Covariant density functions in nuclear physics and their microscopic origin

Ioannina, May 5, 2017

Peter Ring

Technical University Munich Excellence Cluster "Origin of the Universe"

Collaborators: Shihang Shen, Jie Meng Peking University



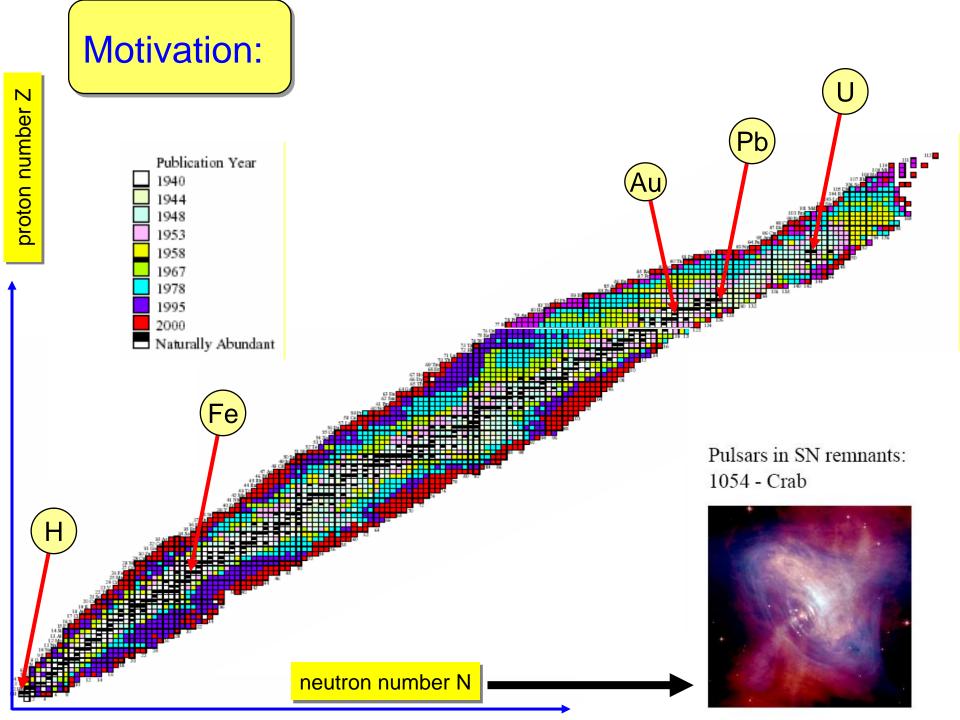




- Motivation
- Phenomenological covariant density functionals
- Semi-microscopic density functionals
- Why do we need a fully microsopic theory ?
- Relativistic Brueckner-Hartree-Fock

In infinite nuclear matter In finite nuclei: local density approximation RBHF-theory in a Dirac-Woods-Saxon Basis Full self-consistent RBHF-theory

- Applications for ⁴He, ¹⁶O, ⁴⁰Ca
- Outlook



Density functional theory for manybody quantum systems

Density functional theory starts from the

Hohenberg-Kohn theorem:

"The exact ground state energy $E[\rho]$ is a universal functional for the local density $\rho(r)$ "

Kohn-Sham theory starts

with a density dependent self-energy:

and the single particle equation:

with the exact density:

$$egin{aligned} h(\mathbf{r}) &= rac{\delta E[
ho]}{\delta
ho(\mathbf{r})} \ h(\mathbf{r}) |arphi_i
angle &= arepsilon_i |arphi_i
angle \
ho(\mathbf{r}) &= \sum_i^A |arphi_i(\mathbf{r})|^2 \end{aligned}$$

In Coulombic systems the functional is derived ab initio

In nuclei DFT has been introduced by **effective Hamiltonians**: by Vautherin and Brink (1972)

$$E = \langle \Psi | H | \Psi \rangle \approx \langle \Phi | \hat{H}_{eff}(\hat{\rho}) | \Phi \rangle = E[\hat{\rho}]$$

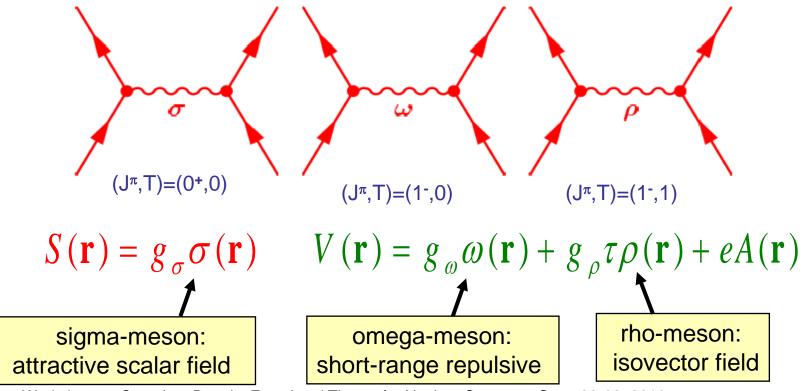
Based on the philosophy of Bethe, Goldstone, and Brueckner one has a in the nuclear interior a density dependen interaction $G(\rho)$

At present the ansatz for $E(\rho)$ is phenomenological:

- Skyrme: non-relativistic, zero range
- Gogny: non-relativistic, finite range (Gaussian)
- CDFT: Covariant density functional theory:
 - Spin-orbit automatically included
 - New saturation mechanism (scalar different from vector density)
 - Proper treatment of time-odd components (nuclear magnetism)
 - Pseudospin symmetry

. . . .

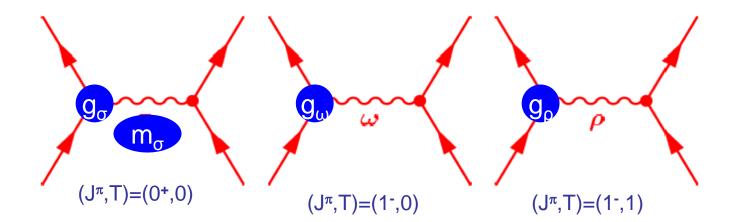
The nuclear fields are obtained by coupling the nucleons through the exchange of effective mesons through an effective Lagrangian.



Workshop on Covariant Density Functional Theory for Nuclear Structure, Sept. 20-22, 2016



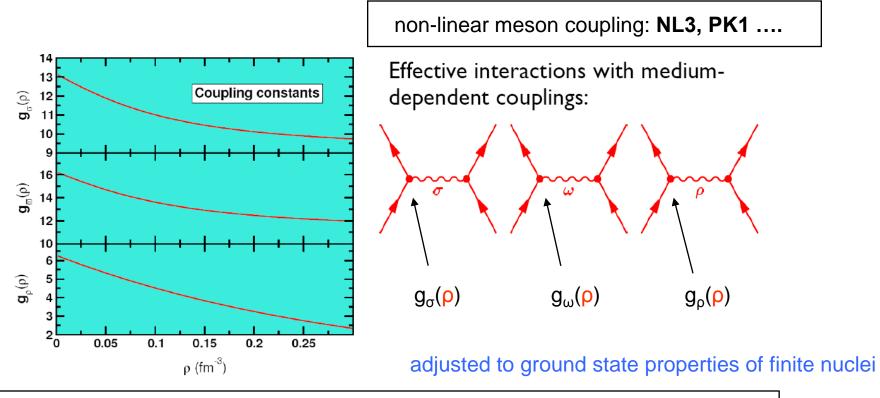
This model has only four parameters:



 $S(r) = g_{\sigma}\sigma(r) \quad V(r) = g_{\omega}\omega(r) + g_{\rho}\rho(r) + eA(r)$

Effective density dependence:

The basic idea comes from ab initio calculations density dependent coupling constants include Brueckner correlations and threebody forces



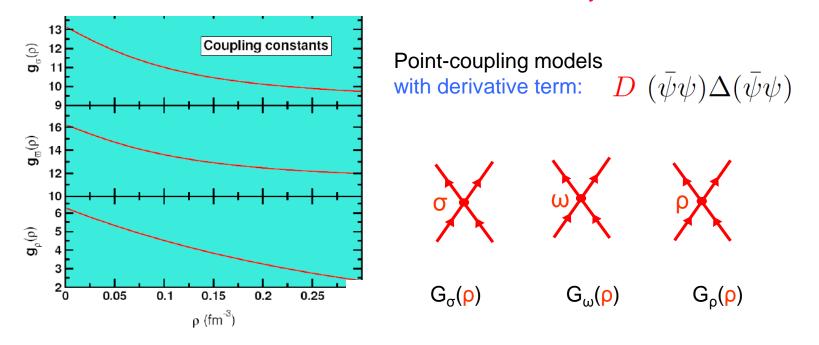
Typel, Wolter, NPA **656**, 331 (1999) Niksic, Vretenar, Finelli, P.R., PRC **66**, 024306 (2002): Lalazissis, Niksic, Vretenar, P.R., PRC 78, 034318 (2008):

DD-ME1 DD-ME2

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Effective density dependence:

The basic idea comes from ab initio calculations density dependent coupling constants include Brueckner correlations and threebody forces



adjusted to ground state properties of finite nuclei

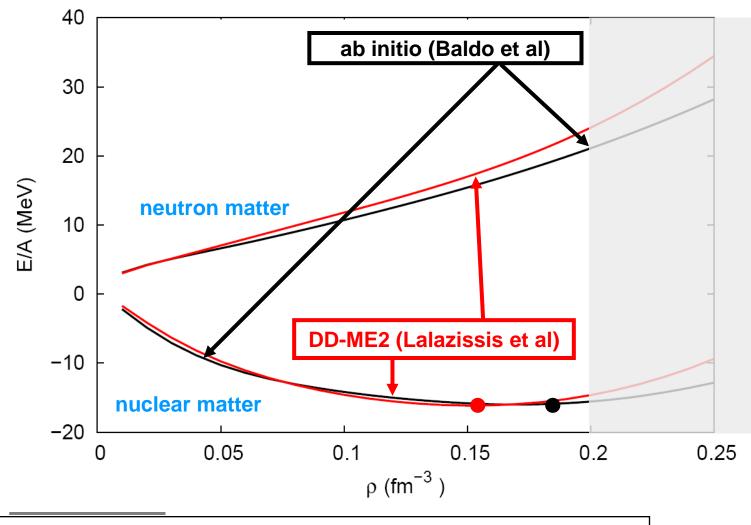
 Manakos and Mannel, Z.Phys. 330, 223 (1988)

 Bürvenich, Madland, Maruhn, Reinhard, PRC 65, 044308 (2002):

 Niksic, Vretenar, P.R., PRC 78, 034318 (2008):

 Zhao, Li, Yao, Meng, J. Meng, PRC 82, 054319 (2010)

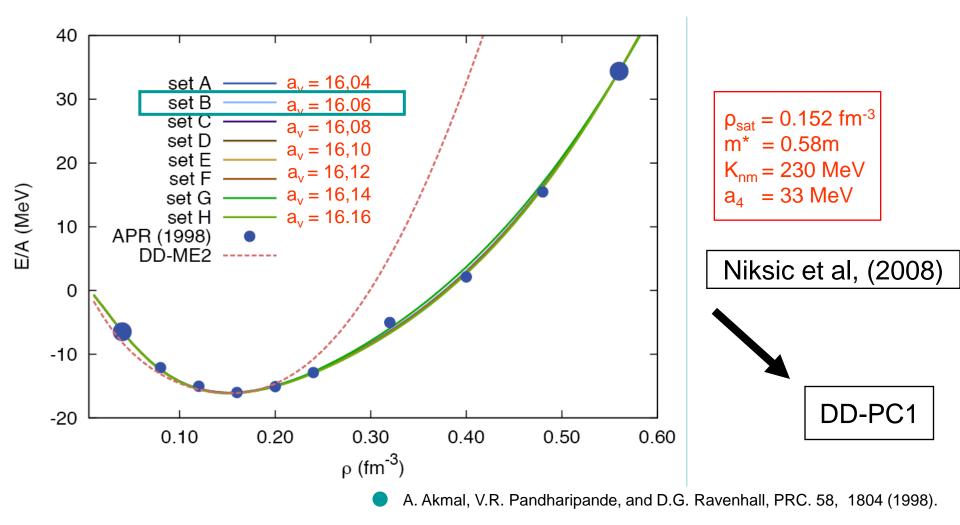
Comparison with ab initio calculations:

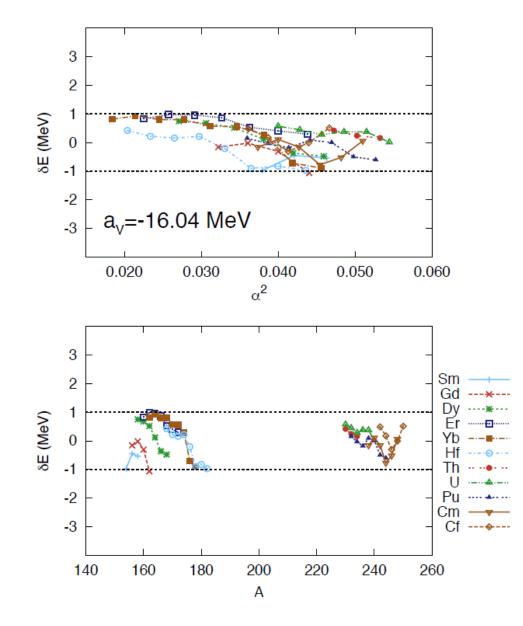


we find excellent agreement with ab initio calculations of Baldo et al.

Semi-microscopic relativistic functionals

point coupling model is fitted to microscopic nuclear matter and to masses of 64 deformed nuclei:





Rare-earth region Sm (Z=62), Gd (Z=64), Dy (Z=66), Er (Z=68), Yb (Z=70), Hf (Z=72)

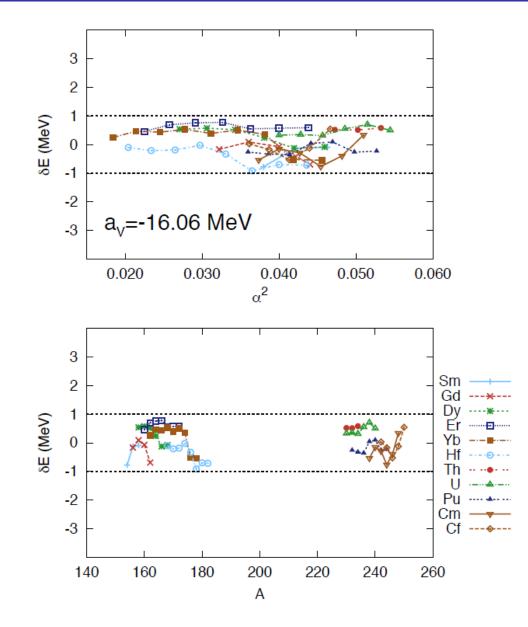
Actinides Th (Z=90), U (Z=92), Pu (Z=94), Cm (Z=96), Cf (Z=98)

Total 64 isotopes

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T. Niksic

4th Workshop on New Aspects and Perspectives in Nuclear Physics, Ioannina University, May 5-6, 2017 12/44



Rare-earth region Sm (Z=62), Gd (Z=64), Dy (Z=66), Er (Z=68), Yb (Z=70), Hf (Z=72)

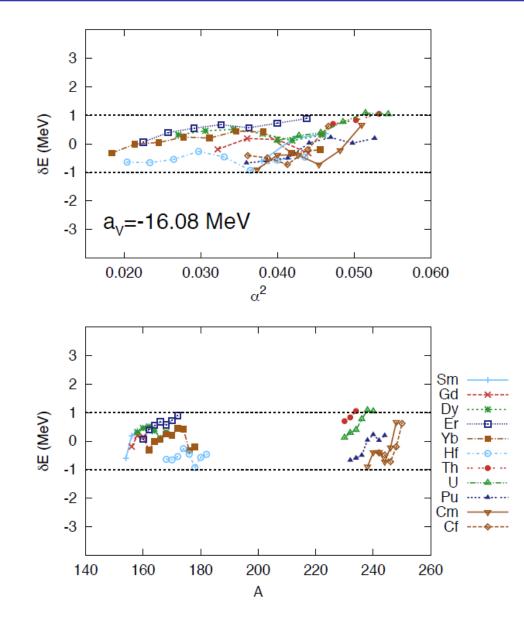
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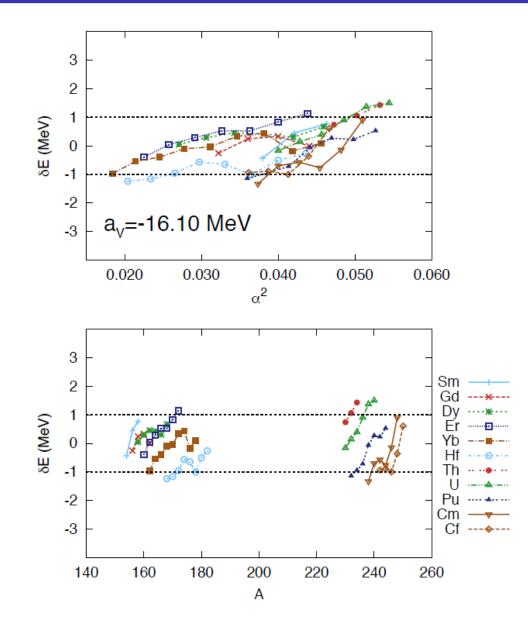
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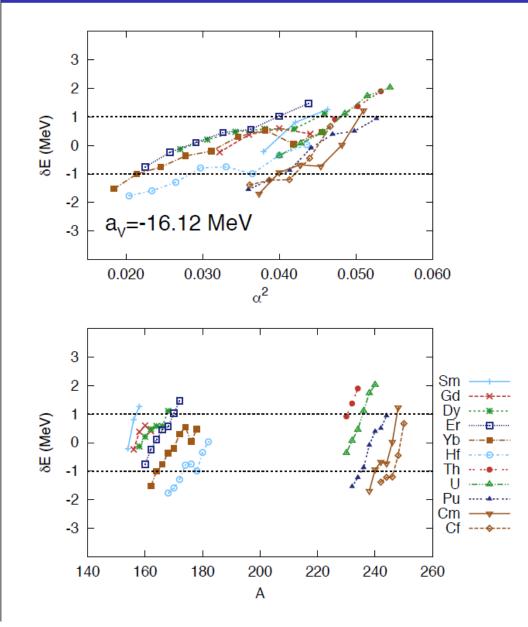
Total 64 isotopes

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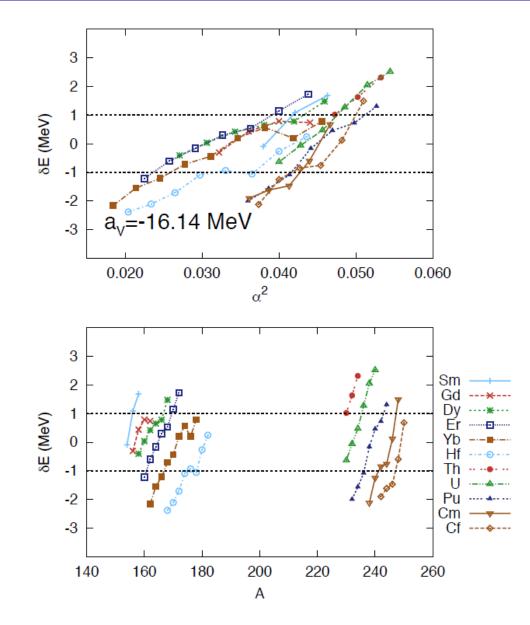
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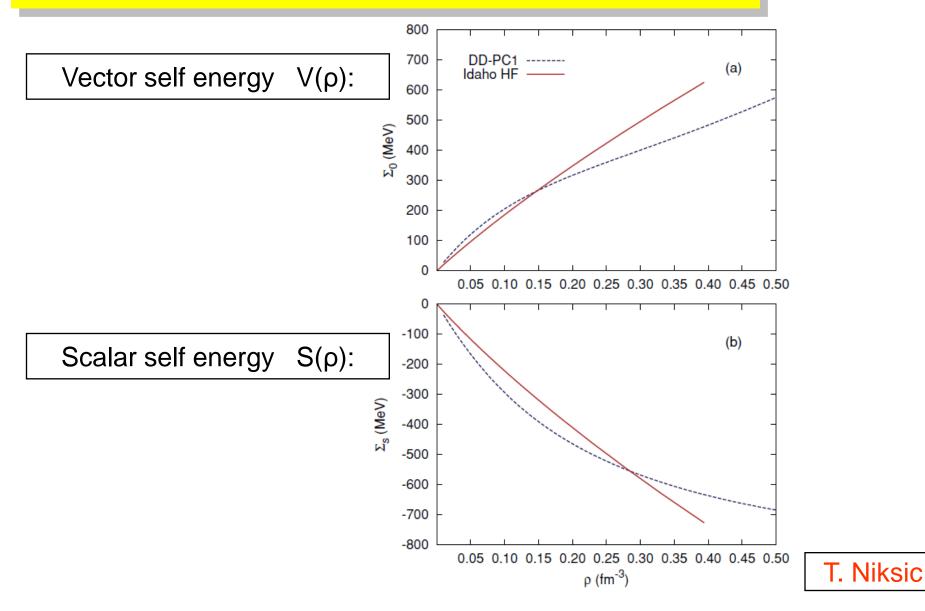
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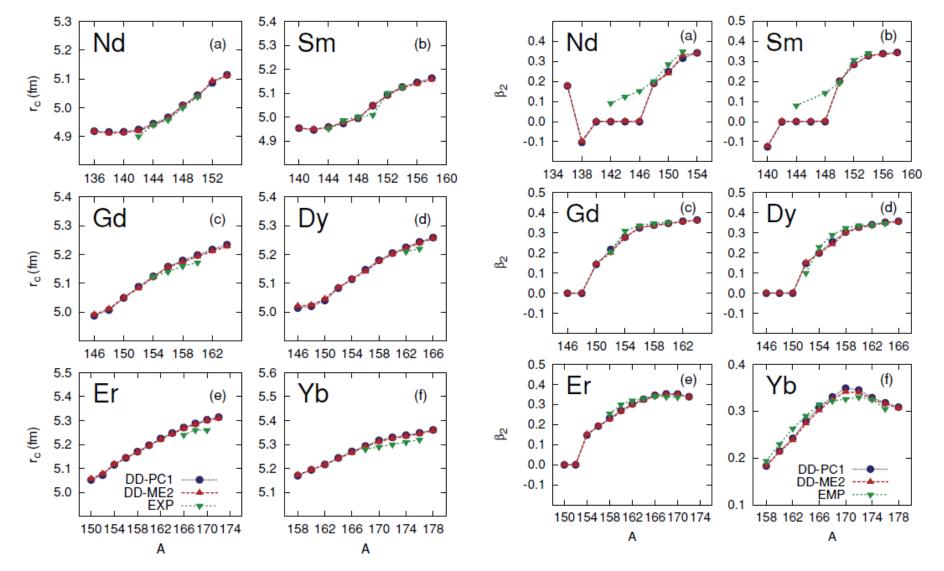
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DD-PC1 and microscopic self energies:



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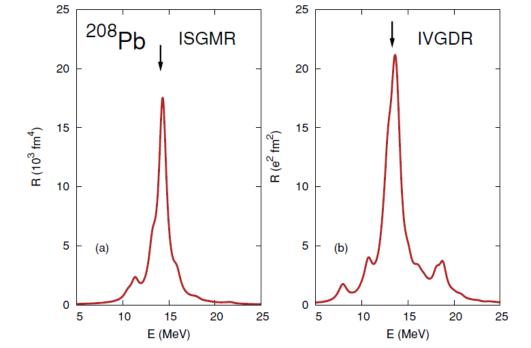
DD-PC1: isotope shifts and deformation parameters

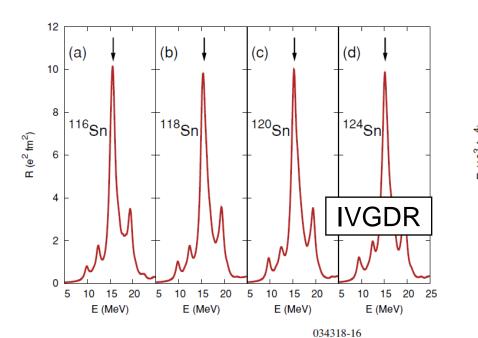


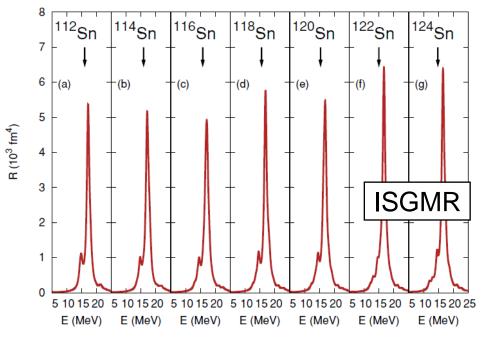
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DD-PC1 Giant resonances:

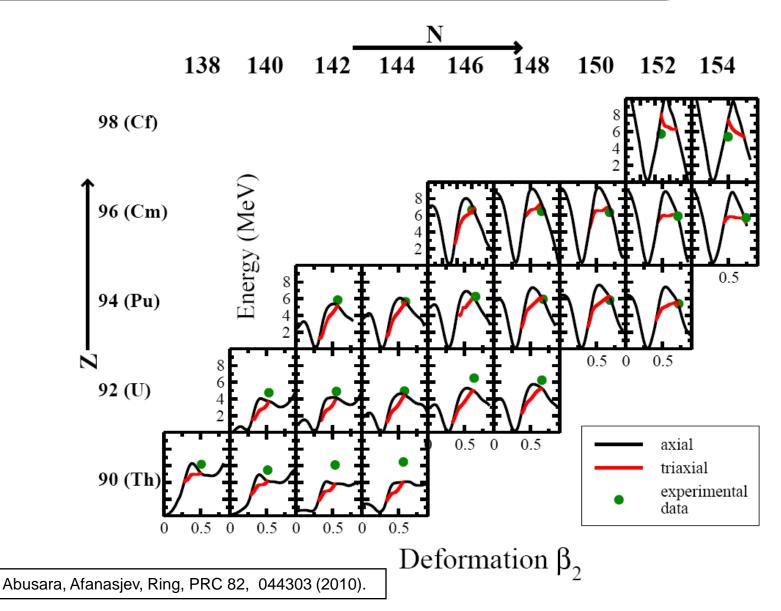
T. Niksic et al, (2008)





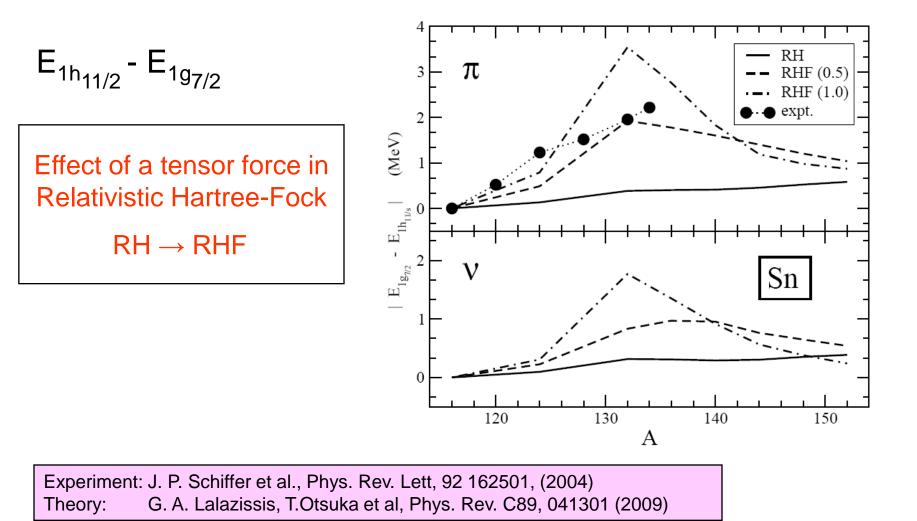


Fission barriers for triaxially deformed shapes:



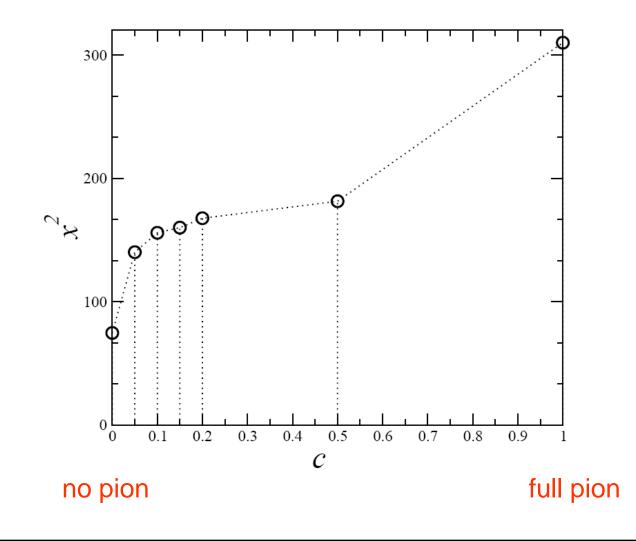
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Single-particle energy splittings in Sn-isotopes:



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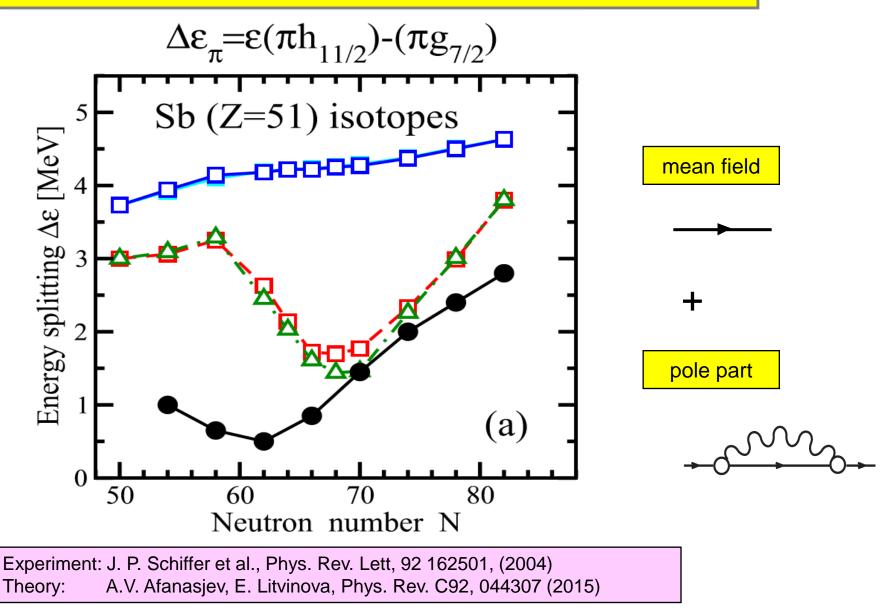
Fit to masses and radii (as DD-ME2) + $c L_{\pi N}$



Lalazissis, Karatzikos, Serra, Otsuka, P.R., PRC 80, 041301(R) (2009)

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Influence of Particle-Vibrational Coupling



Further observations :

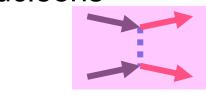
- effective single particle energies in shell-model calculations show specific trends due to the tensor force, which agrees with many data (Otsuka, Schiffer, Greavy)
- the same trend can be found qualitatively if one adds a the pion with an effective coupling constant in fully selfconsistent RHF calculations (Lalazissis, Long)
- particle-vibrational coupling is also important

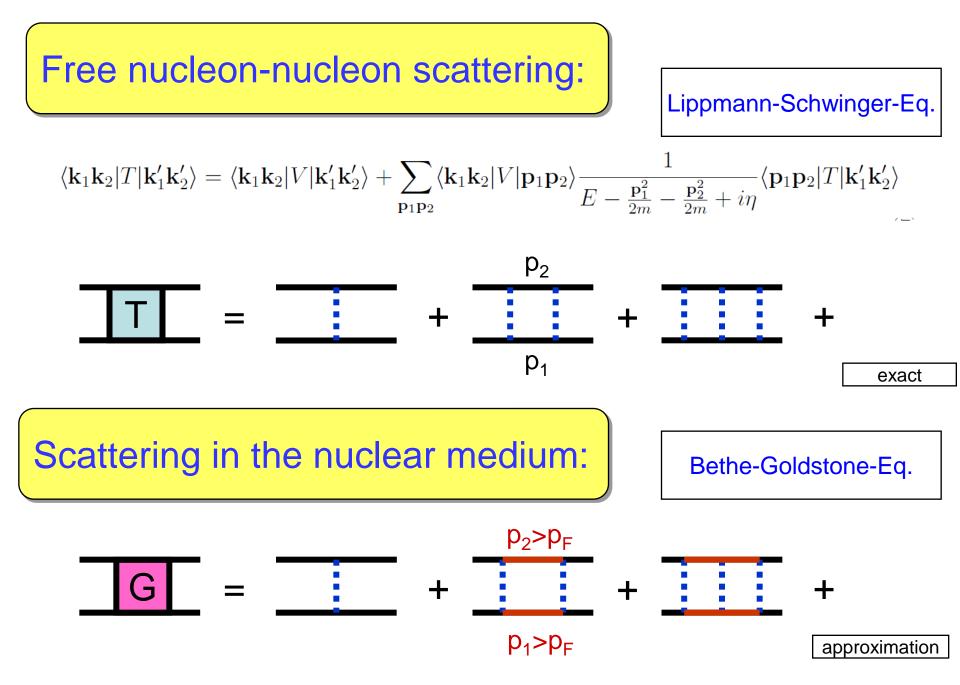
How important are effective tensor forces?

Brueckner theory (1958):

Brueckner, Gammel, Phys. Rev. 109, 1023 (1958)

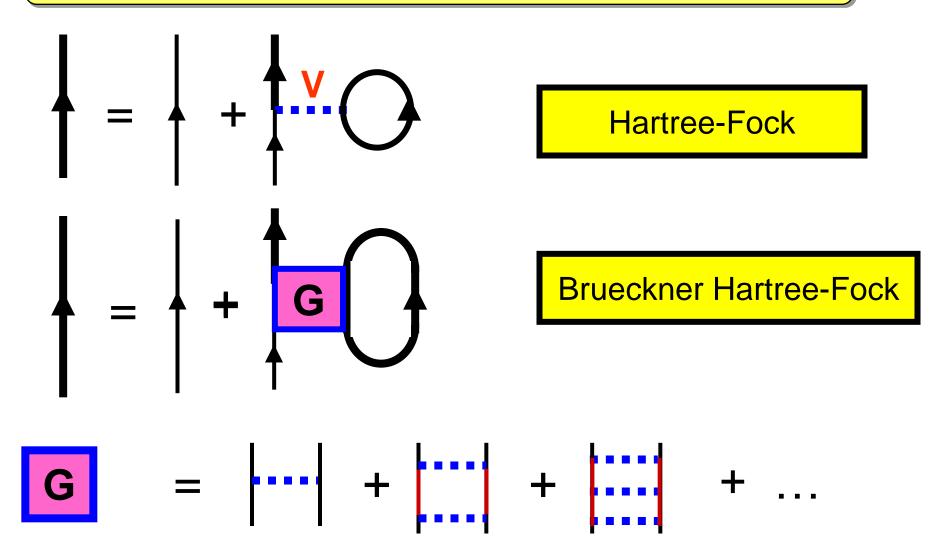
- The nucleons in the interior of the nuclear medium do not feel the same bare force V, as the nucleons feel in free space.
- They feel an effective force G.
- The Pauli principle prohibits the scattering into states, which are already occupied in the medium.
- Therefore this force $G(\rho)$ depends on the density
- This force G is much weaker than bare force V.
- Nucleons move nearly free in the nuclear medium and feel only a strong attraction at the surface (shell model)



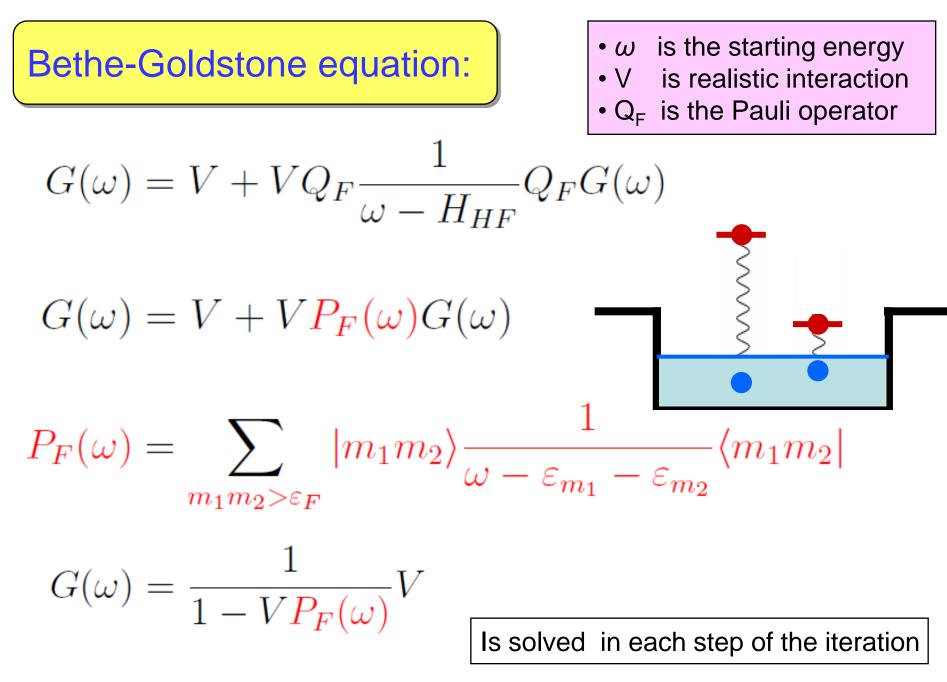


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Ab-initio: Relativistic Brueckner Hartree-Fock:



Summing up all ladder diagramms



Solution of the Bethe-Goldstone equation:

Bethe-Goldstone equation in basis space

$$\langle ab|G(\omega)|a'b'\rangle = \langle ab|\bar{V}^N|a'b'\rangle + \sum_{\varepsilon_m\varepsilon_n > \varepsilon_F} \frac{\langle ab|\bar{V}^N|mn\rangle\langle mn|G(\omega)|a'b'\rangle}{\omega - \varepsilon_m - \varepsilon_n},$$

where \mathcal{E}_F is the Fermi energy, $\mathcal{O} = \mathcal{E}_a + \mathcal{E}_b$ is the starting energy and $|mn\rangle$ are intermediate states.

Bethe-Goldstone equation in plane wave basis

$$G_{ll'}^{\alpha}(kk'K\omega) = V_{ll'}^{\alpha}(kk') + \sum_{ll'} \int \frac{d^3q}{(2\pi)^3} V_{ll'}^{\alpha}(kq) \frac{Q(q,K)}{\omega - H_0} G_{ll'}^{\alpha}(qk'K\omega)$$

where \mathcal{C} is a shorthand notation for J, S, L and T.

Matrix inversion method

$$G = \left(1 - \frac{V}{\omega - H_0}\right)^{-1} V$$

RBHF theory in finite nuclei:

$$(\boldsymbol{\alpha}\boldsymbol{p} + \beta M + U)|a\rangle = \varepsilon_a|a\rangle$$

Normal Hartree-Fock:

Brueckner-Hartree-Fock:

$$U_{ab}^{\rm HF} = \sum_{c=1}^{A} \langle ac | \bar{V} | bc \rangle$$
$$U_{ab}^{\rm BHF} = \sum_{c=1}^{A} \langle ac | \bar{G}(\omega) | bc \rangle$$

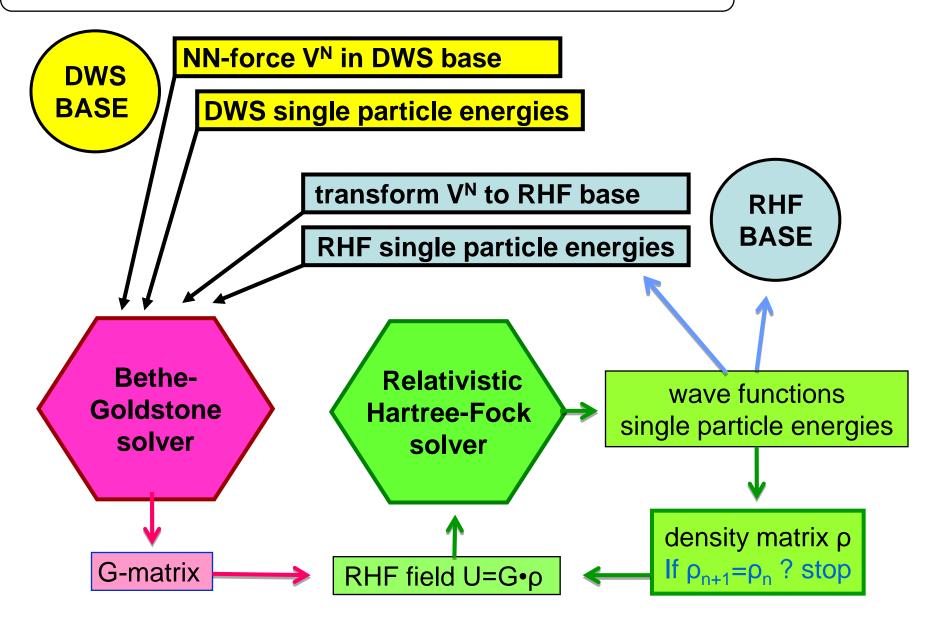
RBHF theory in finite nuclei:

$$(\boldsymbol{\alpha}\boldsymbol{p} + \beta M + U)|a\rangle = \varepsilon_a|a\rangle$$

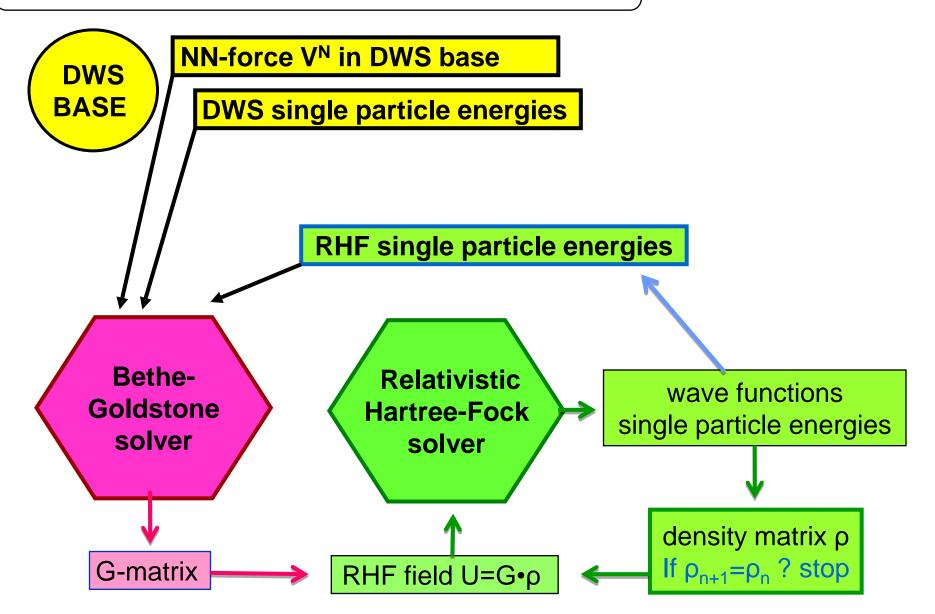
Bethe-Brandow-Petschek Theorem:

$$\begin{split} U_{ab}^{\rm BHF} = & \frac{1}{2} \sum_{c=1}^{A} \langle ac | \bar{G}(\varepsilon_{a} + \varepsilon_{c}) + \bar{G}(\varepsilon_{b} + \varepsilon_{c}) | bc \rangle \\ & \text{for } \varepsilon_{a} < \varepsilon_{F}, \ \varepsilon_{b} < \varepsilon_{F} \\ U_{ab}^{\rm BHF} = & \sum_{c=1}^{A} \langle ac | \bar{G}(\varepsilon_{a} + \varepsilon_{c}) | bc \rangle \\ & \text{for } \varepsilon_{a} < \varepsilon_{F}, \ \varepsilon_{b} > \varepsilon_{F} \\ U_{ab}^{\rm BHF} = & \frac{1}{2} \sum_{c=1}^{A} \langle ac | \bar{G}(\varepsilon_{a}' + \varepsilon_{c}) + \bar{G}(\varepsilon_{b}' + \varepsilon_{c}) | bc \rangle \\ & \text{for } \varepsilon_{a} > \varepsilon_{F}, \ \varepsilon_{b} > \varepsilon_{F} \end{split}$$

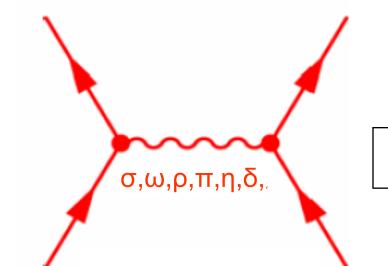
Flow chart for RBHF theory in finite nuclei:



Simpification with fixed DWS-basis:



Bare nucleon-nucleon force:

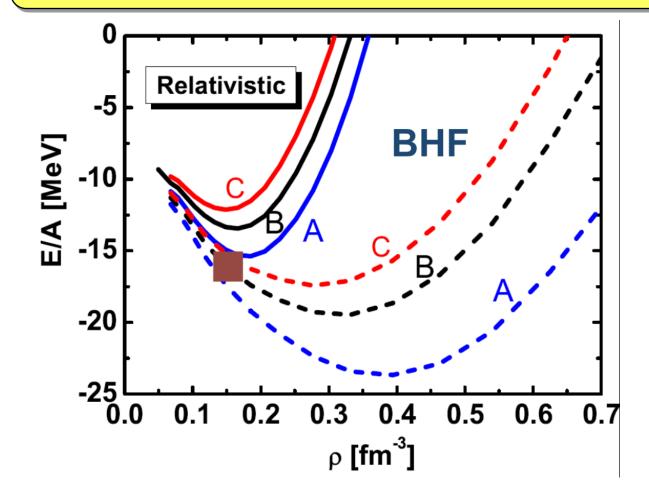


Brockmann and Machleidt, PRC 42, 1965 (1990).

Meson Parameters		Potential A		Potential B		Potential C	
Fai	m_{α} (MeV)	$g_{\alpha}^{2}/4\pi$	Λ_{α} (GeV)	$g_{\alpha}^2/4\pi$	Λ_{α} (GeV)	$g_{\alpha}^{2}/4\pi$	Λ_{α} (GeV)
π	138.03	14.9	1.05	14.6	1.2	14.6	1.3
η	548.8	7	1.5	5	1.5	3	1.5
ρ	769	0.99	1.3	0.95	1.3	0.95	1.3
ω	782.6	20	1.5	20	1.5	20	1.5
δ	983	0.7709	2.0	3.1155	1.5	5.0742	1.5
σ	550	8.3141	2.0	8.0769	2.0	8.0279	1.8

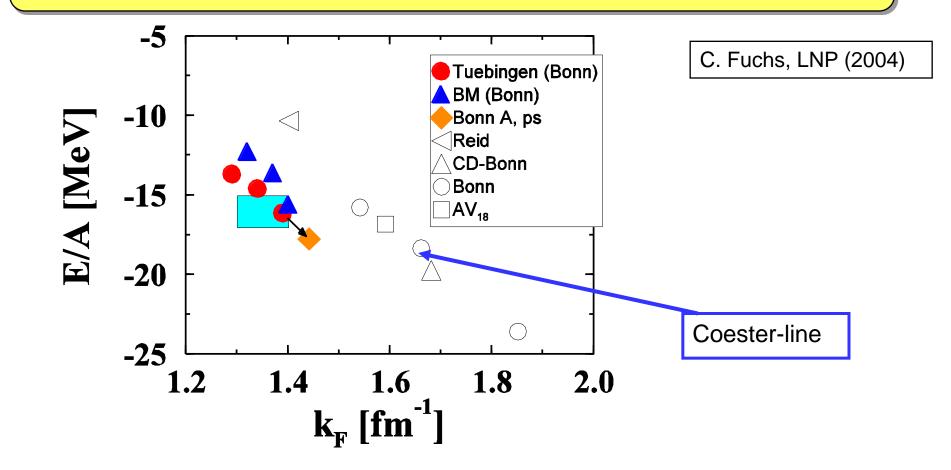
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Dirac-Brueckner-Hartree-Fock in nuclear matter



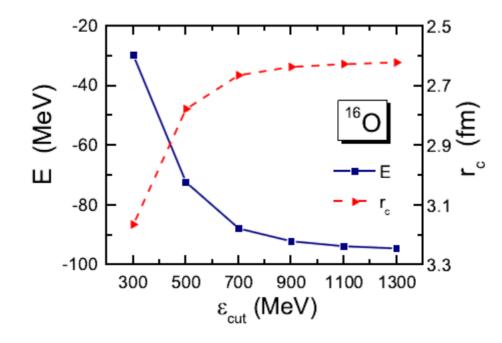
Brockmann and Machleidt, PRC 42, 1965 (1990).

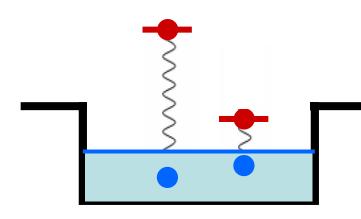
Dirac-Brueckner-Hartree-Fock in nuclear matter





Convergenge with the cuf-off in single particle energy:





Local density approximation:

Bonn A

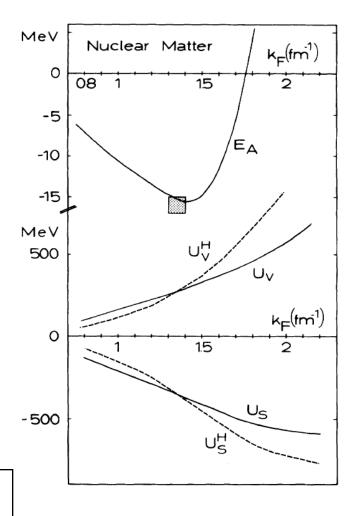
E/A Binding energy per particle in Rel. Brückner-Hartree-Fock

 U_S, U_V : scalar and vector potential $\Sigma(\rho)$ in Rel. Brueckner-Hartree-Fock for nuclear matter with density ρ

 \rightarrow $g_{\sigma}(\rho)$ and $g_{\omega}(\rho)$

U^H_S,U^H_V: scalar and vector potential in Relativistic. Hartree

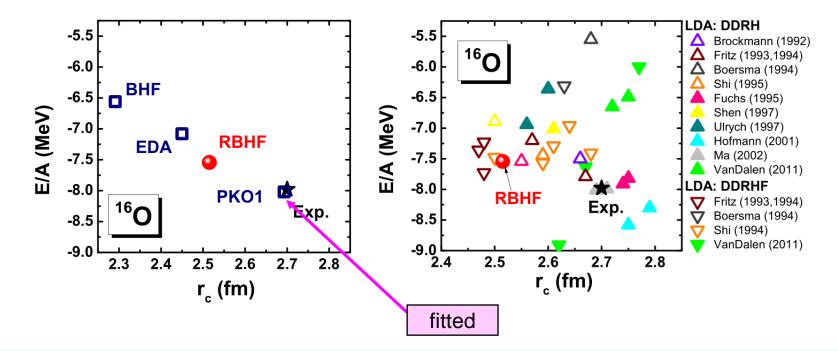
Brockmann and Toki, PRL 68, 3408 (1992).



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Bulk properties of ¹⁶O:

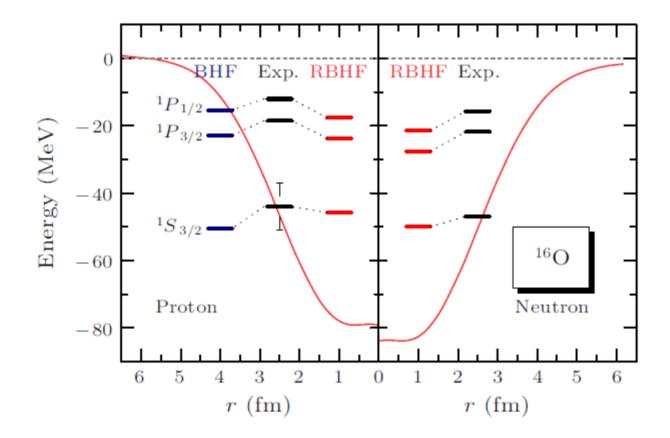
 Energy per particle and charge radius of ¹⁶O calculated by RBHF, nonrelativistic BHF Müther1990PRC, BHF with EDA Müther1990PRC, RHF with PKO1 Long2010PLB (left); and RBHF with LDA (right) :



- Relativistic effect is very important to improve the description.
- There is a big uncertainty between different LDA calculations.

XXIII. Nuclear Physics Workshop in Kazimierz-Dolny, Sept. 27-Oct. 2, 2016

DWS-basis



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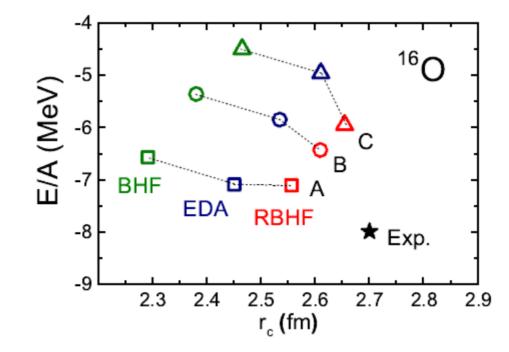
	$E \; (MeV)$	$r_c \ ({\rm fm})$	r_m (fm)	$\Delta E_{\pi 1p}^{ls}$ (MeV)
Exp. [84–87]	-127.6	2.70	2.54(2)	6.3
RBHF, Bonn A	-113.5	2.56	2.42	5.4
RBHF (DWS) [63]	-120.7	2.52	2.38	6.0
DDRHF, PKO1 [70]	-128.3	2.68	2.54	6.4
DDRHF, PKA1 [88]	-127.0	2.80	2.67	6.0
BHF [89], AV18	-134.2		1.95	13.0
CC [91], $N^{3}LO$	-120.9		2.30	
NCSM [90], N^3LO	-119.7(6)			
NLEFT $[93]$, N ² LO	-121.4(5)			

-			
	$E ({\rm MeV})$	$r_c ~({ m fm})$	$r_p \ ({\rm fm})$
Exp. [84, 85, 99]	-28.30	1.68	1.46
RBHF (PBV), Bonn A	-35.05	1.83	1.64
RBHF (PAV), Bonn A	-26.31	1.90	1.73
DDRHF, PKO1 [70]	-28.45	1.90	1.72
DDRHF, PKA1 [88]	-28.28	2.06	1.90
FY [100], CD-Bonn	-26.26		
$FY [101], N^4LO$	-24.27(6)		1.547(2)
$NCSM [102], N^3LO$	-25.39(1)		1.515(2)
NLEFT $[93]$, N ² LO	-25.60(6)		
BHF [89], AV18	-25.90°		

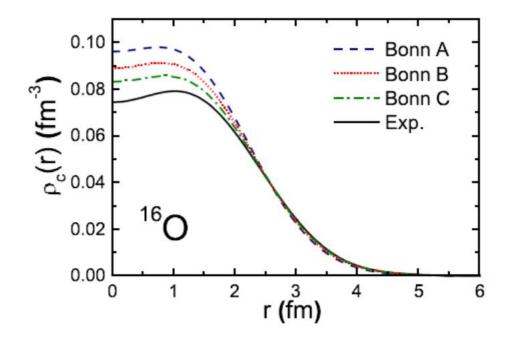
⁴⁰Ca

	$E ({\rm MeV})$	$r_c (\mathrm{fm})$	r_m (fm)	$\Delta E_{\pi 1d}^{ls}$ (MeV)
Exp. [84, 85, 87]	-342.1	3.48		6.6 ± 2.5
RBHF , Bonn A	-290.8	3.23	3.11	5.8
DDRHF, PKO1 [70]	-343.3	3.44	3.33	6.6
DDRHF, PKA1 [88]	-341.7	3.53	3.41	7.2
BHF [89], AV18	-552.1		2.20	24.9
NCSM [105], AV18	-461.8		2.27	
CC [106], AV18	-502.9			
$CC [107], N^{3}LO$	-345.2			

Self-consistent basis



Self-consistent basis



Conclusions

- Covariant density functional theory is very successful
- So far it is phenomenological (8-10 parameters)
- Semi-microscopic with only 4 parameters uses microscopic input (density dependence from BHF...)
- For the required accuracy (10⁻⁴) we need fine-tuning
- In order to understand the phenomenological models and to decide about additional terms in the Lagrangian (e.g. tensor) we need ab-initio derivations
- Rel. Brueckner-Hartree-Fock calculations in finite nuclei:
 - a) Rel. Brueckner-Hartree-Fock in local density approximation
 - b) Rel. Brueckner-Hartree-Fock in a Dirac-Woods-Saxon basis
 - c) Rel. Brueckner-Hartree-Fock fully selfconsistent

we solve the Bethe-Goldstone equation in each step of the iteration selfconsistently.

Outlook

- Heavy nuclei and the tensor force
- Other microscopic forces
- Pairing and open shell nuclei
- Deformed nuclei
- Optical potential
- •
- Three-body forces ?
- Beyond Brueckner ?

Thanks to my collaborators:

Shihang Shen(Beijing)Jie Meng(Beijing)

M. Serra[†] S. Karatzikos (Thessaloniki) G. A. Lalazissis (Thessaloniki) T. Otsuka (Tokyo)

T. Niksic (Zagreb) D. Vretenar (Barcelona)

A. Afanasjev (Mississippi) E. Litvinova (Kalamazoo)

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