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Constrained Fermionic Dynamics of Nuclear Systems: Near Ground State Properties & the Isospin Symmetry

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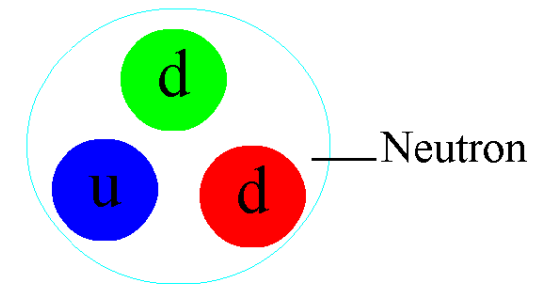
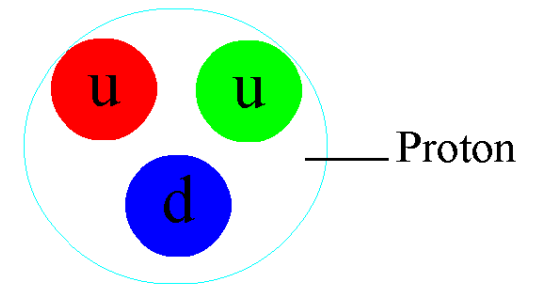
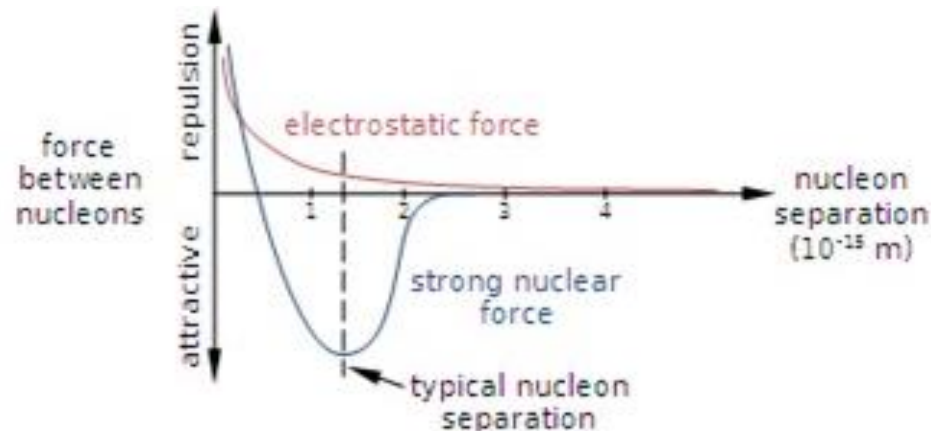
Content Overview

- ❖ **PART I: CoMD & Initial Configurations**
- ❖ **PART II: CoMD & GDR**
- ❖ **PART III: CoMD & Isospin Symmetry**

PART I: “CoMD and Initial Configurations”

The Nuclear N-Body Problem

- ❖ Nuclear Interaction → Fascinating, complicated and still unknown
- ❖ Results from the strong nuclear force between quarks & gluons of the nucleons
- ❖ Can be described by the exchange of pions or phenomenological
- ❖ Includes: central terms, spin-dependence, spin-orbit coupling, tensor terms, momentum and iso-spin dependence

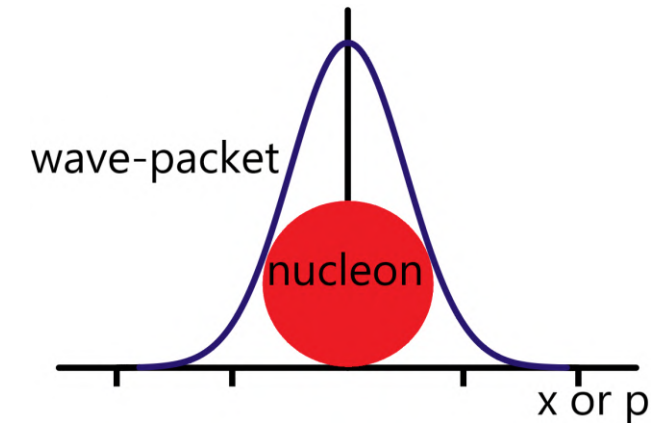


Constrained Molecular Dynamics Approach

- ❖ Hamiltonian: Skyrme Potential → Finite Range of interaction as momentum dependence
- ❖ Volume, surface, Coulomb, 3-Body and symmetry terms

$$V = V^{vol} + V^{surf} + V^{coul} + V^{sym} + V^{(3)}$$

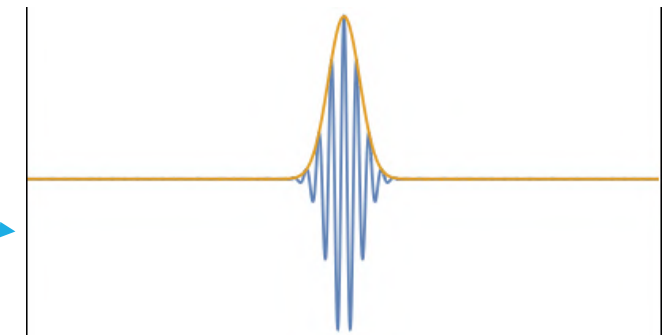
- ❖ Trial wavefunctions → parametrised gaussian wave-packets
→ Time Dependent Variational Principle → Hamiltonian Equations of Motion in Phase Space (Wigner Representation)



$$\frac{d}{dt} \langle \vec{p}_i \rangle = - \frac{\partial H}{\partial \langle \vec{r}_i \rangle}$$

$$\frac{d}{dt} \langle \vec{r}_i \rangle = \frac{\partial H}{\partial \langle \vec{p}_i \rangle}$$

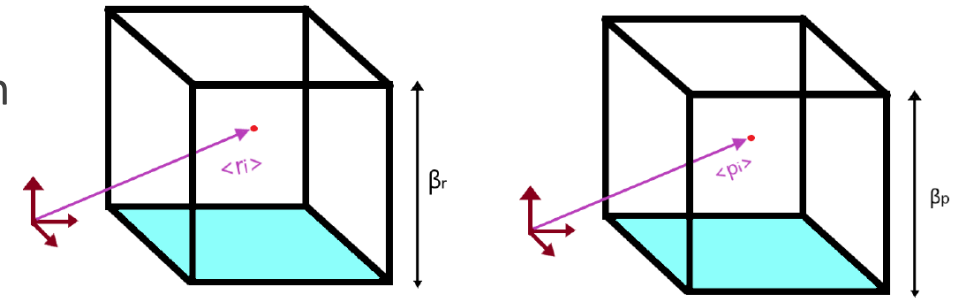
In CoMD σ_r
and σ_p
are
constant!



Constrained Molecular Dynamics Approach

- ❖ Pauli Antisymmetrisation Principle → Constrain of the Phase Space occupation fractions (= probabilities) → Pauli Correlation strength as free parameter $paulm$

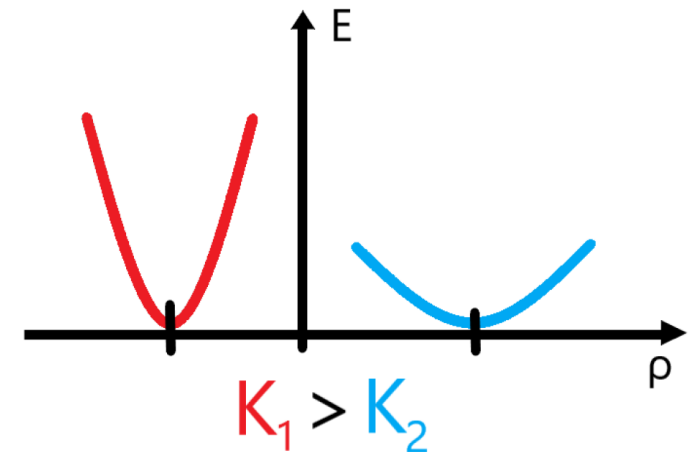
$$\bar{f}_i \leq \frac{paulm}{128}, \forall i$$



- ❖ Compressibility: The nuclear “compression susceptibility” parameter (“hardness of EOS”)

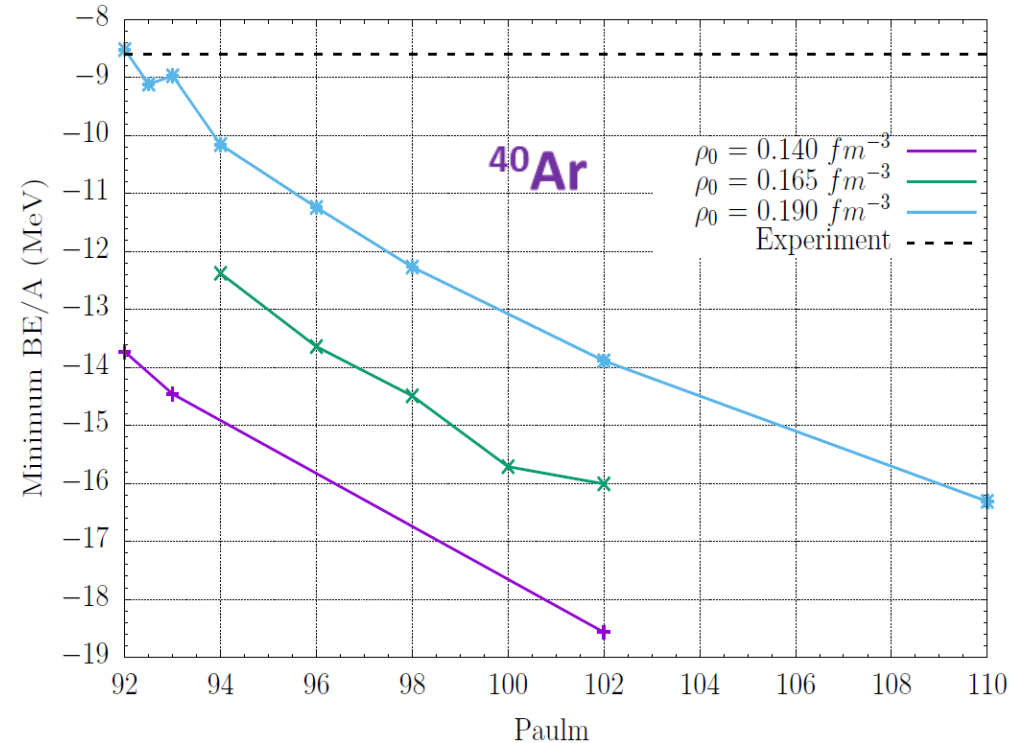
$$K_{NM} \equiv 9\rho_0^2 \left[\frac{\partial^2}{\partial \rho^2} \left(\frac{E}{A} \right) \right]_{\rho_0}$$

- ❖ Equations of Motion → 1st order differential equations → need initial conditions ! Three step calculation → Initialization, Evolution & Data Processing



Initial Configurations and Ground States

- ❖ Statistical Simulated Annealing → Initial Configurations (phase space centroids at t=0)
- ❖ Configuration Space → Depends heavily upon initial parameters
- ❖ Important parameters:
 - ❖ Strength of Pauli Correlations ($paulm$)
 - ❖ Wave-packet width (σ_r)
 - ❖ Saturation Density (ρ_0)
 - ❖ surface parameter (C_{surface})
 - ❖ asymmetry parameter (a_{sym})
 - ❖ Compressibility (K) .
- ❖ Important characteristics: BE/A, neutron skin, rms radii, average density & occupation fraction



$$\frac{E_{Tot}}{A} = -\frac{BE}{A}$$

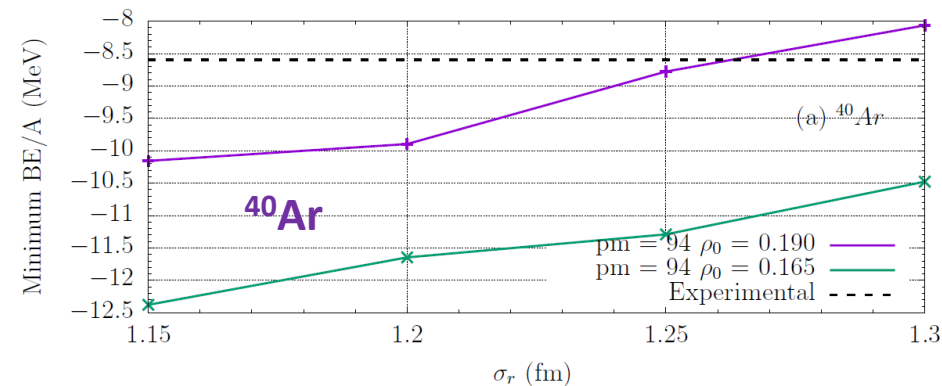
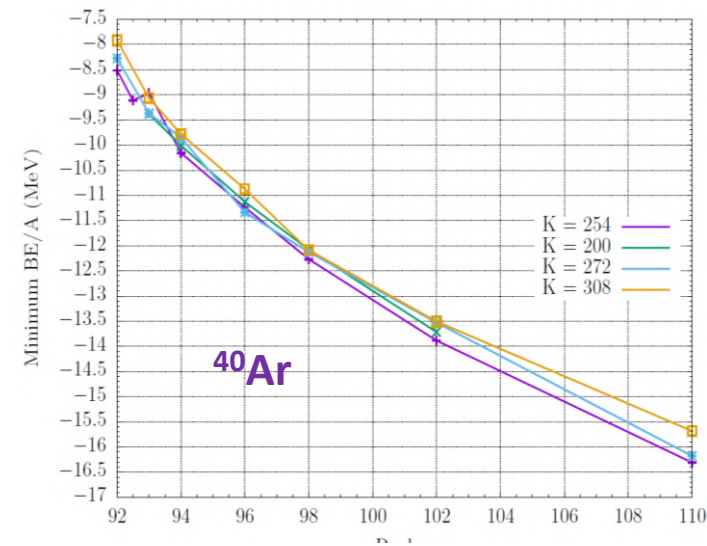
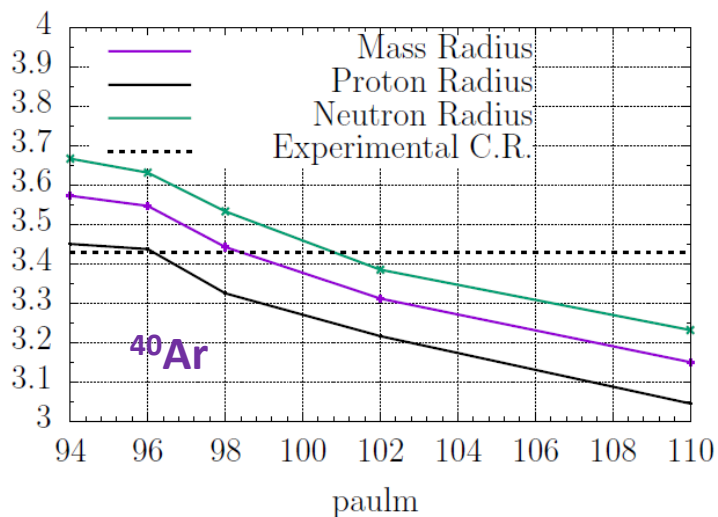
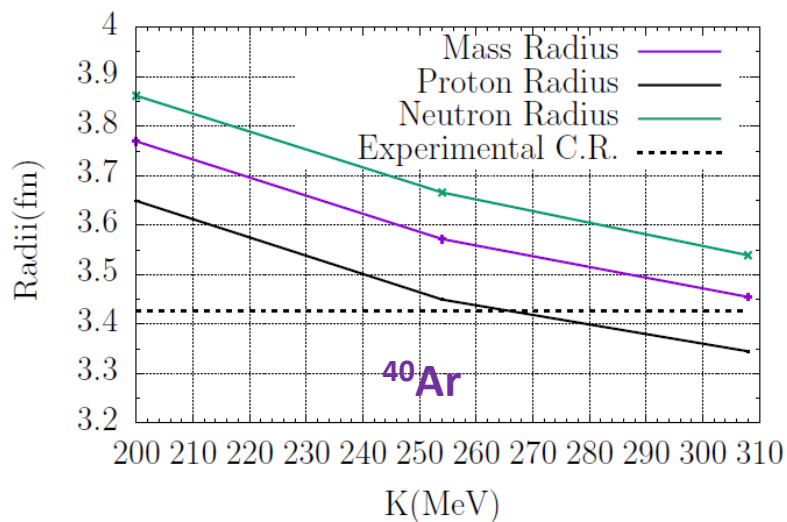
Effect of Important Model Parameters

❖ Increase of $K, \sigma_r, \rho_0, a_{\text{sym}}$

❖ Decrease of $paulm$

} Absolute Decrease of BE/A !

❖ Increase of $K, paulm$ → Absolute Decrease radii & skin !



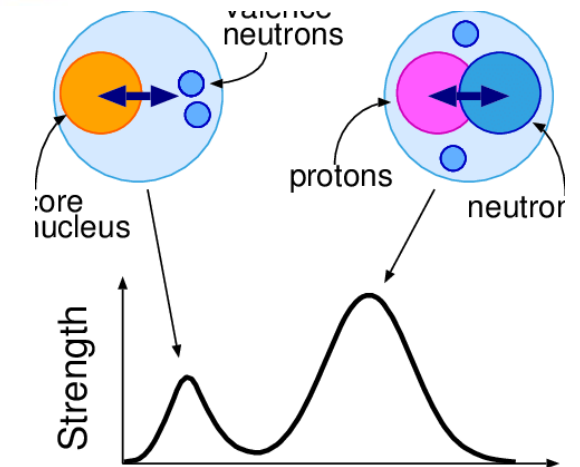
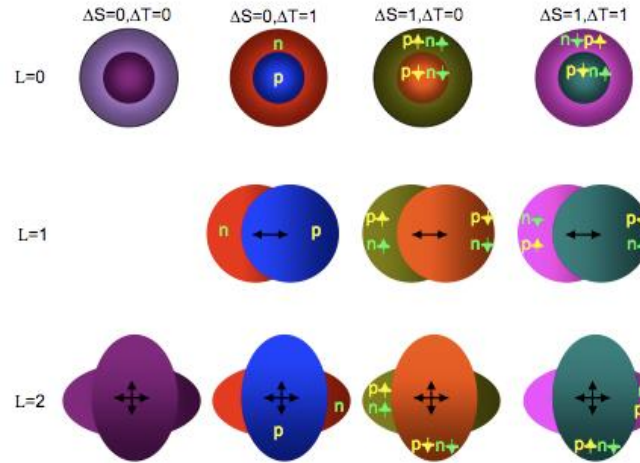
Globally Optimized Configurations

- ❖ Configurations are characterized by:
 - ❖ Binding Energy
 - ❖ Average Occupation Fractions
 - ❖ Radius
 - ❖ Average Density
- ❖ Usually as one characteristic is improved, the others worsen → Effort for Global Optimization !
- ❖ Optimization algorithm → Use of Skyrme-Hartree-Fock calculations, experimental data and empirical parametrizations
- ❖ Results so far: better over-all configuration and characteristics, more stable giant dipole resonances and reactions !

PART II: “CoMD & IVGDR”

Isovector Giant Dipole Resonance (GDR)

- ❖ Low lying excited states → Collective Vibrations! (= modes of rotations and vibrations, like in molecules)
- ❖ GDR → oscillation of proton center of mass against neutron center of mass
- ❖ Results in fission, peripheral, photonuclear reactions, ...
- ❖ Give important information for the nuclear interaction and EoS (= function of energy vs density)



GDR in the CoMD Formalism

- ❖ Perturbation of the initial configurations \rightarrow Damped oscillation (i.e. $D(t)$) \rightarrow Lorentzian Fourier Spectrum
- ❖ Spectrum depends upon the parameters of the effective interaction (K , $paulm$, ...)
- ❖ Width mostly depends on in-medium NN scattering cross-section

$$\sigma_{NN} = a_{redc}(T_{cm}, \bar{\rho}) \sigma_{free}$$

- ❖ Development of a simple model: CoMD based GDR equations of motion

$$\ddot{D} + b\dot{D} + \omega_0^2 D = 0$$

$$b = \frac{\Gamma}{\hbar} \sim \sigma_{NN}$$

$$\omega_0 = \sqrt{\frac{a_{sym}\rho_{np}}{2\rho_0 m \sigma_r^2} \frac{A}{NZ}}$$

$$\omega = \sqrt{\omega_0^2 - \left(\frac{b}{2}\right)^2}$$

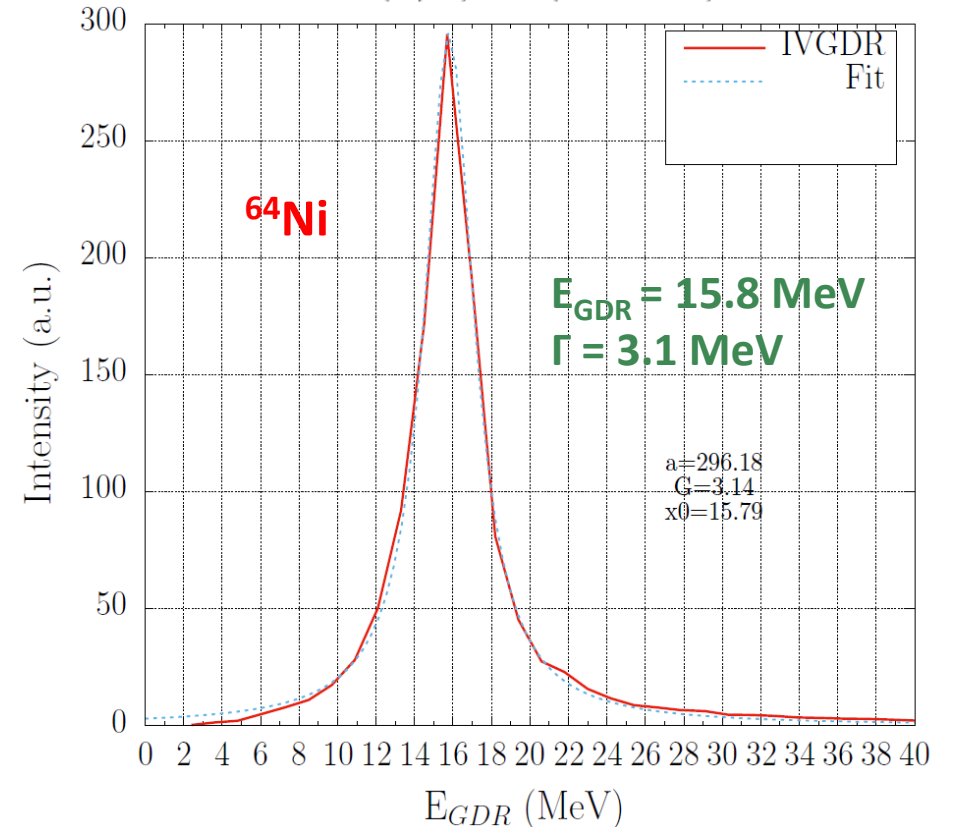
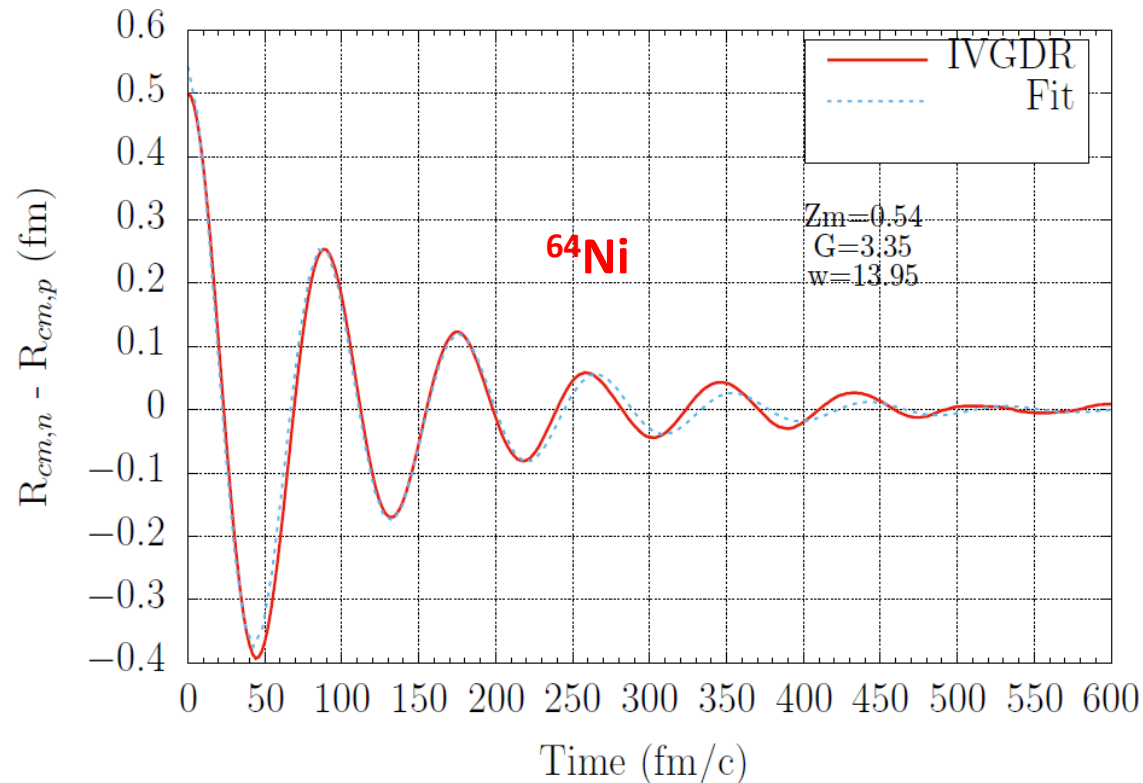
$$\rho_{ij} \sim e^{-\frac{(\langle \vec{r}_i \rangle - \langle \vec{r}_j \rangle)^2}{4\sigma_r^2}} \quad \rho_{np} = \sum_{i \rightarrow p} \sum_{j \rightarrow n} \rho_{ij}$$

GDR in the CoMD Formalism

$$D(t) = D_0 e^{bt/2} \cos \omega t$$

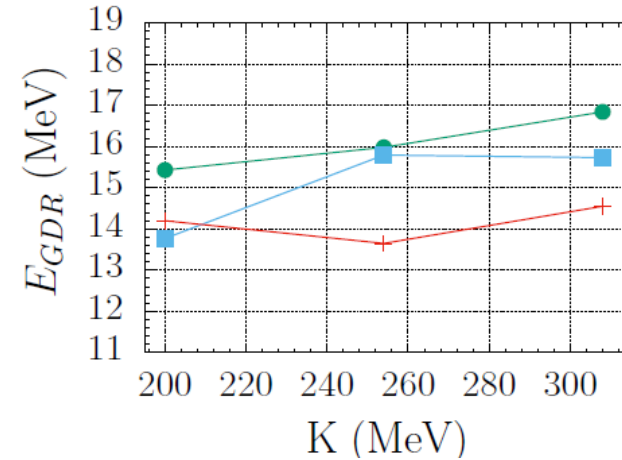
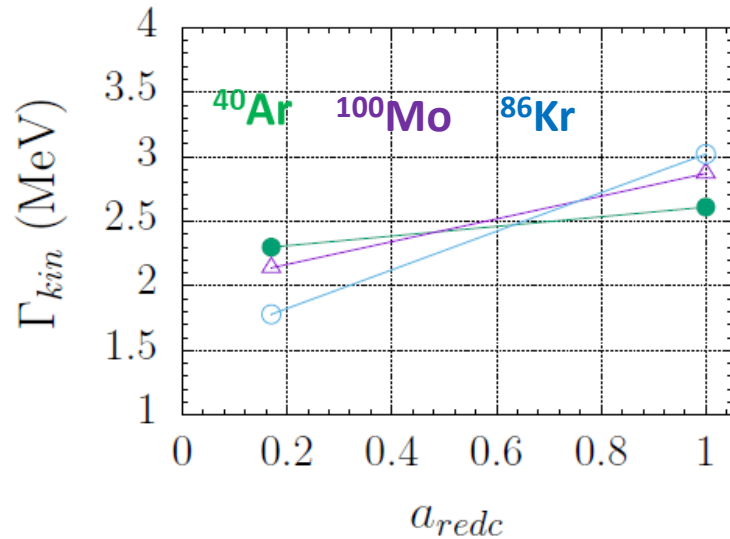
$$b = \frac{\Gamma}{\hbar} \sim \sigma_{NN}$$

$$I = \frac{(\Gamma/2) D_0 \hbar}{(\Gamma/2)^2 + (E - E_D)^2}$$

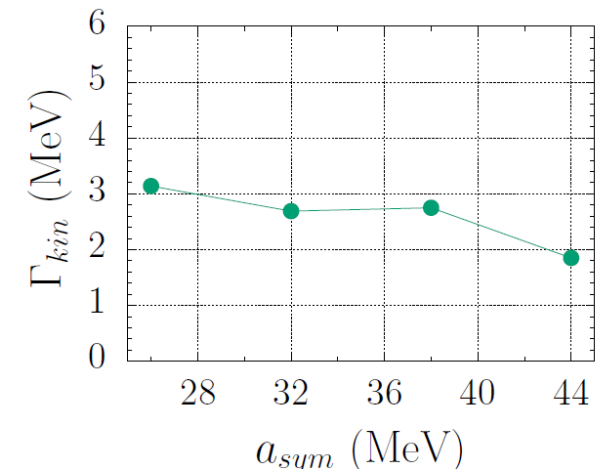
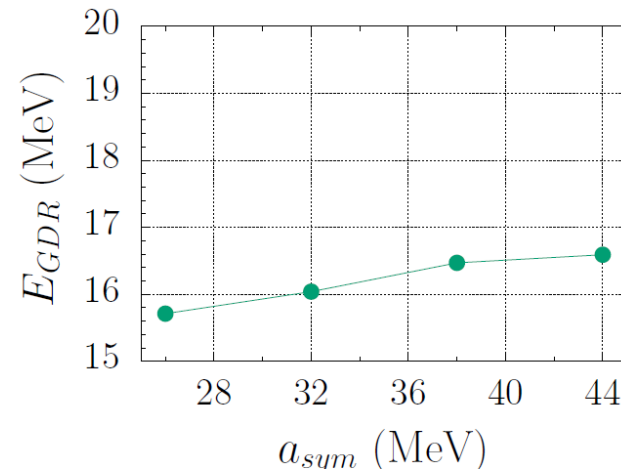


Effect of Model Parameters

- ❖ Increase of K , a_{sym} \longrightarrow Increase of E_{GDR} !
- ❖ Increase of NN cross section \longrightarrow Increase of Γ !



^{40}Ar



Effective Mass and Momentum Dependence

- ❖ Finite Range of interaction → Momentum dependence
- ❖ Skyrme-like Potential of CoMD → Contact force
- ❖ Gaussian ansatz of momentum dependent term:

$$V_{mom}^i = A e^{-c \langle \vec{p}_i \rangle^2}$$

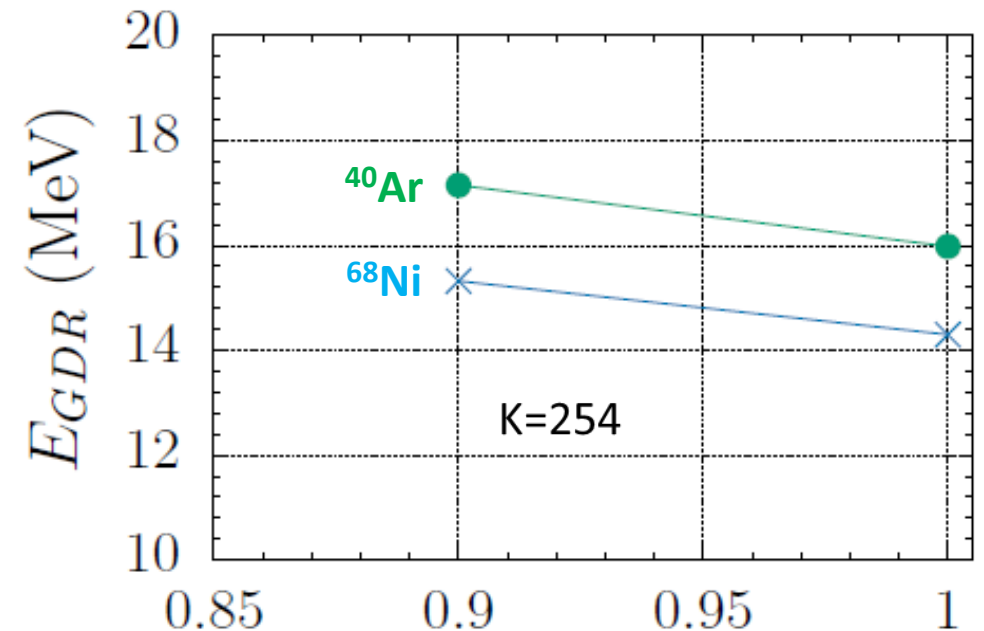
- ❖ Low energy limit → p^2 dependence → “Effective mass”

$$V_{mom}^i \approx A \left(1 - c \langle \vec{p}_i \rangle^2 \right) \longrightarrow T^i = \frac{\langle \vec{p}_i \rangle^2}{2m} (1 - cA) \equiv \frac{\langle \vec{p}_i \rangle^2}{2m^*}$$

- ❖ Parametrization though free parameter:

$$m^* \equiv f_{mass} \cdot m$$

- ❖ Decrease of mass → Increase of E_{GDR}



Mass Dependence of GDR & Skin in CoMD

❖ GDR → Correct dependency upon A

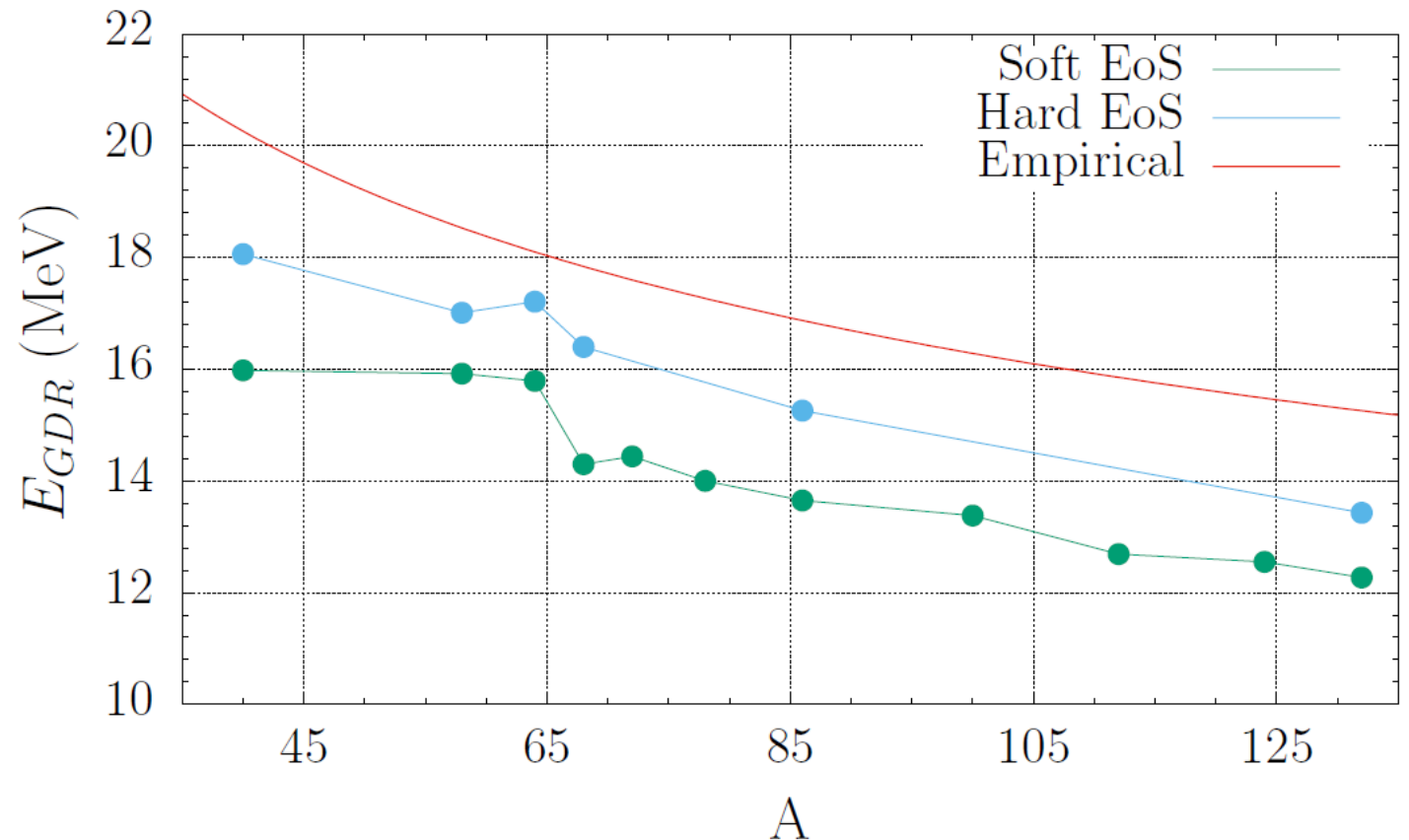
$$E_{GDR} = 31.2A^{-1/3} + 20.6A^{-1/6}$$

❖ Soft EOS → $K=254$, $a_{\text{sym}}=32$ MeV & $m^*/m=1$

❖ Difference from empirical parameterization by about 3-4 MeV

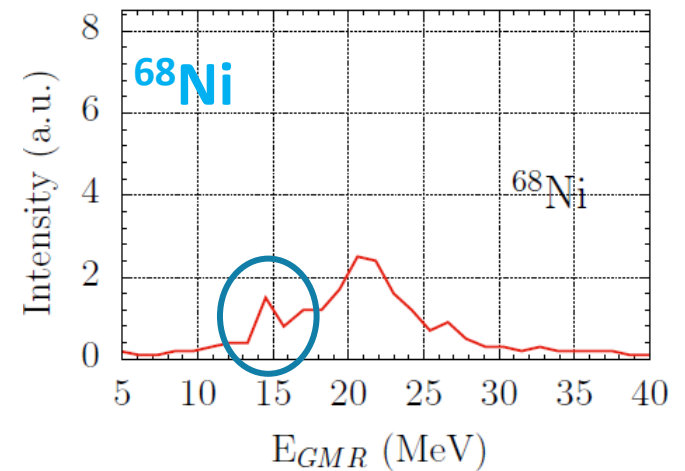
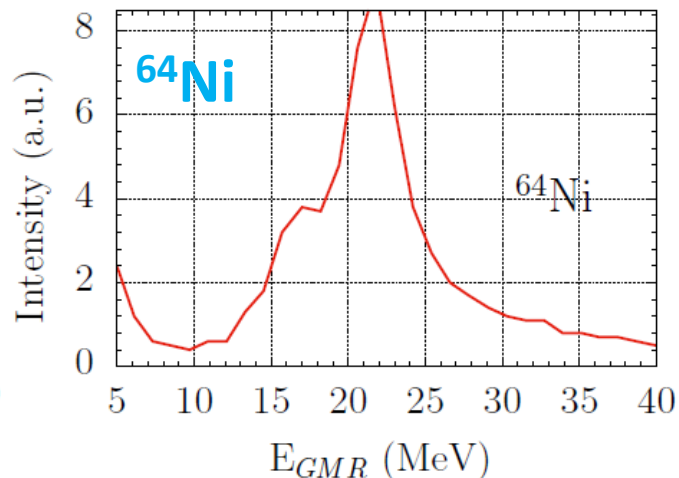
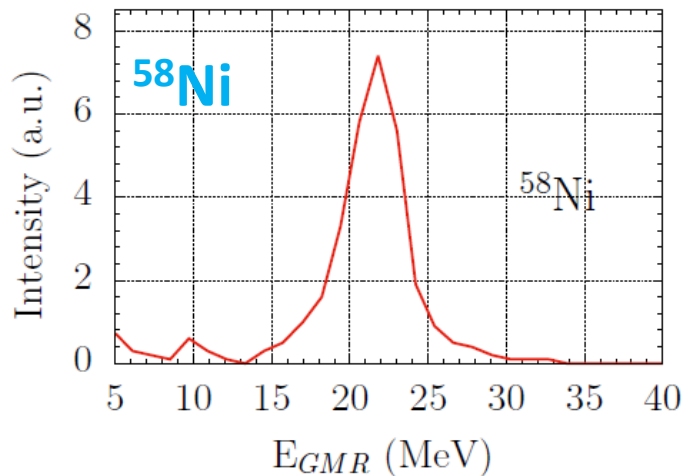
❖ Hard EOS → $K=308$, $a_{\text{sym}}=38$ MeV & $m^*/m < 1$

❖ Difference from empirical parameterization by about 1-2 MeV



Giant Monopole Resonances in CoMD

- ❖ Momentum Space Iso-Scalar Perturbation \rightarrow Temperature
- ❖ Fourier Transform of Radius over time \rightarrow GMR Spectrum !!
- ❖ ${}^x\text{Ni}$ ($x=58,64,68$) $\rightarrow E_{\text{GMR}} \sim 22$ MeV (experimental 21.1 ± 1.9 MeV)
- ❖ Soft monopole for ${}^{68}\text{Ni}$ $\rightarrow E_{\text{soft}} \sim 14.5$ MeV (experimental 12.9 ± 1.0 MeV)



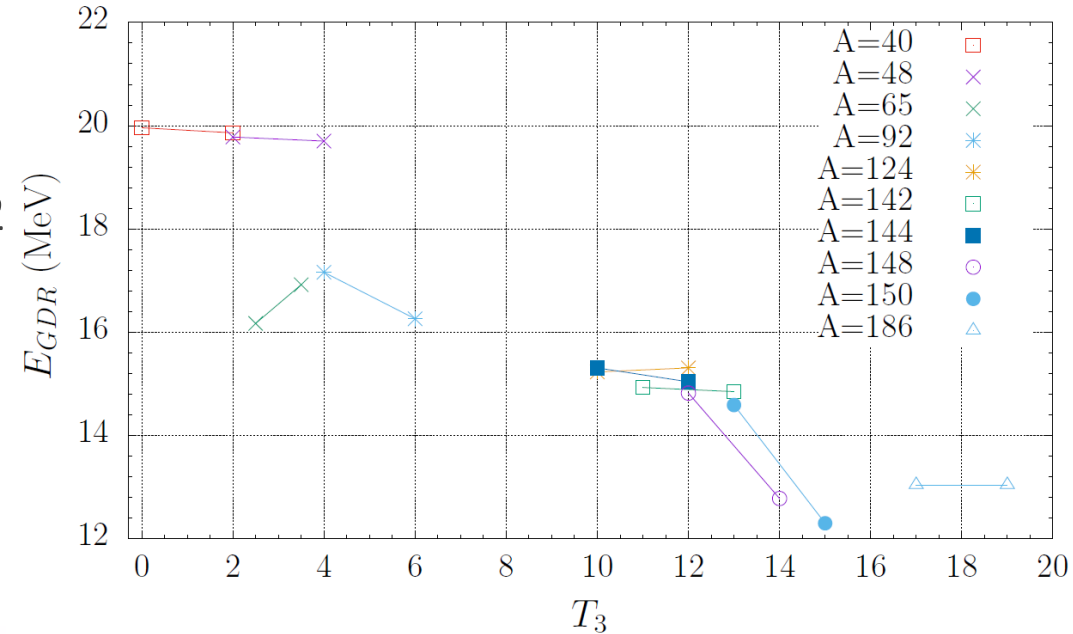
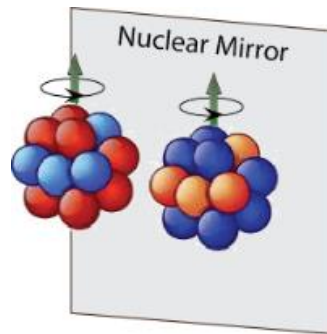
PART III: “CoMD & Isospin Symmetry”

Nuclear Isospin Symmetry

- ❖ Masses of proton-neutron almost equal \rightarrow two states of the same “nucleon particle”
- ❖ Isospin quantum number \rightarrow Identity of the particle $\rightarrow \tau_n = 1/2$ & $\tau_p = -1/2 \rightarrow$ same algebra as $s = 1/2$ particles !
- ❖ Isospin Mirror Transformation \rightarrow exchange of protons and neutrons (i.e. $n \rightarrow p$)
- ❖ Isospin Symmetry \rightarrow Approximate Symmetry of the nuclear systems \rightarrow Breaks due to Coulomb interaction \rightarrow Isospin Symmetry Breaking (ISB), e.g. E_{GDR}

- ❖ Total Coupled Isospin quantum numbers

$$\hat{T}^2|\Psi\rangle = \tau(\tau + 1)|\Psi\rangle \quad \hat{T}_3|\Psi\rangle = \tau_3|\Psi\rangle = \frac{N - Z}{2}|\Psi\rangle$$



Isospin Symmetry of the CoMD

❖ Ground state → Approximate symmetry → Explicit Symmetry Breaking → Coulomb interaction in CoMD Lagrangian

$$V = V^{vol} + V^{surf} + V^{coul} + V^{sym} + V^{(3)}$$

❖ Absolute Neutron Skin → Approximately isospin symmetric

$$|skin| = \left| \sqrt{\langle r_n^2 \rangle} - \sqrt{\langle r_p^2 \rangle} \right|$$

$$L = \frac{e^{bt}}{2} \left(\dot{D}^2 - \omega_0^2 D^2 \right)$$

❖ Extension of GDR theoretical treatment → GDR Lagrangian totally isospin symmetric !

$$\omega_0 = \sqrt{\frac{a_{sym} \rho_{np} A}{2\rho_0 m \sigma_r^2 N Z}}$$

❖ GDR energy depends on A & τ_3^2

$$\frac{A}{NZ} = \frac{A}{4} - \frac{\tau_3^2}{A}$$

$$\rho_{np} = \sum_{i \rightarrow p} \sum_{j \rightarrow n} \rho_{ij}$$

$$\rho_{ij} \sim e^{-\frac{(\vec{r}_i - \vec{r}_j)^2}{4\sigma_r^2}}$$

Isospin Symmetric !!

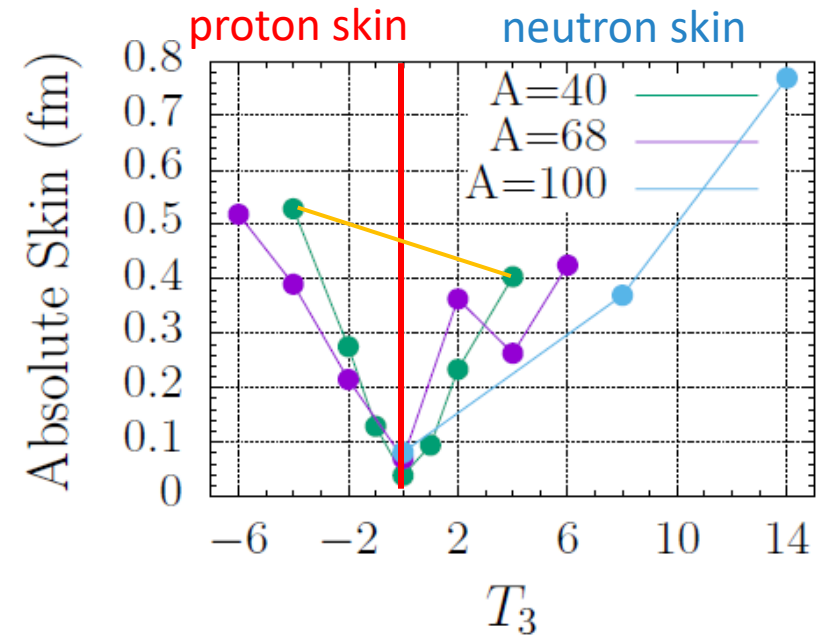
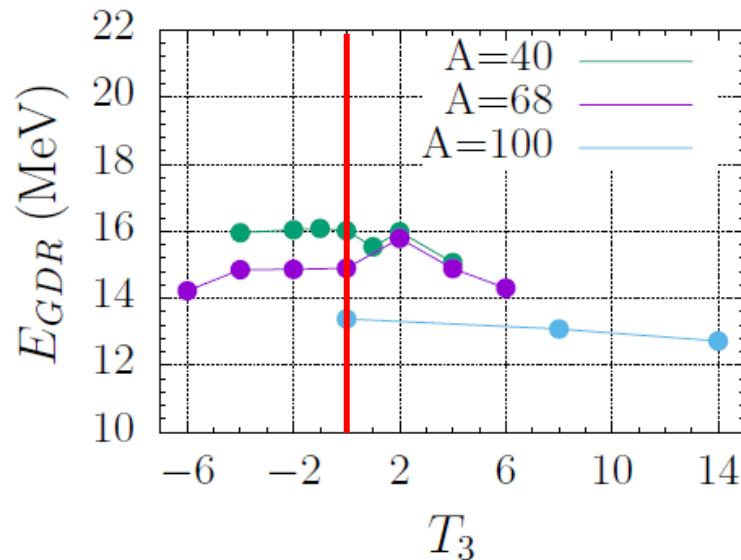
Isospin- T_3 Symmetry Behavior of the CoMD

❖ Absolute Skin \rightarrow Symmetric “well” near $N=Z$ line \rightarrow Explicit Symmetry Breaking

❖ GDR :

❖ Explicit mass dependence around $N=Z$

❖ $T_3 > 0 \rightarrow$ Descending Trend



$$\tau_3 = \frac{N - Z}{2}$$

Summary & Conclusions

- ❖ Extensive use of the CoMD description of the nuclear N-body problem & dynamics
- ❖ Algorithm for selection of optimum configurations → better overall results in calculations
- ❖ Higher K, wave-packet width & Lower *paulm*, saturation density → Higher Total E/A (Less bound nucleus)
- ❖ Higher K, symmetry energy parameter & Lower *paulm*, effective mass → Higher GDR Energy
- ❖ Higher N-N Scattering Cross Section → Greater Width
- ❖ Study of Monopole modes for ^xNi , $x = 58, 64, 68$
- ❖ Examined Isospin symmetry behavior in CoMD:
 - ❖ Protons slightly larger than neutron skins (mirror pairs)
 - ❖ Lower of E_{GDR} with increasing T_3
- ❖ Further Investigations are on going
- ❖ Future Goals: Use of K=308, Explicit Momentum Dependence & Development of E_{GDR} parameterization $f(A,Z)$.



Fin

I WOULD LIKE TO THANK DR. G. SOULIOTIS FOR HIS GUIDANCE AND SUPPORT AND THE OTHER MEMBERS OF THE GROUP AND YOU FOR YOUR ATTENTION !