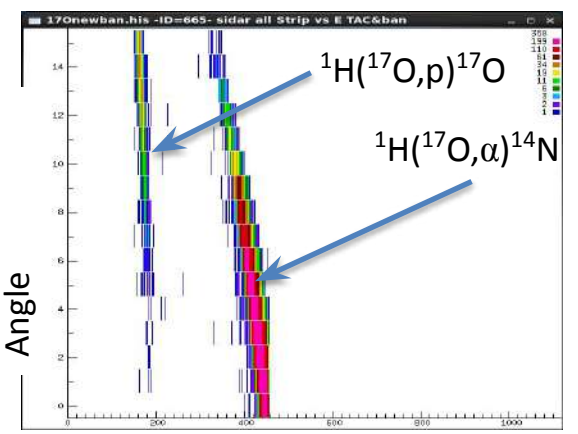
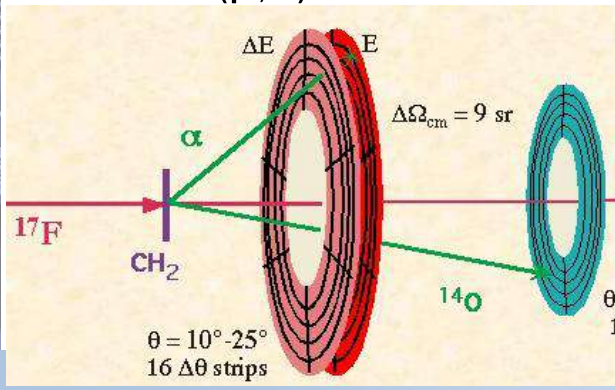
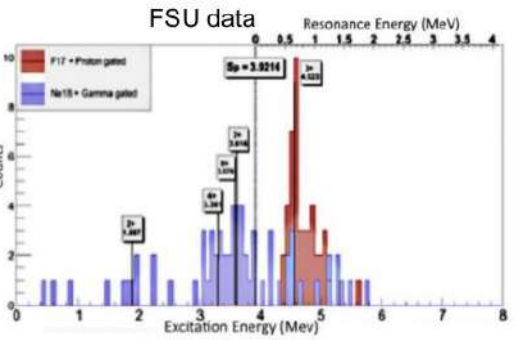
# Origins of $^{18}\text{F}$



Notre Dame data



FSU data



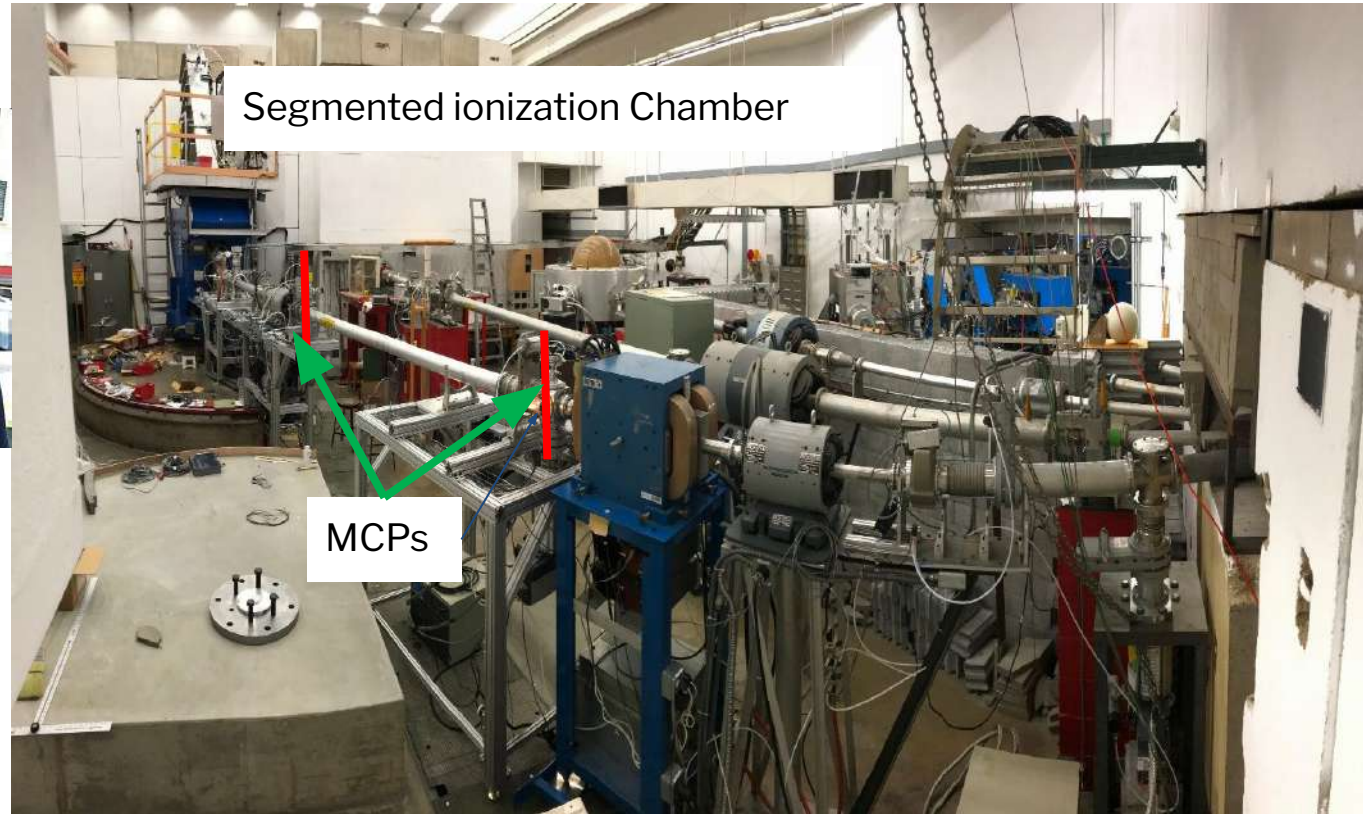
Energy Deposited in Large Detector



# Atomic Mass Spectroscopy



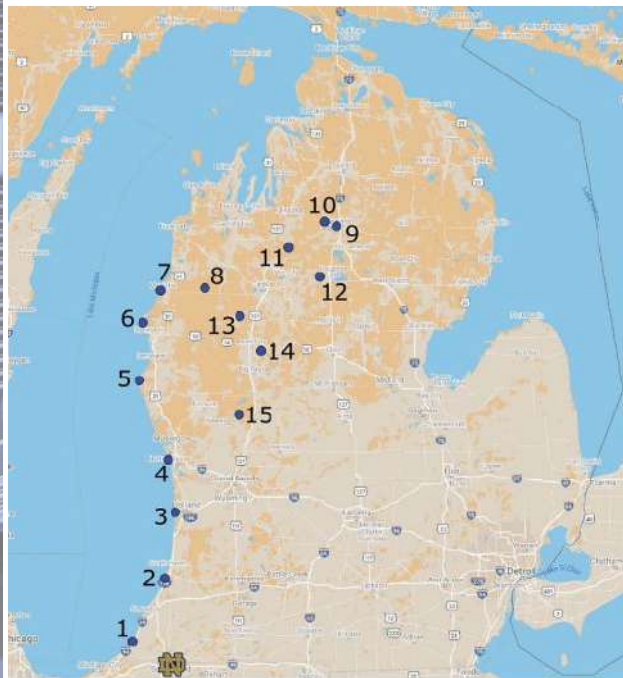
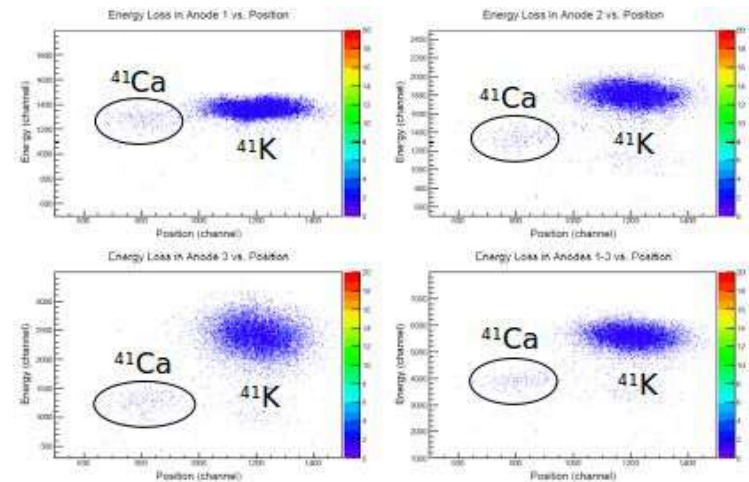
**Philippe Collon,  
Nuclear  
Astrophysics &  
Nuclear  
Applications**



# Range of Applications of AMS



- With a half life of  $9.94 \times 10^4$  yrs,  $^{41}\text{Ca}$  is a useful chronometer for studying the formation of the Early Solar System
- Currently being used to study the cross section for the  $^{40}\text{Ca}(^3\text{He},d)^{41}\text{Ca}$  reaction



- $^{129}\text{I}$  concentrations have been measured for environmental sampling for  $^{129}\text{I}$  contamination in the surrounding Great Lakes region
- Est. Sensitivity limit of  $\sim 3.7 \times 10^{-14}$
- Development towards high sensitivity actinide detection is underway  $^{236}\text{U}/\text{U}$  concentrations from various ore materials



# PIGE Analysis for $^{19}\text{F}$



Use PIGE as a rapid screening tool for chemicals of concern in the environment, consumer products...

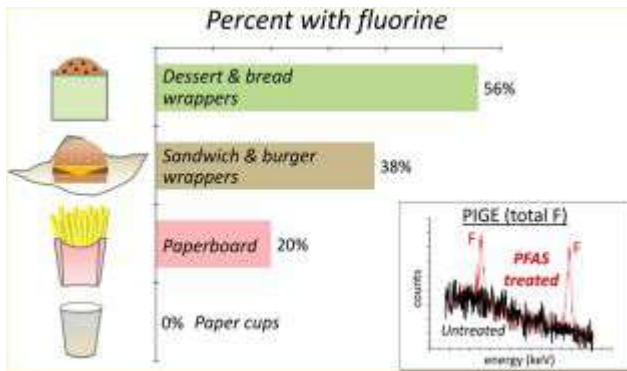
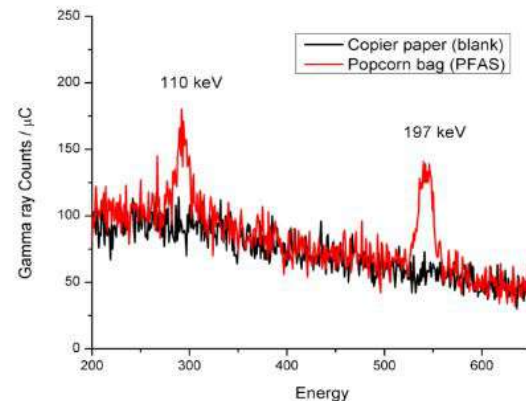
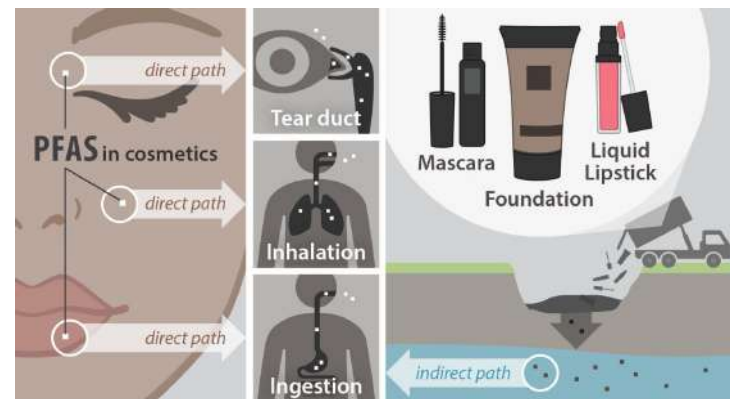
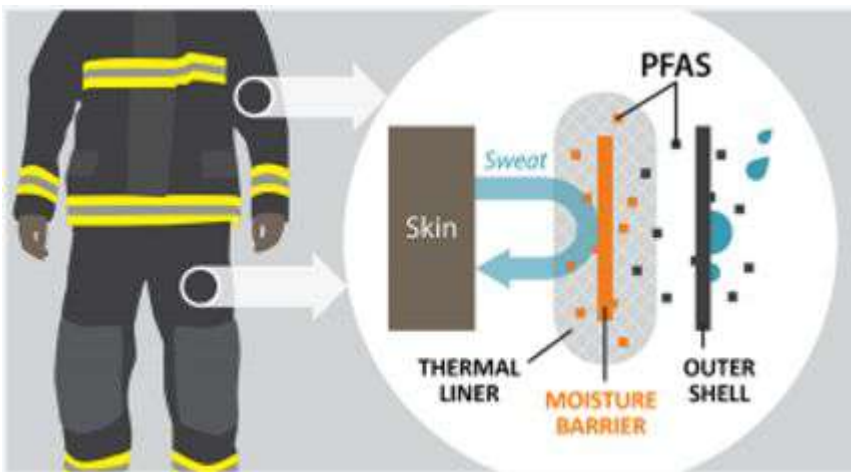


Fig. 3: PFAS-coated paper sample compared with uncoated paper. Irradiation time of 180 second with 9 nA of 3.4 MeV protons.



Graham Peaslee, Nuclear Applications



# But wait, there's more!



- There are many more things happening at Notre Dame
  - SF<sub>6</sub> upgrade of FN tandem
  - TRISOL
  - St. Benedict
  - Enge Split-pole
  - Actinide Target Program



NSF grant (PHY-0959816 & PHY-1062819)



# Questions???

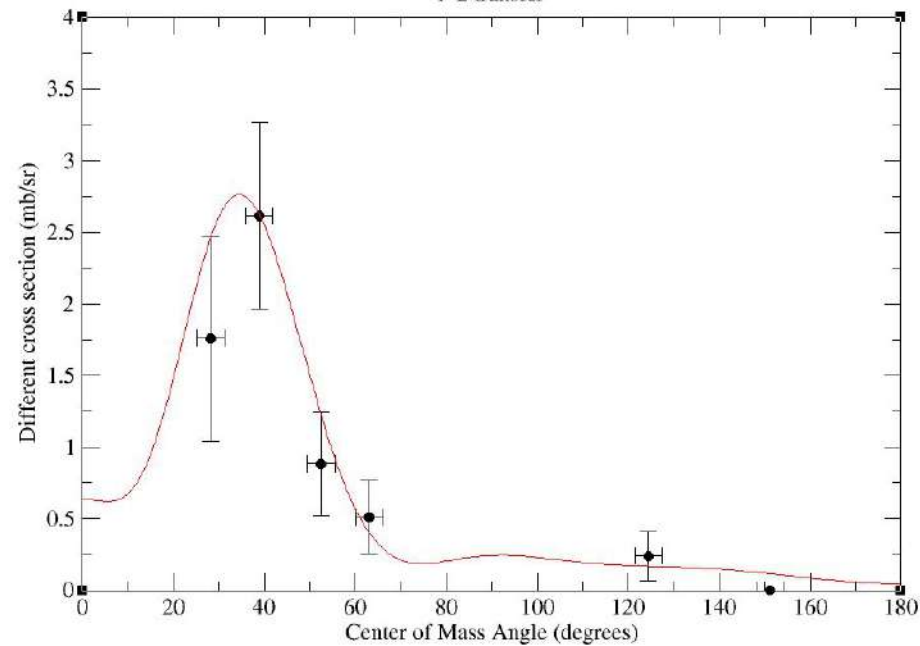




# Angular Distributions

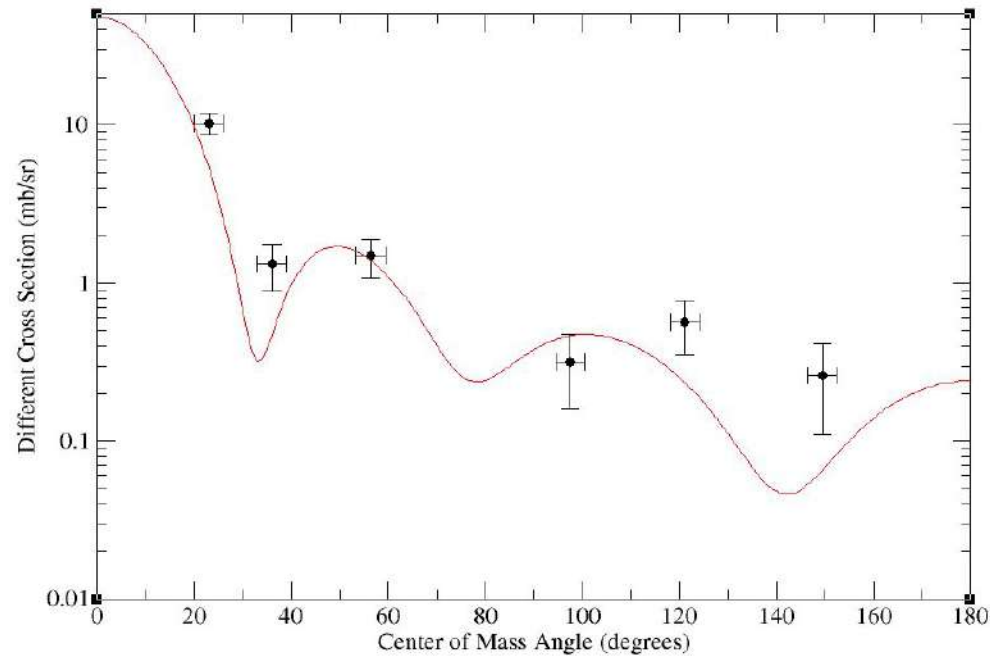
Ground State Angular Distribution (0+)

$l=2$  transfer



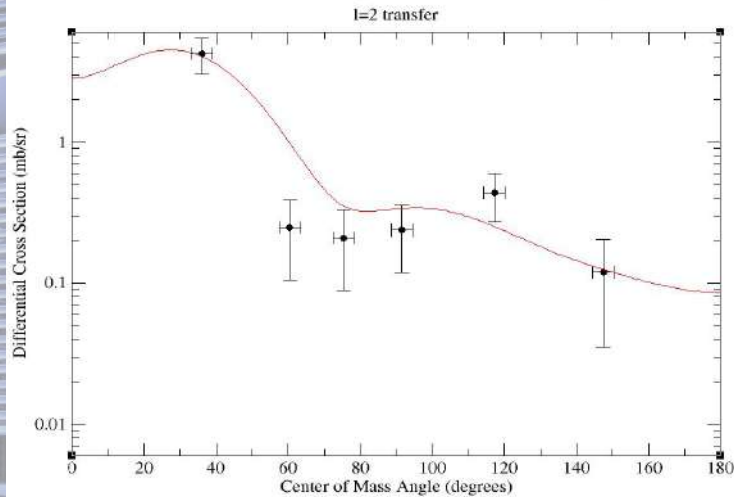
1.887 MeV State Angular Distribution (2+)

$l=0$  transfer

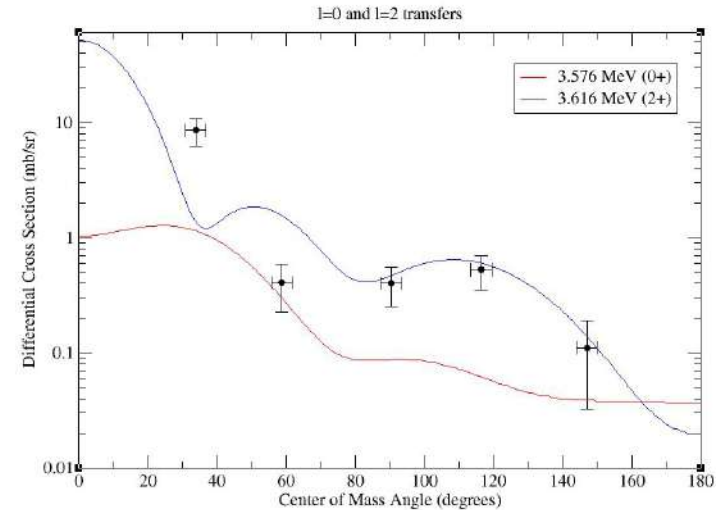


# Angular Distributions

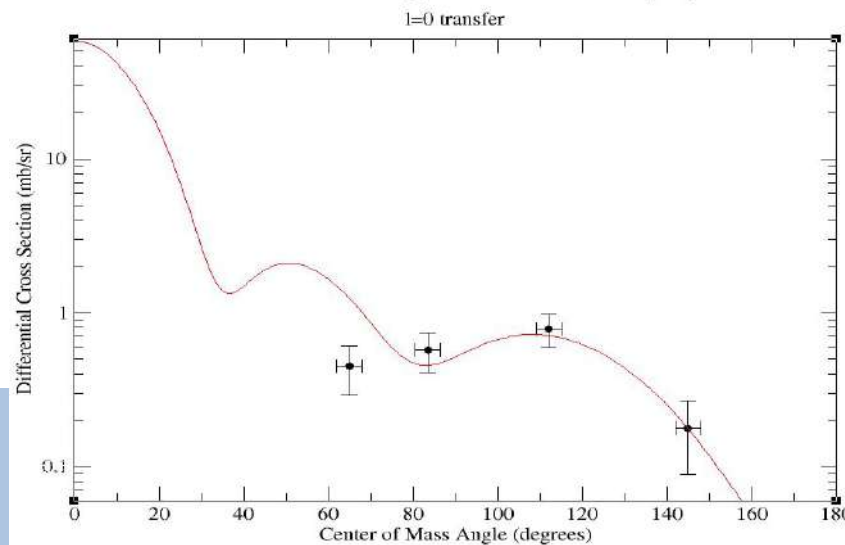
3.376 MeV Angular Distribution (4+)

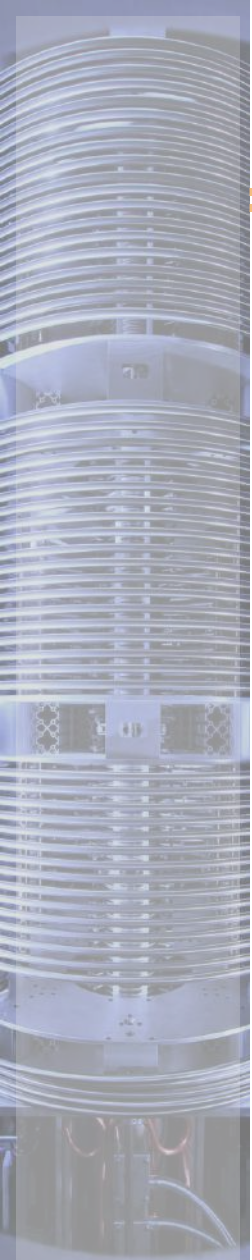


3.6 MeV Doublet Angular Distribution



4.536 MeV Angular Distribution (3+)





# T & R faculty involved in variety of research



The number of nuclear physics teaching and research (T&R) faculty funded by the NSF grant has grown significantly from five in 2010, to seven in 2013, eight in 2016 and finally to nine in the present proposal.



Michael Wiescher,  
Nuclear  
Astrophysics &



Dan Bardayan,  
Nuclear  
Astrophysics



Ani Aprahamian  
Nuclear  
Astrophysics &  
Nuclear Structure  
Physics



Max Brodeur,  
Nuclear  
Astrophysics &  
Fundamental  
Symmetries



Manoel Couder,  
Nuclear  
Astrophysics &  
Nuclear  
Applications



Tan Ahn,  
Nuclear  
Astrophysics &  
Nuclear  
Structure



Philippe Collon,  
Nuclear  
Astrophysics &  
Nuclear  
Applications



Anna Simon,  
Nuclear  
Astrophysics &  
Nuclear  
Reactions



Umesh Garg,  
Nuclear  
Astrophysics &  
Nuclear  
Structure



Graham  
Peaslee,  
Nuclear  
Applications



# NSL research faculty



## JINA



James DeBoer  
R-matrix  
development

## Port



Khachatur Manukyan  
Materials,  
Applications, &  
Targets

## NSF



Daniel Robertson  
CASPAR  
Development



Patrick O'Malley  
TwinSol/St.

## Project based DOE



Georg  
Berg  
SECAR/  
HRS



Jay  
Laverne  
Rad.  
Chemistry

## University support



Joachim  
Görres  
Research  
Support



Ed Stech  
NSL  
Operation



Wanpeng  
Tan  
User  
Support

