## Production of Neutron-Rich Rare Isotopes toward the astrophysical r-proceess

## **Georgios A. Souliotis**

Laboratory of Physical Chemistry, Department of Chemistry, National and Kapodistrian University of Athens, Athens, Greece

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# Outline

- Introduction
- Chart of Nuclides
- Nucleosynthesis processes: focus on trans-Fe elements
- Research Results and Directions
- Rare Isotope Production
- Nuclear Fission Spallation
- Summary-Conclusions



#### The Nuclear Landscape and the r-process path



## Rapid Neutron capture process (r-process)



<sup>59</sup>Fe  $t_{1/2} = 44.5 \text{ d},$ <sup>68</sup>Fe  $t_{1/2} = 0.019 \text{ s},$  <sup>71</sup>Co  $t_{1/2} = 0.09 \text{ s},$  <sup>76</sup>Ni  $t_{1/2} = 0.24 \text{ s}$ 

\* e.g. C.A. Bertulani, T. Kajino, Prog. Nucl. Part. Physics 89, 56 (2016)





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#### **Production Methods of Rare Isotope Beams**<sup>[1]</sup>



\*http://www.sc.doe.gov/np/nsac/nsac.html

## RIB facility @ S. KOREA: Raon

Raon: 라온 meaning happy/joyful in Korean **Raon:** Heavy ion accelerator planned to be constructed by 2021 in South Korea

**RISP** (Rare Isotope Science Project): Research facility of the **IBS** (Institute for Basic Science), Daejeon, South Korea



https://www.ibs.re.kr/eng/sub01\_05.do

#### **FUTURE ACCELERATOR COMPLEX IN RAON**

Innovative feature: ISOL + IF facilities (ideally coupled)



## Characteristics of Nuclear Reactions: Energy Regimes

• Low Energy:	< 10	MeV/u	( u < 0.15 c )
• Fermi Energy:	10 - 50 I	MeV/u	( u = 0.15 - 0.3 c )
Medium Energy:	50 -200	MeV/u	( u = 0.3 - 0.6 c )
• High Energy:	> 200	MeV/u	( u > 0.6 c )





#### **Collisions between Heavy Ions at Fermi Energies (E/A < 50MeV)**



## Microscopic Calculations: Constrained Molecular Dynamics Model (CoMD)

#### **CoMD:** Quantum Molecular Dynamics (Semi classical)

- Nucleons considered as gaussian wavepackets
- Solution of the classical Hamilton's equations for the centroids
- Phenomenological N-N interaction (Skyrme-type)
- Symmetry potential depending on the nuclear density
- Surface potential (can be isospin dependent)
- Emulation of Pauli principle through appropriate restriction in phase space (phase space constraint)
- Recognition of cluster (fragment) formation (RN-N = 2.4 fm)
- Simulation of a large number of events (Monte Carlo approach)
- Continue evolution for 300-500 fm/c.
- Obtain properties of primary fragments





#### **Microscopic Calculations: Constrained Molecular Dynamics (CoMD)\***





\*M. Papa, A. Bonasera et al., Phys. Rev. C 64, 024612 (2001)

#### CoMD Calculations: <sup>86</sup>Kr (15 MeV/nucleon) + <sup>124</sup>Sn



**CoMD:** Constraint Molecular Dynamics; M. Papa, A. Bonasera, Phys. Rev. C 64, 024612 (2001)



#### Cyclotron Institute at Texas A&M University









#### Comparison: Data, Calculations: <sup>86</sup>Kr (15 MeV/nucleon) + <sup>64</sup>Ni



\* P.N. Fountas, G.A. Souliotis et al., Phys. Rev C 90, 064613 (2014)

• <sup>86</sup>Kr + <sup>64</sup>Ni (15 MeV/u)\*

- CoMD/SMM

----- CoMD/GEMINI

\*data: G.A. Souliotis et al., Phys. Rev. C 84, 064607 (2011)

- **CoMD:** Constrained Molecular Dynamics: M.Papa et. al., Phys. Rev. C 64, 024612 (2001)
- GEMINI: Binary Decay Code: R. Charity, Nucl. Phys. A 483, 391 (1988)
- SMM: Statistical Multifragmentation Model: A. Botvina et al., Phys. Rev. C 65, 044610 (2002); Nucl. Phys. A 507, 649 (1990)

#### <sup>92</sup>Kr RNB Cross Sections and Rate estimates

Rare isotope	Reaction Channel	Cross Section (mb)	Rates (s⁻¹)
<sup>93</sup> Kr	- 0p + 1n	18.8	$1.1 \times 10^{4}$
<sup>94</sup> Kr	- 0p + 2n	2.3	$1.3 \times 10^{3}$
<sup>95</sup> Kr	- 0p + 3n	0.63	$3.8 \times 10^{2}$
<sup>96</sup> Kr	- 0p + 4n	0.20	$1.2 \times 10^{2}$
<sup>92</sup> Br	- 1p + 1n	4.5	2.7 × 10 <sup>3</sup>
<sup>93</sup> Br	- 1p + 2n	0.75	$4.5 \times 10^{2}$
<sup>94</sup> Br	- 1p + 3n	0.078	47
<sup>95</sup> Br	- 1p + 4n	0.039	23
<sup>96</sup> Br	- 1p + 5n	0.008	5
<sup>90</sup> Se	- 2p + 0n	2.70	$1.6 \times 10^{3}$
<sup>91</sup> Se	- 2p + 1n	0.60	$3.5 \times 10^{2}$
<sup>92</sup> Se	- 2p + 2n	0.12	70
<sup>93</sup> Se	- 2p + 3n	0.039	23

- Cross sections and rate estimates of neutron-rich isotopes from <sup>92</sup>Kr (15 MeV/nucleon) + <sup>64</sup>Ni
- Beam of <sup>92</sup>Kr with intensity 0.5 pnA (3 x 10<sup>9</sup> particles/s)
- <sup>64</sup>Ni target with a thickness of 20 mg/cm<sup>2</sup>

\* P.N. Fountas, G.A. Souliotis, M. Veselsky, et al, Phys. Rev. C 90, 064613 (2014)

#### **Example of nuclide production in DIC with RIBs:**



Rate estimates: <sup>92</sup>Kr from RAON at 0.5pnA (~3x10<sup>9</sup>pps), <sup>64</sup>Ni (20mg/cm<sup>2</sup>):

#### 1mb => 600 pps

\* P.N. Fountas, G.A. Souliotis et al, Phys. Rev. C 90, 064613 (2014)



## Experimental setup: The MAGNEX spectrometer



$$\theta_{\text{MAGNEX}} = 9^{\circ}$$
  $\Delta \theta = 4^{\circ} - 16^{\circ}$ 

#### **Experimental setup: MAGNEX target chamber**



#### Calculations: DIT/SMM: <sup>70</sup>Zn (15 MeV/nucleon) + <sup>64</sup>Ni, <sup>208</sup>Pb



**Mass-resolved angular distributions** 

For this test run : one angle setting  $\theta_{MAGNEX} = 9^{\circ}$   $\Delta \theta = 4^{\circ} - 16^{\circ}$ 

#### Preliminary DIT/SMM Calculations: <sup>70</sup>Zn (15 MeV/nucleon) + <sup>64</sup>Ni

#### Mass distributions of near-projectile isotopes

Mass Number A



Expected rate: 1µb => ~ 4 cpm



 Experimental study of peripheral reactions at energy ~10-30 MeV/nucleon Beams: <sup>48</sup>Ca, <sup>70</sup>Zn, <sup>82</sup>Se (targets: <sup>64</sup>Ni, <sup>124</sup>Sn, <sup>208</sup>Pb, <sup>238</sup>U, look ~ θ<sub>gr</sub>)

Projectile fission of <sup>238</sup>U (15 MeV/nucleon)

Excitation-energy reconstruction and study (10-30 MeV/nucleon data)

Experience for experimental studies using neutron-rich RIBs (FRIBS/LNS, FRIB/MSU, RISP/Korea)

Access the neutron-drip line near Z~26-28 with RIBs of Zn, Ni

## Nuclear Fission Studies with CoMD

## p (30 MeV) + <sup>235</sup>U

- lur235\_sur0\_t10k\_animate.dist\_out' using 5:6:( ev(\$1)\*te(\$2)\*nn(\$4)\* \$7 ) •
- lur235\_sur0\_t10k\_animate.dist\_out' using 5:6:( ev(\$1)\*te(\$2)\*pp(\$4)\* \$7 ) •



\*N. Vonta, G.A. Souliotis, et al., Phys. Rev C 87, 014001 (2015).



## Comparison between Theoretical and Experimental Results: p (30 MeV) + <sup>235</sup>U



Red line: standard V<sub>sym</sub> ~  $\rho$ Blue line: soft V<sub>sym</sub> ~  $\rho^{1/2}$ 

Grey points: experimental data: S.I. Mulgin et al., Nucl. Phys. A 824, 1 (2009)



\*N. Vonta, G.A. Souliotis, et al., Phys. Rev C 87, 014001 (2015).

#### Fission Calculations: <sup>238</sup>U (20 MeV/nucleon) + <sup>64</sup>Ni



Production cross sections of neutron-rich nuclides

\*N. Vonta, G.A. Souliotis, *et al., Phys. Rev C* (Dec.2016).

## Overview of spallation and its applications

- Astrophysics (reactions induced by cosmic rays in interstellar medium)
- Accelerator-driven systems (ADS) ("Energy amplification")
- Transmutation of nuclear waste
- Sources (Facilities) of spallation neutrons
- Production of rare isotopes (ISOL facilities)

## Spallation reaction mechanism



# Spallation: p(500MeV) +<sup>208</sup>Pb



\*A. Assimakopoulou, G.A. Souliotis, *A. Bonasera, M. Veselsky, in preparation* (Aug. 2016).

# Comparison between theoretical results and experimental data: $p(1000 \text{ MeV}) + {}^{208}Pb$



#### Exp. Data: T. Enqvist et al., Nucl. Phys. A 686, 481 (2001)

\*A. Assimakopoulou, G.A. Souliotis, et al.,

## Fission Cross Section of <sup>208</sup>Pb

□CoMD calculations: Red line: standard □ Blue line: soft

Black points: experimental data



1)J. L. Rodriguez et al., Phys. Rev. C 90, 064606 (2014)

2)T. Enqvist et al., Nucl. Phys. A 686, 481-524 (2001)

3)B. Fernandez et al., Nucl. Phys. A 747, 227-267 (2005) 4)K. -H. Schmidt et al., Phys. Rev. C 87, 034601 (2013)

5) A. V. Prokofiev, Nucl. Instr. Meth. A 463 557-575 (2001)

## Fission cross section/residue cross section



1)T. Enqvist et al., Nucl. Phys. A 686, 481-524 (2001)

2)B. Fernandez-Dominguez et al., Nucl. Phys. A 747, 227-267 (2005)

3)M. Bernas et al., Nucl. Phys. A 725, 213-253 (2003) 5)F. Rejmund et al., Nucl. Phys. A 683 540-565 (2001) 4)J. Benlliure, P. Armbruster et al., Nucl. Phys. A 700 469-491 (2002)

# Summary

- Overview of Nuclear Studies
- Nuclear Chart-Nuclear Stability
- Overview of Nucleosynthesis. Focus on r-process
- Research Results and Directions
- Rare Isotope Production
- Nuclear Multifragmentation: hot nuclei
- Nuclear Fission, Spallation





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